## THE

# PRINCIPLES OF DYNAMICS 

## BY

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## FOREWORD

The content of the present little book is based upon a lecture that I gave as an extension of the general lecture on mechanics in Spring-Summer 1918. My goal was to show how the various principles of dynamics develop from each other and are connected together. In particular, it seemed to me that a simple, but rigorous, explanation of how the various forms of the principle of least action come about would be desirable today since that principle has proved to be a powerful heuristic force in Einstein's theory of gravitation. I have entered into the HamiltonJacobi theory because it has likewise approached the sphere of interest of physicists by the work of Schwarzschild, Sommerfeld, and Epstein.

I wish to express my heartfelt thanks to Adolf Kneser for his interest and advice.
Breslau, April 1919

## C. Schaefer.

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## § 1.

## The principle of virtual displacements for holonomic and scleronomic equations of constraint.

As is known, the equations of equilibrium for a system of $n$ completely-free mass-points (so one with $3 n$ degrees of freedom) read:

$$
\left\{\begin{array}{l}
X_{v}=0,  \tag{1}\\
Y_{v}=0, \\
Z_{v}=0
\end{array} \quad(v=1,2, \ldots, n)\right.
$$

in which $X_{v}, Y_{v}, Z_{v}$ mean the rectangular components of the total force that acts upon the $v^{\text {th }}$ mass-particle. Those three equations admit a simple combination when we multiply each of equations (1) by an arbitrary function, which we would like to denote by $\delta x_{v}, \delta y_{v}, \delta z_{v}$, respectively, and then add the entire system:

$$
\begin{equation*}
\sum_{v=1}^{N}\left(X_{v} \delta x_{v}+Y_{v} \delta y_{v}+Z_{v} \delta z_{v}\right)=0 \tag{2}
\end{equation*}
$$

Due to the arbitrariness in the choice of the $3 n$ functions $\delta x_{v}, \delta y_{v}, \delta z_{v}$, that equation is completely equivalent to the system (1); it says nothing more or less than the latter. That is because we can choose, e.g., $\delta x_{1} \neq 0$, and choose all of the remaining $\delta x_{v}$, as well as all of $\delta y_{v}$ and $\delta z v$, to be equal to zero. Equation (2) will then produce the result:

$$
X_{1} \delta x_{1}=0
$$

or since we have $\delta x_{1} \neq 0$, by assumption, it will follow that $X_{1}=0$. We will once more obtain all equations in the system (1) in succession in that way.

If one interprets the quantities $\delta x_{v}, \delta y_{v}, \delta z_{v}$ as the components of a displacement that one imagines to be infinitely small, and which all mass-points suffer, then the expression $X_{v} \delta x_{v}+$ $Y_{v} \delta y_{v}+Z_{v} \delta z_{v}$ will represent the work $\delta A_{\nu}$ that the forces $X_{v}, Y_{v}, Z_{v}$ do under the displacement that one imagines. The left-hand side of (2), which can be written as $\sum_{v} \delta A_{v}=\delta A$, will then represent the total work $\delta A$ that is done by all forces when each of the $n$ mass-points suffers a displacement $\delta \mathfrak{s v}$. We can then formulate the equilibrium condition as follows:

For a completely-free system of mass-points, the work done by an infinitely-small displacement of the system will be equal to zero in equilibrium.

Equation (2) represents the simplest case of the so-called principle of virtual displacements. Namely, the imagined displacements that are assigned to the system, which one cares to denote
by the symbol $\delta$ from the calculus of variations, in contrast to the actual displacements, which we will characterize by the symbol " $d$," will be called by the medieval term virtual (i.e., possible) displacements, in order to emphasize the fact that they differ from the actual displacements. That definition, which is sufficient for the present purposes, must be subjected to suitable restrictions in what follows. Here, it is important to point out that in the case of complete freedom of motion for the mass-system, which we have indeed assumed here, any conceivable displacement is referred to as virtual. In particular, the actual displacements are also "possible," i.e., virtual ones.

In the case of a completely unconstrained system that was just considered, the principle of virtual displacements that is included in equation (2) says nothing more than the system of equations (1) and has no advantage over the latter. Things will be different when we shift our attention to the fact that the freedom of motion might be restricted by certain equations of constraint (there might be $m$ of them with $m<3 n$ ) that exist between the coordinates $x_{1}, y_{1}, z_{1}, x_{2}$, $y_{2}, z_{2}, \ldots, x_{n}, y_{n}, z_{n}$. For instance, we might prescribe that each mass-point might displace only on a certain surface, and we would then ask what equilibrium in the system, thus-restricted, would be.

There would then exist equations of the form:

$$
\begin{equation*}
\varphi_{\kappa}\left(x_{1}, y_{1}, z_{1}, \ldots, x_{v}, y_{v}, z_{v}, \ldots, x_{n}, y_{n}, z_{n}\right)=0 \quad(\kappa=1,2, \ldots, m) . \tag{3}
\end{equation*}
$$

What form will the equilibrium conditions take now?
The fact that this problem is essentially more difficult than the previous one is illuminated when one makes the following clear: Instead of restricting the degrees of freedom of the masspoints by the constraint equations (3), one can also combine the old forces $X_{v}, Y_{v}, Z_{v}$ with suitably-chosen new ones $\Xi_{v}, \mathrm{H}_{v}, \mathrm{Z}_{v}$, with the stipulation that the additional forces should be arranged such that they will prevent the mass-points from leaving the prescribed surface. Their ultimate effect should then be to replace the constraint equations (3). One will indeed know the effect of those additional forces then, but not their magnitudes. In order to have a suitable name for those additional forces, we would like to call them the forces of constraint.

For the sake of simplicity, we shall first consider a mass-point (mass $m_{v}$, coordinates $x_{v}, y_{v}$, $z_{v}$ ) and a constraint equation:

$$
\varphi_{\kappa}\left(x_{v}, y_{v}, z_{v}\right)=0 .
$$

We can always go on to the general case of several mass-points and constraints later by summation then. The simple case that we have taken up here makes it possible for us to give it a geometric interpretation in ordinary three-dimensional space. The constraint equation $\varphi_{\kappa}=0$ represents a surface in it upon which the point must remain. If we replace the constraint equation $\varphi_{\kappa}=0$ with the constraint forces $\Xi_{V}, \mathrm{H}_{\nu}, \mathrm{Z}_{\nu}$ then our mass-point will once more be completely free, except that it will no longer be acted upon by the forces $X_{v}, Y_{v}, Z_{v}$, but by the forces $X_{v}+\Xi_{v}$, $Y_{v}+\mathrm{H}_{v}, Z_{v}+\mathrm{Z}_{v}$. We can then apply the equilibrium condition (1) for a free system and have:

$$
\begin{equation*}
X_{v}+\Xi_{v}=0, \quad Y_{v}+\mathrm{H}_{v}=0, \quad Z_{v}+\mathrm{Z}_{v}=0 \tag{4}
\end{equation*}
$$

or when we go over to the domain of the principle of virtual displacements, i.e., we subject our mass-point to a displacement $\overline{\delta \mathfrak{s}_{v}}$, then we will find from (2) that:

$$
\begin{equation*}
\left(X_{v}+\Xi_{v}\right) \overline{\delta x_{v}}+\left(Y_{v}+\mathrm{H}_{v}\right) \overline{\delta y_{v}}+\left(Z_{v}+\mathrm{Z}_{v}\right) \overline{\delta z_{v}}=0 \tag{5}
\end{equation*}
$$

or also:

$$
\begin{equation*}
X_{v} \overline{\delta x_{v}}+Y_{v} \overline{\delta y_{v}}+Z_{v} \overline{\delta z_{v}}=-\left(\Xi_{v} \overline{\delta x_{v}}+\mathrm{H}_{v} \overline{\delta y_{v}}+\mathrm{Z}_{v} \overline{\delta z_{v}}\right) \tag{6}
\end{equation*}
$$

i.e., the work done by the explicit forces $\Xi_{v}, \mathrm{H}_{v}, \mathrm{Z}_{v}$ under the imagined motion will be equal and opposite to the work done by the forces of constraint $\Xi_{v}, \mathrm{H}_{v}, \mathrm{Z}_{v}$ in equilibrium. That equation is really of no use, since the $\Xi_{v}, \mathrm{H}_{v}, \mathrm{Z}_{v}$ are generally unknown. Therefore, in order to go any further so that we can make a definite statement about equilibrium in this case of restricted degrees of freedom, we must add something new. That new thing takes the form of adding a further statement about the nature of the forces of constraint that can be quite plausible in certain concrete cases, but not proved in general, and it can therefore be introduced only as a hypothesis in the general case. We then address the following situation: If we observe that the forces of constraint $\Xi_{v}, \mathrm{H}_{v}, \mathrm{Z}_{v}$ are only supposed to prevent the mass-point from leaving the surface then that suggests that we might assume that the forces of constraint are directed normal to the surface $\varphi_{\kappa}=0$, i.e., that $\Xi_{\nu}, H_{\nu}, Z_{\nu}$ should be proportional to the direction cosines of the surface:

$$
\frac{\frac{\partial \varphi_{\kappa}}{\partial x_{v}}, \frac{\partial \varphi_{\kappa}}{\partial y_{v}}, \frac{\partial \varphi_{\kappa}}{\partial z_{v}}}{\sqrt{\left(\frac{\partial \varphi_{\kappa}}{\partial x_{v}}\right)^{2}+\left(\frac{\partial \varphi_{\kappa}}{\partial y_{v}}\right)^{2}+\left(\frac{\partial \varphi_{\kappa}}{\partial z_{v}}\right)^{2}}}
$$

If $\lambda_{\kappa}$ means the proportionality factor then from the assumption above, the forces of constraint can be written:

$$
\Xi_{v}=\frac{\partial \varphi_{\kappa}}{\partial x_{v}}, \quad \mathrm{H}_{\nu}=\frac{\partial \varphi_{\kappa}}{\partial y_{v}}, \quad \mathrm{Z}_{v}=\frac{\partial \varphi_{\kappa}}{\partial z_{v}} .
$$

The equilibrium conditions for our mass-point would then be:

$$
\left\{\begin{array}{l}
X_{v}+\lambda_{\kappa} \frac{\partial \varphi_{\kappa}}{\partial x_{v}}=0 \\
Y_{v}+\lambda_{\kappa} \frac{\partial \varphi_{\kappa}}{\partial y_{v}}=0  \tag{7}\\
Z_{v}+\lambda_{\kappa} \frac{\partial \varphi_{\kappa}}{\partial z_{v}}=0
\end{array}\right.
$$

or also when we use (6) to go over to the principle of virtual displacements:

$$
\begin{equation*}
X_{\nu} \overline{\delta x_{v}}+Y_{v} \overline{\delta y_{v}}+Z_{v} \overline{\delta z_{v}}=-\lambda_{\kappa}\left(\frac{\partial \varphi_{\kappa}}{\partial x_{v}} \overline{\delta x_{v}}+\frac{\partial \varphi_{\kappa}}{\partial y_{v}} \overline{\delta y_{v}}+\frac{\partial \varphi_{\kappa}}{\partial z_{v}} \overline{\delta z_{v}}\right) \tag{8}
\end{equation*}
$$

The $\overline{\delta x_{v}}, \overline{\delta y_{v}}, \overline{\delta z_{v}}$ in that are completely independent of each other, since we have indeed suppressed the equation of constraint. In order to distinguish those completely-free displacements $\overline{\delta x_{v}}, \ldots$ from more specialized ones that are likewise introduced, we have provided them with an overbar. The four unknowns $x_{v}, y_{v}, z_{v}, \lambda_{\kappa}$ can be determined from the equilibrium condition (7) or (8), in conjunction with the equation of the surface $\varphi_{\kappa}=0$. We then see how the introduction of our hypothesis proves to be practicable.

Up to now, as we have emphasized, the quantities $\overline{\delta x_{v}}, \overline{\delta y_{v}}, \overline{\delta z_{v}}$ could be chosen to be completely arbitrary and independent of each other. For that reason, the right-hand side of (8) cannot be made to vanish, in general. However, if that right-hand side of (8) can be made to vanish then we would gain a great advantage due to the fact that we would get the same expression for the equilibrium condition in the case of restricted degrees of freedom that we would get in the case of a completely-free system. That is because (8) and (2) would be equal to each other in that case.

Now, how can one succeed in making the right-hand side of (8) equal to zero? Obviously, to that end, the $\overline{\delta x_{v}}, \overline{\delta y_{v}}, \overline{\delta z_{v}}$ cannot be completely arbitrary, but the components of the displacement must be suitably restricted, and indeed as follows:

Before the displacement, the coordinates $x_{v}, y_{v}, z_{v}$ of the mass-point will obey the condition $\varphi_{\kappa}=0$ that was imposed upon them. By contrast, up to now, the coordinates $x_{v}+\overline{\delta x_{v}}, y_{v}+\overline{\delta y_{v}}$, $z_{v}+\overline{\delta z_{v}}$ did not need to do that after the displacement since I can indeed choose the displacement to be, say, perpendicular to the surface so it points away from the surface. That is because the displacement $\overline{\delta x_{v}}, \ldots$ was chosen to be completely free. I shall now restrict my choice of displacement in such a way that I demand that the mass-point should also still lie on the surface after the displacement, i.e., If I now impose the condition upon the displacement $\overline{\delta x_{v}}$ , ... that it must be compatible with the equilibrium conditions then we will have:

$$
\left\{\begin{array}{l}
\varphi_{\kappa}\left(x_{v}, y_{v}, z_{v}\right)=0  \tag{9}\\
\varphi_{\kappa}\left(x_{v}+\delta x_{v}, y_{v}+\delta y_{v}, z_{v}+\delta z_{v}\right)=0
\end{array}\right.
$$

if the displacements, thus-specialized, are denoted by $\delta x_{v}, \delta y_{v}, \delta z_{v}$ (with no overbar). Thus, when we develop the second equation in a Taylor series and truncate it after the linear terms, we will get:

$$
\varphi_{\kappa}\left(x_{v}, y_{v}, z_{v}\right)+\frac{\partial \varphi_{\kappa}}{\partial x_{v}} \delta x_{v}+\frac{\partial \varphi_{\kappa}}{\partial y_{v}} \delta y_{v}+\frac{\partial \varphi_{\kappa}}{\partial z_{v}} \delta z_{v}=0 .
$$

Finally, upon subtracting the first equation (9) from the one that was just obtained, we will get:

$$
\begin{equation*}
\frac{\partial \varphi_{k}}{\partial x_{v}} \delta x_{v}+\frac{\partial \varphi_{\kappa}}{\partial y_{v}} \delta y_{v}+\frac{\partial \varphi_{\kappa}}{\partial z_{v}} \delta z_{v}=0 \tag{10}
\end{equation*}
$$

We can now say that: We choose our displacements components $\delta x_{v}, \delta y_{v}, \delta z_{v}$ in such a way that they will obey equation (10). Naturally, the right-hand side of equation (8) would vanish then, as we will get the principle of virtual displacements in the old form:

$$
\begin{equation*}
X_{v} \delta x_{v}+Y_{v} \delta y_{v}+Z_{v} \delta z_{v}=0 \tag{11}
\end{equation*}
$$

but it will no longer be true for arbitrary displacements $\overline{\delta x_{v}}, \overline{\delta y_{v}}, \overline{\delta z_{v}}$, but only for ones $\delta x_{v}$, $\delta y_{v}, \delta z_{v}$ that obey the varied constraint equation $\varphi_{\kappa}=0$, i.e., equation (10).

From now on, we shall restrict the term "virtual displacement" in the following way:
We now understand the virtual displacements to be only those imagined displacements that are compatible with the equilibrium conditions. In what follows, we shall always denote them by the symbol $\delta$.

The equilibrium conditions for our mass-point are included in (11), which is valid simultaneously with (10). Naturally, we can now no longer conclude that $X_{v}=Y_{\nu}=Z_{v}=0$ from (11), i.e., the principle of virtual displacements, as we could in the case of complete degrees of freedom, since it would indeed also be false, according to equation (7), but rather since the three displacement components are no longer independent of each other, we must proceed as follows: We multiply (10) by an arbitrary factor $\lambda_{\kappa}$ and add it to (11). It would then follow that:

$$
\begin{equation*}
\left(X_{v}+\lambda_{\kappa} \frac{\partial \varphi_{\kappa}}{\partial x_{v}}\right) \delta x_{v}+\left(Y_{v}+\lambda_{\kappa} \frac{\partial \varphi_{\kappa}}{\partial y_{v}}\right) \delta y_{v}+\left(Z_{v}+\lambda_{\kappa} \frac{\partial \varphi_{\kappa}}{\partial z_{v}}\right) \delta z_{v}=0 . \tag{12}
\end{equation*}
$$

From equation (10), we can regard one of the three displacements in that as being determined by the other two, say, $\delta x_{v}$ is determined in terms of $\delta y_{v}$ and $\delta z_{v}$. Therefore, whereas $\delta y_{v}$ and $\delta z_{v}$ can be chosen arbitrarily, $\delta x_{v}$ will be determined by them from (10). We now choose $\lambda_{\kappa}$, which is still undetermined, but it can still be chosen freely, in such a way that the coefficient of $\delta x_{v}$ in equation (12) will vanish. All that will then remain are the terms with mutually-independent displacements $\delta y_{v}$ and $\delta z_{v}$, and their coefficients must vanish in their own right due to just that independence of $\delta y_{v}$ and $\delta z_{v}$. We will then get the following three equations, which are completely symmetric, although the argument that leads to the first one is different from the one that leads to the last two:

$$
\begin{aligned}
& X_{v}+\lambda_{\kappa} \frac{\partial \varphi_{\kappa}}{\partial x_{v}}=0, \\
& Y_{v}+\lambda_{\kappa} \frac{\partial \varphi_{\kappa}}{\partial y_{v}}=0
\end{aligned}
$$

$$
Z_{v}+\lambda_{\kappa} \frac{\partial \varphi_{\kappa}}{\partial z_{v}}=0
$$

They are once more equations (7), which suffice to determine the equilibrium state in conjunction with constraint equation $\varphi_{\kappa}=0$, as was mentioned above.

What was carried out for one mass-point and one constraint equation can be easily generalized to $n$ mass-points and $m$ constraint equations. A geometric representation is also possible then that is completely analogous to the one above, except that we must now go to $3 n$ dimensional space, i.e., we must regard the coordinates ( $x_{1}, y_{1}, z_{1}, \ldots, x_{v}, y_{v}, z_{v}, \ldots, x_{n}, y_{n}, z_{n}$ ) that determine the configuration of the system as the coordinates of one point in $3 n$-dimensional space. Every constraint equation of the form $\varphi_{\kappa}\left(x_{1}, \ldots, z_{n}\right)=0$ will then represent a hypersurface in that space, and when $m$ such constraint equations are present, that will mean that the point ( $x_{1}$, $\left.y_{1}, z_{1}, \ldots, x_{n}, y_{n}, z_{n}\right)$ that represents the configuration of the system must be found on the common intersection of those $m$ hypersurfaces. Moreover, we can then proceed precisely as we did in the foregoing, and indeed we can even use the same words as before.

Thus, if $\lambda_{\kappa}(\kappa=1,2, \ldots, m)$ denote the so-called Lagrange multipliers then we will get:

$$
\begin{equation*}
\sum_{v=1}^{n}\left(X_{v} \delta x_{v}+Y_{v} \delta y_{v}+Z_{v} \delta z_{v}\right)=0 \tag{13}
\end{equation*}
$$

as the equilibrium condition for those displacements $\delta x_{v}, \delta y_{v}, \delta z_{v}$ that satisfy the following equations:

$$
\left\{\begin{array}{lrrl}
\frac{\partial \varphi_{1}}{\partial x_{1}} \delta x_{1}+\frac{\partial \varphi_{1}}{\partial y_{1}} \delta y_{1}+\frac{\partial \varphi_{1}}{\partial z_{1}} \delta z_{1}+\cdots+\frac{\partial \varphi_{1}}{\partial x_{n}} \delta x_{n}+\frac{\partial \varphi_{1}}{\partial y_{n}} \delta y_{n} & +\frac{\partial \varphi_{1}}{\partial z_{n}} \delta z_{n} & =0  \tag{14}\\
\frac{\partial \varphi_{\kappa}}{\partial x_{1}} \delta x_{1}+ & +\frac{\partial \varphi_{\kappa}}{\partial z_{n}} \delta z_{n} & =0 \\
\frac{\partial \varphi_{m}}{\partial x_{1}} \delta x_{1}+ & +\frac{\partial \varphi_{m}}{\partial z_{n}} \delta z_{n} & =0
\end{array}\right.
$$

i.e., more briefly: for all virtual displacements (in the sense that was just established). If equations (14) were successively extended with the factors $\lambda_{1}, \ldots, \lambda_{m}$ and added to (13) then that would produce the equation:

$$
\sum_{v=1}^{n}\left[\left(X_{v} \delta x_{v}+Y_{v} \delta y_{v}+Z_{v} \delta z_{v}\right)+\sum_{\kappa=1}^{m} \lambda_{\kappa}\left(\frac{\partial \varphi_{\kappa}}{\partial x_{v}} \delta x_{v}+\frac{\partial \varphi_{\kappa}}{\partial y_{v}} \delta y_{v}+\frac{\partial \varphi_{\kappa}}{\partial z_{v}} \delta z_{v}\right)\right]=0
$$

from which, the same argument that we applied to equation (11) would produce the following system:

$$
\left\{\begin{array}{l}
X_{v}+\sum_{\kappa=1}^{m} \lambda_{\kappa} \frac{\partial \varphi_{\kappa}}{\partial x_{v}}=0,  \tag{15}\\
Y_{v}+\sum_{\kappa=1}^{m} \lambda_{\kappa} \frac{\partial \varphi_{\kappa}}{\partial y_{v}}=0, \\
Z_{v}+\sum_{\kappa=1}^{m} \lambda_{\kappa} \frac{\partial \varphi_{\kappa}}{\partial z_{v}}=0
\end{array}\right.
$$

Together with the $m$ constraint equations $\varphi_{\kappa}=0(\kappa=1,2, \ldots, m)$, they will suffice to determine the $(3 n+m)$ unknowns $\left(x_{1}, y_{1}, z_{1}, \ldots, x_{n}, y_{n}, z_{n}, \lambda_{1}, \ldots, \lambda_{m}\right)$. Equilibrium is then established by the formula for the principle of virtual displacements.

The constraint equations $\varphi_{\kappa}\left(x_{1}, \ldots, z_{n}\right)=0$ that were considered here do not include time explicitly, and for that reason, Boltzmann called them scleronomic equations, in order to distinguish them from equations of the form $\varphi_{\kappa}\left(x_{1}, \ldots, z_{n}, t\right)=0$, which Boltzmann called rheonomic, due to the explicit appearance of time in them. Both types of equation have in common that they represent relations between the finite quantities $x_{1}, \ldots, z_{n}$, and possibly time $t$. However, more general forms that represent relations between the differentials $d x_{1}, \ldots, d z_{n}$, and possibly the time differential $d t$ are conceivable:

$$
\begin{equation*}
\sum_{\nu=1}^{n}\left(a_{\kappa \nu} d x_{v}+b_{\kappa v} d y_{v}+c_{\kappa v} d z_{v}\right)=0 \tag{16}
\end{equation*}
$$

and

$$
\begin{equation*}
\sum_{v=1}^{n}\left(a_{\kappa v}^{\prime} d x_{v}+b_{\kappa v}^{\prime} d y_{v}+c_{\kappa v}^{\prime} d z_{v}\right)=0 . \tag{17}
\end{equation*}
$$

In case those equations are not integrable, i.e., the coefficients $a, b, c\left(a^{\prime}, b^{\prime}, c^{\prime}\right.$, resp.) do not fulfill the integrability conditions, the latter equations cannot be brought into the form that was considered up to now. One distinguishes the two forms of the constraint equations from each other by the names holonomic and non-holonomic, which go back to H. Hertz. (The terminology goes back to the Greek roots ó ó $о \varsigma=$ whole and vó $\mu о \varsigma=$ law.) Equation (16) will then be referred to as non-holonomic-scleronomic, while equation (17) is non-holonomic-rheonomic.

Up to now, we have exhibited the principle of virtual displacements only for holonomicscleronomic constraint equations, and in the following section, we will investigate how it is formulated for non-holonomic and rheonomic ones. That will come down to establishing what we understand virtual displacements to mean in that case.

## The principle of virtual displacements for rheonomic and non-holonomic equations of constraint.

We next ask what the physical meaning of a holonomic and rheonomic equation of constraint:

$$
\begin{equation*}
\varphi_{\kappa}\left(x_{1}, y_{1}, z_{1}, \ldots, x_{n}, y_{n}, z_{n}, t\right)=0 \tag{18}
\end{equation*}
$$

might be. Obviously, for every constant value of $t$, it represents a hypersurface in $3 n$-dimensional space. If we let $t$ vary then we will get a surface moving in that space. A simple example of that is the following one: Let a mass-point be constrained to move on the surface of a sphere whose center moves with constant velocity $c$ parallel to the $x$-axis. We will examine that example in more detail later. Is the principle of virtual displacements also valid here in the form (13)?

In order to answer that question, we once more start from the fact that the equations of constraint can be replaced with forces of constraint $\Xi_{v}, \mathrm{H}_{v}, \mathrm{Z}_{v}$. We again make the hypothesis that the force of constraint is perpendicular to the surface, i.e., that $\Xi_{v}, H_{v}, Z_{v}$ are proportional to the direction cosines of the surface. For one mass-point and one equation of constraint, we will then get the old relations:

$$
\Xi_{v}=\lambda_{\kappa} \frac{\partial \varphi_{\kappa}}{\partial x_{v}}, \mathrm{H}_{v}=\lambda_{\kappa} \frac{\partial \varphi_{\kappa}}{\partial y_{v}}, \mathrm{Z}_{v}=\lambda_{\kappa} \frac{\partial \varphi_{\kappa}}{\partial z_{v}} .
$$

The equilibrium condition:

$$
\begin{equation*}
X_{v}+\lambda_{\kappa} \frac{\partial \varphi_{\kappa}}{\partial x_{v}}=0, Y_{v}+\lambda_{\kappa} \frac{\partial \varphi_{\kappa}}{\partial y_{v}}=0, \quad Z_{v}+\lambda_{\kappa} \frac{\partial \varphi_{\kappa}}{\partial z_{v}}=0 \tag{19}
\end{equation*}
$$

will then follow in this case, just as before, or when we focus upon a completely-arbitrary displacement $\overline{\delta \mathfrak{s}_{v}}$ with the independent components $\overline{\delta x_{v}}, \overline{\delta y_{v}}, \overline{\delta z_{v}}$ :

$$
\begin{equation*}
X_{v} \overline{\delta x_{v}}+Y_{v} \overline{\delta y_{v}}+Z_{v} \overline{\delta z_{v}}=-\lambda_{\kappa}\left(\frac{\partial \varphi_{\kappa}}{\partial x_{v}} \overline{\delta x_{v}}+\frac{\partial \varphi_{\kappa}}{\partial y_{v}} \overline{\delta y_{v}}+\frac{\partial \varphi_{\kappa}}{\partial z_{v}} \overline{\delta z_{v}}\right) \tag{20}
\end{equation*}
$$

Just as before, we will now return to the old form (13) of the principle of virtual displacements when we restrict the allowable displacements in such a way that they will satisfy the equation:

$$
\begin{equation*}
\frac{\partial \varphi_{\kappa}}{\partial x_{v}} \delta x_{v}+\frac{\partial \varphi_{\kappa}}{\partial y_{v}} \delta y_{v}+\frac{\partial \varphi_{\kappa}}{\partial z_{v}} \delta z_{v}=0 \tag{21}
\end{equation*}
$$

in which we now drop the overbar again.
Those constraint equations arise formally in such a way that one varies the constraint equation $\varphi_{\kappa}\left(x_{1}, y_{1}, z_{1}, t\right)=0$ while holding time constant. We then establish that we understand rheonomic constraints to mean virtual displacements that obey (21), i.e., the varied constraint equation $\varphi_{\kappa}\left(x_{1}, y_{1}, z_{1}, t\right)=0$ when time is not varied. The old equation:

$$
\begin{equation*}
X_{v} \delta x_{v}+Y_{v} \delta y_{v}+Z_{v} \delta z_{v}=0 \tag{22}
\end{equation*}
$$

in conjunction with (21), will again produce the equilibrium conditions.
That result generalizes directly from one mass-point and one constraint to a system of $n$ mass-points and $m$ constraints in the way that was given in the first section, which does not need to be specified in more detail here. For us, it is enough that we can formulate the result once more as:

For holonomic constraints, whether scleronomic or rheonomic, the principle of virtual displacements is valid in the form (13) for all virtual displacements that are defined by equations (14) in each case.

On first glance, it might seem remarkable that time does not need to be varied along with everything else, because time will naturally be varied by an actual displacement, which would indeed result in time. If we denote the components of such a thing (in the simple case of one mass-point and one constraint) by, say, $d x_{v}, d y_{v}, d z_{v}$ then that must naturally correspond to the equation:

$$
\begin{equation*}
\frac{\partial \varphi_{\kappa}}{\partial x_{v}} d x_{v}+\frac{\partial \varphi_{\kappa}}{\partial y_{v}} d y_{v}+\frac{\partial \varphi_{\kappa}}{\partial z_{v}} d z_{v}+\frac{\partial \varphi_{\kappa}}{\partial t} d t=0 \tag{23}
\end{equation*}
$$

since the surface $\varphi_{\kappa}\left(x_{v}, y_{v}, z_{v}, t\right)=0$ will indeed advance accordingly in the time interval $d t$. That implies that here the actual displacements do not belong with the virtual ones, although that was the case for scleronomic constraints. The confusion will be resolved, moreover, when we consider the fact that in the case of rheonomic constraints, the partial derivatives $\frac{\partial \varphi_{\kappa}}{\partial x_{v}}, \frac{\partial \varphi_{\kappa}}{\partial y_{v}}, \ldots$ generally include time as a parameter, so it is by no means true that the system must obey the same constraints in the scleronomic and rheonomic cases.

The aforementioned example will make the situation clearer, in which one asks what equilibrium would be for a point that is supposed to remain on a sphere of radius $R$ about the point $\left(x_{0}, y_{0}, z_{0}\right)$ when gravity acts upon it at the same time. Let the sphere be initially at rest, so the constraint is scleronomic. It the $z$-axis points vertically upwards then the explicit forces will be:

$$
\begin{equation*}
X=Y=0, \quad Z=-m g, \tag{24}
\end{equation*}
$$

and the constraint will further read:

$$
\begin{equation*}
\varphi(x, y, z) \equiv\left(x-x_{0}\right)^{2}+\left(y-y_{0}\right)^{2}+\left(z-z_{0}\right)^{2}-R^{2}=0 . \tag{25}
\end{equation*}
$$

The virtual displacements are then subject to the following constraint here:

$$
\begin{equation*}
\left(x-x_{0}\right) \delta x+\left(y-y_{0}\right) \delta y+\left(z-z_{0}\right) \delta z=0 . \tag{26}
\end{equation*}
$$

Along with a factor $\lambda$, in conjunction with the principle of virtual displacements that was formulated in (13), that will yield the equilibrium conditions:

$$
\left\{\begin{array}{r}
0+\lambda\left(x-x_{0}\right)=0,  \tag{27}\\
0+\lambda\left(y-y_{0}\right)=0, \\
-m g+\lambda\left(z-z_{0}\right)=0 .
\end{array}\right.
$$

The equilibrium position $(\bar{x}, \bar{y}, \bar{z})$ is calculated from that as follows: The first two equations in (27) imply that:

$$
\left\{\begin{array}{l}
\bar{x}=x_{0},  \tag{28.a}\\
\bar{y}=y_{0} .
\end{array}\right.
$$

Thus, from the constraint equation (25), one further has:

$$
\begin{equation*}
\bar{z}-z_{0}= \pm R \tag{28.b}
\end{equation*}
$$

i.e., the point is at either highest or lowest point in the sphere at rest, which is explained immediately. (Naturally, the equilibrium is labile in the first case, which is established by a special examination in each case.) With those values of $\bar{x}, \bar{y}, \bar{z}$, one finally finds the value of $\lambda$ from the third equation in (27):

$$
\begin{equation*}
\lambda= \pm \frac{m g}{R} \tag{28.c}
\end{equation*}
$$

in which the upper or lower sign will be true according to whether the mass-point is at the highest or lowest point, resp.

The center of the sphere might now move along the positive $x$-axis with a velocity of $c$. The constraint will then be rheonomic, and indeed one will have:

$$
\begin{equation*}
\varphi(x, y, z, t) \equiv\left(x-x_{0}-c t\right)^{2}+\left(y-y_{0}\right)^{2}+\left(z-z_{0}\right)^{2}-R^{2}=0 . \tag{29}
\end{equation*}
$$

From our convention, the virtual displacements are subject to the following constraint here:

$$
\begin{equation*}
\left(x-x_{0}-c t\right) \delta x+\left(y-y_{0}\right) \delta y+\left(z-z_{0}\right) \delta z=0 \tag{30}
\end{equation*}
$$

and although time is not also varied, we see that this says something totally different from the constraint (26) in the scleronomic case. Together with the principle of virtual displacements that was formulated in (13), along with a factor $\lambda$, we will get the following equilibrium equations here:

$$
\left\{\begin{align*}
0+\lambda\left(x-x_{0}-c t\right) & =0  \tag{31}\\
0+\lambda\left(y-y_{0}\right) & =0 \\
-m g+\lambda\left(y-y_{0}\right) & =0
\end{align*}\right.
$$

which correspondingly differ from (27).
The equilibrium position $(\bar{x}, \bar{y}, \bar{z})$ and the factor $\lambda$ follow in the same way as before:

$$
\left\{\begin{align*}
\bar{x} & =x_{0}+c t  \tag{32}\\
\bar{y} & =y_{0} \\
\bar{z}-z_{0} & = \pm R \\
\lambda & = \pm \frac{m g}{R}
\end{align*}\right.
$$

i.e., the point can once more be at the highest or lowest point of the moving sphere in equilibrium.

We shall now move on to the non-holonomic (scleronomic or rheonomic) constraints. It is easy to see now that the principle of virtual displacements will again be valid in the form (13) when we define the virtual displacements as follows: They must obey the constraints:

$$
\begin{array}{ll}
\text { for scleronomic constraints: } & \sum_{\nu}\left(a_{\kappa \nu} \delta x_{v}+b_{\kappa \nu} \delta y_{v}+c_{\kappa \nu} \delta z_{v}\right)=0,  \tag{33}\\
\prime \text { rheonomic } \quad " \quad \sum_{v}\left(a_{\kappa \nu}^{\prime} \delta x_{v}+b_{\kappa \nu}^{\prime} \delta y_{v}+c_{\kappa \nu}^{\prime} \delta z_{v}\right)=0 .
\end{array}
$$

Those equations will be obtained from equations (16) and (17), which typically represent non-holonomic constraints, when one switches the symbol " $d$ " with " $\delta$ ' and sets $d t$ equal to zero, i.e., time is not varied. Namely, that will follow directly from the fact that in our previous explanation of the equations of constraint, we did not actually use them in their "holonomic" form $\varphi_{\kappa}\left(x_{1}, \ldots, z_{n}, t\right)=0$, but only in their varied form, in which they represented differential relations between the differentials (variations, resp.). The fact that they fulfilled the integrability conditions was not used at all.

At the same time, let it be remarked here that such non-holonomic constraints always appear when one deals with the rolling of spheres and similar bodies on a certain surface, e.g., a plane. If the surface on which the sphere should roll moves in space then we would simultaneously have a non-holonomic-rheonomic constraint.

As the final result, we can then establish that:
The principle of virtual displacements is true for every type of constraint - holonomic, nonholonomic, rheonomic, scleronomic - in the form (13), in which virtual displacements are understood in the most general case to mean that they satisfy equations of the form (34).

That is because the non-holonomic, rheonomic equations of constraint (17) reduce to scleronomic ones in the special case in which time does not appear explicitly and to holonomic ones when the integrability conditions are fulfilled.

## § 3.

## D'Alembert's principle.

As is known from Newton's second axiom, the equations of motion for a system of $n$ completely-free mass-points (so $3 n$ degrees of freedom) read:

$$
\left\{\begin{align*}
& X_{v}-m_{v} \ddot{x}_{v}=0,  \tag{35}\\
& Y_{v}-m_{v} \ddot{y}_{v}=0, \\
& Z_{v}-m_{v} \ddot{z}_{v}=0
\end{align*} \quad(v=1,2, \ldots, n)\right.
$$

If we denote the left-hand sides briefly by $X_{v}^{\prime}, Y_{v}^{\prime}, Z_{v}^{\prime}$ then we can write those equations as:

$$
\left\{\begin{array}{l}
X_{v}^{\prime}=0,  \tag{36}\\
Y_{v}^{\prime}=0, \\
Z_{v}^{\prime}=0 .
\end{array} \quad(v=1,2, \ldots, n)\right.
$$

They then take the same form as equations (1) for equilibrium in a free system of $n$ masspoints, as we considered it in $\S 1$.

The fruitful interpretation of the equations of motion, for which we have d'Alembert to thank, is based upon this formal agreement between equations (36) and (1), and is what goes by the name of d'Alembert's principle.

Obviously, one can formulate (36) as:

Every mass-point that exhibits an accelerated motion under the influence of the explicit forces $X_{v}, Y_{v}, Z_{v}$ will be brought into equilibrium under the common influence of the forces $X_{v}$ $m_{v} \ddot{x}_{v}=X_{v}^{\prime}, Y_{v}-m_{v} \ddot{y}_{v}=Y_{v}^{\prime}, Z_{v}-m_{v} \ddot{z}_{v}=Z_{v}^{\prime}$,
or:

If one adds the so-called d'Alembertian inertial forces $-m_{v} \ddot{x}_{v},-m_{v} \ddot{y}_{v},-m_{v} \ddot{z}_{v}$ to the impressed forces $X_{v}, Y_{v}, Z_{v}$ then those forces taken together, i.e., the forces $X_{v}-m_{v} \ddot{x}_{v}, Y_{v}-$ $m_{v} \ddot{y}_{v}, Z_{v}-m_{v} \ddot{z}_{v}$, will keep the mass-point in equilibrium.

D'Alembert's principle then reduces the problems of dynamics to problems of equilibrium.
Now, in his Mécanique analytique, Lagrange took the further step by saying that d'Alembert's principle, as the simplest expression that we can consider, viz., (35) [(36), resp.],
should be coupled with the principle of virtual displacements. In fact, if we multiply the $3 n$ equations (36) by the arbitrary functions $\delta x_{v}, \delta y_{v}, \delta z_{v}$ and add them then it will follow that:

$$
\sum_{v=1}^{n}\left(X_{v}^{\prime} \delta x_{v}+Y_{v}^{\prime} \delta y_{v}+Z_{v}^{\prime} \delta z_{v}\right)=0
$$

or, from (35):

$$
\begin{equation*}
\sum_{v=1}^{n}\left[\left(X_{v}-m_{v} \ddot{x}_{v}\right) \delta x_{v}+\left(Y_{v}-m_{v} \ddot{y}_{v}\right) \delta y_{v}+\left(Z_{v}-m_{v} \ddot{z}_{v}\right) \delta z_{v}\right]=0 . \tag{37}
\end{equation*}
$$

In analytical mechanics, that particular equation is called d'Alembert's principle. We will adopt that terminology here and always understand that principle to mean equation (37).

Equation (37) is precisely equivalent to equations (35) for a completely-free mass-system, because the quantities $\delta x_{v}, \delta y_{v}, \delta z_{v}$, which we naturally once more consider to be infinitelysmall displacements $\delta \mathfrak{s}$, are mutually independent. However, if certain constraints are present, say, $m(<3 n)$ of them, that we will assume to be non-holonomic and rheonomic here, for the sake of generality:

$$
\begin{equation*}
\sum_{v=1}^{n}\left(a_{\kappa v}^{\prime} d x_{v}+b_{\kappa v}^{\prime} d y_{v}+c_{\kappa v}^{\prime} d z_{v}\right)+a_{v}^{\prime} d t=0 \quad(\kappa=1,2, \ldots, m), \tag{38}
\end{equation*}
$$

then it will follow in analogy to the foregoing section that equation (37) will retain its validity when we understand $\delta x_{v}, \delta y_{v}, \delta z_{v}$ to mean virtual displacements, i.e., displacements that obey the following conditions:

$$
\begin{equation*}
\sum_{v=1}^{n}\left(a_{\kappa v}^{\prime} \delta x_{v}+b_{\kappa v}^{\prime} \delta y_{v}+c_{\kappa \nu}^{\prime} \delta z_{v}\right)=0 \quad(\kappa=1,2, \ldots, m) \tag{39}
\end{equation*}
$$

With the help of $m$ Lagrange multipliers, upon combining equations (37) for d'Alembert's principle and the $m$ equations (38), one will get the equation:

$$
\begin{equation*}
\sum_{v=1}^{n}\left[\left(X_{v}-m_{v} \ddot{x}_{v}+\sum_{\kappa=1}^{m} \lambda_{\kappa} a_{\kappa v}^{\prime}\right) \delta x_{v}+\left(Y_{v}-m_{v} \ddot{y}_{v}+\sum_{\kappa=1}^{m} \lambda_{\kappa} b_{\kappa v}^{\prime}\right) \delta y_{v}+\left(Z_{v}-m_{v} \ddot{z}_{v}+\sum_{\kappa=1}^{m} \lambda_{\kappa} c_{\kappa v}^{\prime}\right) \delta z_{v}\right]=0, \tag{40}
\end{equation*}
$$

which will then give the following system of Lagrange equations of motion of the first kind, in just the same way as what followed from equation (12) in § 1:

$$
\left\{\begin{array}{l}
X_{v}-m_{v} \ddot{x}_{v}+\sum_{\kappa=1}^{m} \lambda_{\kappa} a_{\kappa v}^{\prime}=0  \tag{41}\\
Y_{v}-m_{v} \ddot{y}_{v}+\sum_{\kappa=1}^{m} \lambda_{\kappa} b_{\kappa v}^{\prime}=0 \\
Z_{v}-m_{v} \ddot{z}_{v}+\sum_{\kappa=1}^{m} \lambda_{\kappa} c_{\kappa v}^{\prime}=0
\end{array}\right.
$$

Those $3 n$ equations, together with the $m$ equations (38), will suffice to determine the ( $3 n+m$ ) unknowns $x_{1}, y_{1}, z_{1}, \ldots, x_{n}, y_{n}, z_{n}, \lambda_{1}, \lambda_{2}, \ldots, \lambda_{m}$.

As we have repeatedly emphasized, virtual displacements are determined in the same way here as they were in the principle of virtual displacements. In particular, time is not varied for rheonomic constraints, as we will see, e.g., when we go from (38) to (39).

Naturally, equations (41) also subsume the case of scleronomic and holonomic constraints. If the constraints are initially scleronomic then the coefficients $a_{\kappa \nu}^{\prime}, b_{\kappa \nu}^{\prime}, c_{\kappa \nu}^{\prime}$ will not include time explicitly, which we would like to suggest by dropping the prime, as with our previous notation. If the equations are holonomic, in addition, then the coefficients $a_{\lambda v}, b_{\lambda v}, c_{\lambda v}$ will be equal to the partial derivatives $\frac{\partial \varphi_{\kappa}}{\partial x_{v}}, \frac{\partial \varphi_{\kappa}}{\partial y_{v}}, \frac{\partial \varphi_{\kappa}}{\partial z_{v}}$, resp., of a certain function $\varphi_{\kappa}$.

## § 4.

## The energy principle for scleronomic and rheonomic equations of constraint.

This is the place for us to establish the circumstances under which the energy principle will be included in the equations of motion of mechanics (41). It will be shown that this is the case only when the equations of constraint are scleronomic, i.e., they do not include time explicitly, whereas holonomity or non-holonomity is irrelevant to that.

We take the most-general case of non-holonomic-rheonomic constraints, whose equations of motion were exhibited in (41). Thus, perhaps $m$ equations of constraint of the form (38) might exist:

$$
\sum_{v=1}^{n}\left(a_{\kappa v}^{\prime} d x_{v}+b_{\kappa v}^{\prime} d y_{v}+c_{\kappa v}^{\prime} d z_{v}\right)+a_{v}^{\prime} d t=0 \quad(\kappa=1,2, \ldots, m),
$$

from which, dividing by $d t$ will produce the following relations for the velocity components $\dot{x}_{v}$, $\dot{y}_{v}, \dot{z}_{v}$ :

$$
\sum_{v=1}^{n}\left(a_{\kappa v}^{\prime} \dot{x}_{v}+b_{\kappa v}^{\prime} y_{v}^{\prime}+c_{\kappa v}^{\prime} \ddot{z}_{v}\right)+a_{v}^{\prime}=0 \quad(\kappa=1,2, \ldots, m)
$$

If we successively extend the equations of motion by the velocity components $\dot{x}_{v}, \dot{y}_{v}, \dot{z}_{v}$, add them, and arrange them then we will get:

$$
\begin{equation*}
\sum_{v=1}^{n}\left(X_{v} \dot{x}_{v}+Y_{v} \dot{y}_{v}+Z_{v} \dot{z}_{v}\right)-\sum_{v=1}^{n} m_{v}\left(\ddot{x}_{v} \dot{x}_{v}+\ddot{y}_{v} \dot{y}_{v}+\ddot{z}_{v} \dot{z}_{v}\right)+\sum_{\kappa=1}^{m} \lambda_{\kappa}\left(a_{\kappa v}^{\prime} \dot{x}_{v}+b_{\kappa v}^{\prime} \dot{y}_{v}+c_{\kappa v}^{\prime} \dot{z}_{v}\right)=0 . \tag{43}
\end{equation*}
$$

The individual summands in that equation have simple meanings. The first sum represents the work done by the impressed forces $X_{v}, Y_{\nu}, Z_{v}$ per unit time; we would like to denote it by $d A^{\prime} / d t$. The differential symbol is given a prime in order to suggest that $d^{\prime} A$ is does not necessarily need to be an exact differential of a pure function of the coordinates, which will be true only in special cases that are generally very important. The second summand can be written:

$$
\begin{equation*}
\sum_{v=1}^{n} m_{v}\left(\ddot{x}_{v} \dot{x}_{v}+\ddot{y}_{v} \dot{y}_{v}+\ddot{z}_{v} \dot{z}_{v}\right)=\frac{1}{2} \sum_{v=1}^{n} m_{v}\left(\dot{x}_{v}^{2}+\dot{y}_{v}^{2}+\dot{z}_{v}^{2}\right)=\frac{d}{d t} \sum_{v=1}^{n} L_{v}=\frac{d L}{d t} . \tag{44}
\end{equation*}
$$

$L_{v}$ means the kinetic energy of the $v^{\text {th }}$ mass-particle, and $\sum L_{v}=L$ means that of the total system. Finally, based upon equations (42), the third sum can be written:

$$
\begin{equation*}
\sum_{\kappa=1}^{m} \lambda_{\kappa}\left(a_{\kappa v}^{\prime} \dot{x}_{v}+b_{\kappa v}^{\prime} \dot{y}_{v}+c_{\kappa \nu}^{\prime} \dot{z}_{v}\right)=-\sum_{\kappa=1}^{m} \lambda_{\kappa} a_{\kappa}^{\prime} . \tag{45}
\end{equation*}
$$

Along with (44) and (45), equation (43) will become the following one:

$$
\begin{equation*}
\frac{d L}{d t}=\frac{d^{\prime} A}{d t}-\sum_{\kappa=1}^{m} \lambda_{\kappa} a_{\kappa}^{\prime}, \tag{46}
\end{equation*}
$$

while the energy principle is known to demand that the time derivative of the kinetic energy $d L / d t$ should be equal to the work done per unit time $d A^{\prime} / d t$.

We then see that the energy principle is not true in the case of rheonomic constraints, which are indeed characterized by the appearance of the coefficients $a_{\kappa}^{\prime}$, and in fact it is irrelevant whether the rheonomic constraints are holonomic or non-holonomic. That is because if we assume that the constraint (42) can be converted into a holonomic one (due to the integrability conditions being valid) then we will indeed have:

$$
a_{\kappa v}^{\prime}=\frac{\partial \varphi_{\kappa}}{\partial x_{v}}, \quad b_{\kappa v}^{\prime}=\frac{\partial \varphi_{\kappa}}{\partial y_{v}}, \quad c_{\kappa v}^{\prime}=\frac{\partial \varphi_{\kappa}}{\partial z_{v}}, \quad a_{\kappa}^{\prime}=\frac{\partial \varphi_{\kappa}}{\partial t}
$$

and equation (46) would go to:

$$
\begin{equation*}
\frac{d L}{d t}=\frac{d^{\prime} A}{d t}-\sum_{\kappa=1}^{m} \lambda_{\kappa} \frac{\partial \varphi_{\kappa}}{\partial t}, \tag{46.a}
\end{equation*}
$$

which proves the assertion above.
By contrast, if the constraints are scleronomic, i.e., the coefficients $a_{\kappa}^{\prime}$ vanish (in the holonomic case, $\partial \varphi_{\kappa} / \partial t=0$, resp.), then the energy principle will, in fact, follow:

$$
\begin{equation*}
\frac{d L}{d t}=\frac{d^{\prime} A}{d t} . \tag{47}
\end{equation*}
$$

The energy principle will then be true for only scleronomic (holonomic or non-holonomic) constraint equations.

In special cases, the forces can be represented as the negatives of partial derivatives of a single function $\Phi$ of the coordinates $x_{1}, y_{1}, z_{1}, \ldots, x_{n}, y_{n}, z_{n}$ :

$$
\begin{equation*}
X_{v}=-\frac{\partial \Phi}{\partial x_{v}}, \quad Y_{v}=-\frac{\partial \Phi}{\partial y_{v}}, \quad Z_{v}=-\frac{\partial \Phi}{\partial z_{v}} \quad(v=1,2, \ldots, n), \tag{48}
\end{equation*}
$$

and the expression:

$$
\frac{d^{\prime} A}{d t}=\sum_{v=1}^{n}\left(X_{v} \dot{x}_{v}+Y_{v} \dot{y}_{v}+Z_{v} \dot{z}_{v}\right)
$$

will go to:

$$
\begin{equation*}
\frac{d^{\prime} A}{d t}=-\sum_{v=1}^{n}\left(\frac{\partial \Phi}{\partial x_{v}} \dot{x}_{v}+\frac{\partial \Phi}{\partial y_{v}} \dot{y}_{v}+\frac{\partial \Phi}{\partial z_{v}} \dot{z}_{v}\right)=-\frac{d \Phi}{d t} \tag{49}
\end{equation*}
$$

which will make the energy principle (47) assume the simple form:

$$
\left\{\begin{align*}
\frac{d}{d t}(L+\Phi) & =0  \tag{50}\\
\text { or } \quad L+\Phi & =h
\end{align*}\right.
$$

in which $h$ means a constant. One calls $\Phi$ the potential energy, while $h=L+\Phi$ means the total energy.

In the latter case, i.e., when a potential energy $\Phi$ exists, we would like to refer to equation (50) as the energy integral, which we would like to characterize as the narrower form of the law of energy, to distinguish it from its broader form (47).

The result that was obtained that the energy principle is not valid for rheonomic constraints is important in the investigation of the least action principle (§ 12), among other things.

## § 5.

## True, transitional, and varied motion.

The fundamental difference between holonomic and non-holonomic constraints can be highlighted in a different way, which we would now like to explain.

We first consider the true motion of one mass-point. (Everything that is essential here will become clear when we restrict ourselves to this simplest system. Moreover, the statements that will be made here can be adapted verbatim to a system of $n$ mass-points when we go to $3 n$ dimensional space.) We can think of it as being represented in, say, the form:

$$
\left\{\begin{array}{l}
x_{v}=x_{v}(t)  \tag{51}\\
y_{v}=y_{v}(t) \\
z_{v}=z_{v}(t)
\end{array}\right.
$$



Figure 1.

We associate each point $x_{v}, y_{v}, z_{v}$ with of the true path at time $t$ with another neighboring point $\left(x_{v}+\delta x_{v}, y_{v}+\delta y_{v}, z_{v}+\delta z_{v}\right)$ at the same time $t$ by means of the virtual displacements $\delta x_{v}$, $\delta y_{v}, \delta z_{v}$. The set of all those latter points will likewise define a continuous path, at least when $\delta x_{v}, \delta y_{v}, \delta z_{v}$ are continuous functions of time, which we assume. We shall call that path the varied path. Naturally, we can associate the true path with a varied path in very different ways, such as when we let the point $x_{v}, y_{v}, z_{v}$ at time $t$ on the true path correspond to the point ( $x_{v}+$ $\left.\Delta x_{v}, y_{v}+\Delta y_{v}, z_{v}+\Delta z_{v}\right)$ at time $t+\Delta t$ on the varied path. However, the displacements $\Delta x_{v}, \Delta y_{v}$, $\Delta z_{v}$ would not be virtual displacements then, in general, which is why they are also not denoted by the symbol $\delta$. Nonetheless, we shall only deal with virtual displacements, so for our considerations, the varied path will emerge from the true path in the first way that was given.

We can perhaps depict that graphically by considering only the $x_{v} t$-plane (Fig. 1). Naturally, we would then get only the projection of the true and varied paths onto the $x$-axis.

As is shown in the figure, a point $P$ on the true path is associated with a point $P^{\prime}$ that corresponds to the same value of time $t$. We call $\overline{P P^{\prime}}$ the transitional path. The components of $\overline{P P^{\prime}}$ are $\delta x_{v}, \delta y_{v}, \delta z_{v}$.

Now let a holonomic constraint exist that we would initially like to assume is scleronomic, for simplicity:

$$
\begin{equation*}
\varphi\left(x_{v}, y_{v}, z_{v}\right)=c_{1}, \tag{52}
\end{equation*}
$$

in which the constant $c_{1}$ can naturally also have the value 0 , but it does not need to. For the sake of clarity, we prefer that the value of the constant should remain undetermined.

We shall now address the issue of what sort of constraints the true, the transitional, and the varied paths should be subject to.

For the true path, according to the problem that was posed, equation (52) must be fulfilled for every position of the material point along it, e.g., for the point ( $x_{v}, y_{v}, z_{v}$ ), as well as for its neighboring point $\left(x_{v}+d x_{v}, y_{v}+d y_{v}, z_{v}+d z_{v}\right)$ (which it will assume at time $t+d t$, although that will not be an issue here). Thus, along with (52), one will also have the equation:

$$
\varphi\left(x_{v}+d x_{v}, y_{v}+d y_{v}, z_{v}+d z_{v}\right)=0
$$

for the neighboring point, or:

$$
\varphi\left(x_{v}, y_{v}, z_{v}\right)+\frac{\partial \varphi}{\partial x_{v}} d x_{v}+\frac{\partial \varphi}{\partial y_{v}} d y_{v}+\frac{\partial \varphi}{\partial z_{v}} d z_{v}=c_{1}
$$

or after subtracting equation (52):

$$
\begin{equation*}
\frac{\partial \varphi}{\partial x_{v}} d x_{v}+\frac{\partial \varphi}{\partial y_{v}} d y_{v}+\frac{\partial \varphi}{\partial z_{v}} d z_{v} \equiv d \varphi=0 . \tag{53}
\end{equation*}
$$

The components of the true displacement $d x_{v}, d y_{v}, d z_{v}$ then obey the relation (53), which says just that the value $c_{1}$ of $\varphi$ does not change along the true motion.

How does that work for the transitional motion?
During it, which is indeed produced by virtual displacements $\delta x_{v}, \delta y_{v}, \delta z_{v}$ as a result of d'Alembert's principle, the latter displacements must obey the equation:

$$
\begin{equation*}
\frac{\partial \varphi}{\partial x_{v}} \delta x_{v}+\frac{\partial \varphi}{\partial y_{v}} \delta y_{v}+\frac{\partial \varphi}{\partial z_{v}} \delta z_{v} \equiv \delta \varphi=0 \tag{54}
\end{equation*}
$$

as was shown before, i.e., the change in the value $c_{1}$ of $\varphi$ will also be zero during the transitional motion, that is, the transitional motion will also obey the constraint equation (52).

A statement about the behavior of the varied path will now follow directly from that.
$\varphi$ has the value $c_{1}$ at the point $P$ on the true path. $\varphi$ will not change during the transitional motion that leads to the point $P^{\prime}(\delta \varphi=0)$. Therefore, the point $P^{\prime}$ on the varied path will also be assigned the same value $\varphi=c_{1}$, i.e., the varied path will also obey the constraint (52) under our assumptions.

We have then arrived at the following result:

For holonomic constraints, the true, the transitional, and the varied path will obey the equations of constraint.

As we had already anticipated in the formulation above, that will also be true when the holonomic constraint equations are rheonomic since $t$ is treated as a constant parameter during the transitional motion.

In the above, we concluded the behavior of the varied path from that of the true and the transitional paths. Had we demanded from the outset that the varied path should also satisfy the holonomic constraints, along with the true path, then we could have concluded in the same way that the transitional path would have to do so, as well, i.e., that the displacements $\delta x_{v}, \delta y_{v}, \delta z_{v}$ that are thus constituted would have to be virtual ones.

That behavior will change when we now move on to non-holonomic equations.
Therefore, let a non-holonomic (but scleronomic) constraint be given now:

$$
\begin{equation*}
d^{\prime} \varphi \equiv a_{\kappa v} d x_{v}+b_{\kappa \nu} d y_{v}+c_{\kappa v} d z_{v}=0, \tag{55}
\end{equation*}
$$

in which the $a_{\kappa v}, \ldots$ are functions of $x_{v}, y_{v}, z_{v}$ that do not fulfill the integrability conditions. We have denoted the left-hand side of the constraint by $d^{\prime} \varphi$, in which the prime suggests that it should not be treated as an exact differential of a function $\varphi$.

We will now first present the conditions for the transitional path to obey the constraint (55), i.e., for the displacements to be virtual. Previously, that followed for holonomic constraints from the fact that the varied path also obeyed the equations of constraint. Secondly, we will formulate the condition for the varied path to obey equation (55) and see whether those two determinations are also compatible with each other now ( ${ }^{1}$ ).

We shall first present the equations that have to be valid along the transitional path!
The virtual displacements $\delta x_{v}, \delta y_{v}, \delta z_{v}$ that characterize that motion have to obey the constraint that is given by (55) when we switch $d$ with $\delta$, i.e., the equation:

$$
\begin{equation*}
\delta^{\prime} \varphi \equiv a_{\kappa v} \delta x_{v}+b_{\kappa v} \delta y_{v}+c_{\kappa v} \delta z_{v}=0 \tag{56}
\end{equation*}
$$

whose left-hand side we logically denote by $\delta^{\prime} \varphi$. Now, that equation will be as valid when we complete the transition to the varied path from the point $\left(x_{v}, y_{v}, z_{v}\right)$ on the true path as it is when

[^0]we go to the neighboring point $\left(x_{v}+d x_{v}, y_{v}+d y_{v}, z_{v}+d z_{v}\right)$ on the true path, i.e., along with (56) [i.e., along with $\delta^{\prime} \varphi\left(x_{v}, y_{v}, z_{v}\right)=0$ ], it must also be true that:
$$
\delta^{\prime} \varphi\left(x_{v}+d x_{v}, y_{v}+d y_{v}, z_{v}+d z_{v}\right)=0,
$$
or after developing that in a Taylor series and subtracting equation (56):
\[

$$
\begin{equation*}
d\left(\delta^{\prime} \varphi\right) \equiv d\left(a_{\kappa \nu} \delta x_{\nu}+b_{\kappa \nu} \delta y_{\nu}+c_{\kappa \nu} \delta z_{\nu}\right)=0 . \tag{57}
\end{equation*}
$$

\]

Let us make note here of the special case that will emerge from (57) for a holonomic constraint equation:

$$
\begin{equation*}
d(\delta \varphi) \equiv d\left(\frac{\partial \varphi}{\partial x_{v}} \delta x_{v}+\frac{\partial \varphi}{\partial y_{v}} \delta y_{v}+\frac{\partial \varphi}{\partial z_{v}} \delta z_{v}\right)=0 . \tag{57.a}
\end{equation*}
$$

(57) is the condition that is implied by the demand that the transitional path should obey the prescribed equation.

We shall now formulate the other demand that the varied path should also do that. However, that obviously means that whereas the equation $\delta^{\prime} \varphi\left(x_{v}, y_{v}, z_{v}\right)=0$ is fulfilled for the true motion, the same equation should be valid for the varied coordinates $\left(x_{v}+\delta x_{v}, y_{v}+\delta y_{v}, z_{v}+\delta z_{v}\right)$, so:

$$
d^{\prime} \varphi\left(x_{v}+\delta x_{v}, y_{v}+\delta y_{v}, z_{v}+\delta z_{v}\right)=0,
$$

or upon developing that it in a Taylor series and subtracting equation (55):

$$
\begin{equation*}
\delta\left(d^{\prime} \varphi\right) \equiv \delta\left(a_{\kappa v} \delta x_{v}+b_{\kappa v} \delta y_{v}+c_{\kappa v} \delta z_{v}\right)=0, \tag{58}
\end{equation*}
$$

which will again specialize to:

$$
\begin{equation*}
\delta(d \varphi) \equiv \delta\left(\frac{\partial \varphi}{\partial x_{v}} d x_{v}+\frac{\partial \varphi}{\partial y_{v}} d y_{v}+\frac{\partial \varphi}{\partial z_{v}} d z_{v}\right)=0 \tag{58.a}
\end{equation*}
$$

for a holonomic constraint. We now ask whether the two conditions that were thus obtained (which we already know to be identical in the case of holonomity, which will also be shown again soon) are compatible with each other for non-holonomic equations.

By performing the differentiations that are suggested in (57), it will follow that:

$$
\begin{equation*}
\left(\frac{\partial a_{\kappa v}}{\partial x_{v}} d x_{v}+\cdots\right) \delta x_{v}+\left(\frac{\partial b_{\kappa v}}{\partial x_{v}} d x_{v}+\cdots\right) \delta y_{v}+\left(\frac{\partial c_{\kappa v}}{\partial x_{v}} d x_{v}+\cdots\right) \delta z_{v}+\left(a_{\kappa v} d \delta x_{v}+\cdots\right)=0, \tag{59}
\end{equation*}
$$

and likewise from (58):

$$
\begin{equation*}
\left(\frac{\partial a_{\kappa v}}{\partial x_{v}} \delta x_{v}+\cdots\right) d x_{v}+\left(\frac{\partial b_{\kappa v}}{\partial x_{v}} \delta x_{v}+\cdots\right) d y_{v}+\left(\frac{\partial c_{\kappa v}}{\partial x_{v}} \delta x_{v}+\cdots\right) d z_{v}+\left(a_{\kappa v} d \delta x_{v}+\cdots\right)=0 \tag{60}
\end{equation*}
$$

Subtraction will give the following relation that must exist between the two equations (57) and (58) in the case of compatibility:

$$
\left\{\begin{array}{c}
\left(\frac{\partial a_{\kappa v}}{\partial y_{v}}-\frac{\partial b_{\kappa v}}{\partial x_{v}}\right)  \tag{61}\\
\left(d y_{v} \delta x_{v}-d x_{v} \delta y_{v}\right)+\left(\frac{\partial b_{\kappa v}}{\partial z_{v}}-\frac{\partial c_{\kappa v}}{\partial y_{v}}\right)\left(d z_{v} \delta y_{v}-d y_{v} \delta z_{v}\right) \\
+\left(\frac{\partial c_{\kappa v}}{\partial x_{v}}-\frac{\partial a_{\kappa v}}{\partial x_{v}}\right)\left(d x_{v} \delta z_{v}-d z_{v} \delta x_{v}\right)=0 .
\end{array}\right.
$$

On the other hand, the following proportion can be inferred from (55) and (56):

$$
\begin{equation*}
a_{\kappa v}: b_{\kappa v}: c_{\kappa v}=\left(d z_{v} \delta y_{v}-d y_{v} \delta z_{v}\right):\left(d x_{v} \delta z_{v}-d z_{v} \delta x_{v}\right):\left(d x_{v} \delta y_{v}-d y_{v} \delta x_{v}\right), \tag{62}
\end{equation*}
$$

or when we appeal to a proportionality factor $A$ :

$$
\begin{align*}
& a_{\kappa v}=A \cdot\left(d x_{v} \delta y_{v}-d y_{v} \delta x_{v}\right), \\
& b_{\kappa v}=A \cdot\left(d x_{v} \delta z_{v}-d z_{v} \delta x_{v}\right),  \tag{63}\\
& c_{\kappa v}=A \cdot\left(d y_{v} \delta x_{v}-d x_{v} \delta y_{v}\right),
\end{align*}
$$

with which (61) can be written:

$$
\begin{equation*}
\left(\frac{\partial a_{\kappa v}}{\partial y_{v}}-\frac{\partial b_{\kappa v}}{\partial x_{v}}\right) c_{\kappa v}+\left(\frac{\partial b_{\kappa v}}{\partial z_{v}}-\frac{\partial c_{\kappa v}}{\partial y_{v}}\right) a_{\kappa v}+\left(\frac{\partial c_{\kappa v}}{\partial x_{v}}-\frac{\partial a_{\kappa v}}{\partial x_{v}}\right) b_{\kappa v}=0 . \tag{64}
\end{equation*}
$$

For a holonomic constraint equation, that relation will specialize to:

$$
\begin{equation*}
\left(\frac{\partial^{2} \varphi}{\partial x_{v} \partial y_{v}}-\frac{\partial^{2} \varphi}{\partial y_{v} \partial x_{v}}\right) \frac{\partial \varphi}{\partial z_{v}}+(\cdots) \frac{\partial \varphi}{\partial x_{v}}+(\cdots) \frac{\partial \varphi}{\partial y_{v}}=0 \tag{64.a}
\end{equation*}
$$

and we will see that:

This is always fulfilled identically, i.e., in the case of holonomic constraints, the demands that, on the one hand, the varied path and on the other, the transitional path should satisfy the constraints on the motion will be identical to each other, as we know already since the one demand inevitably implies the other.

However, (64) cannot be fulfilled, in general, since the coefficients $a_{\kappa v}, b_{\kappa v}, c_{\kappa \nu}$ do not vanish simultaneously, and likewise, the expressions in the parentheses cannot be simultaneously zero
due to non-holonomity. That is because the vanishing of those quantities would mean that the integrability conditions are fulfilled, which is contrary to the assumption of non-holonomity.

That then implies this result:
For non-holonomic equations of constraint, the true and the transitional motions indeed fulfill the equations of constraint, but not the varied path, in general.

Since, on the one hand, mechanical principles (e.g., d'Alembert's principle) demand that the transitional motion will be composed of virtual displacements, so the transitional motion must always fulfill the equations of constraint, if one would like to determine the displacements $\overline{\delta x_{v}}$, $\overline{\delta y_{v}}, \overline{\delta z_{v}}$ from that principle in such a way that the varied path fulfills the constraints on the motion then one would arrive at false equations of motion, in general. Both determinations will be equivalent only in the case of holonomic constraints. That distinction is not always stressed clearly enough, e.g., Hertz ${ }^{1}$ ) had violated it in a known example in his theory of mechanics.

The state of affairs might be explained once more with that Hertzian example. A ball rolls without slipping on a horizontal plane under the action of arbitrary forces. As we have already mentioned before, the equations of constraint in this problem that express rolling without slipping are non-holonomic. Two neighboring points ( $x_{v}, y_{v}, z_{v}$ ) and ( $x_{v}+d x_{v}, y_{v}+d y_{v}, z_{v}+d z_{v}$ ) along the true path naturally emerge from each other by a pure rolling motion.

The principles of mechanics further demand that the transitional motions that start from the point ( $x_{v}, y_{v}, z_{v}$ ), as well as the point ( $x_{v}+d x_{v}, y_{v}+d y_{v}, z_{v}+d z_{v}$ ), should lead to the respective final positions $\left(x_{v}+\delta x_{v}, y_{v}+\delta y_{v}, z_{v}+\delta z_{v}\right)$ and $\left(x_{v}+\delta x_{v}+d\left(x_{v}+\delta x_{v}\right), \ldots\right)$, i.e., that the displacements should be virtual. If that were the case then the points ( $x_{v}+\delta x_{v}, y_{v}+\delta y_{v}, z_{v}+$ $\left.\delta z_{v}\right)$ and $\left(x_{v}+\delta x_{v}+d x_{v}+d \delta x_{v}, y_{v}+\delta y_{v}+d y_{v}+d \delta y_{v}, \ldots\right)$ on the varied path would no longer emerge from each other by a pure rolling motion, but a slide would necessarily have to enter in.

Hölder ( ${ }^{2}$ ) has carried out the calculations in this example thoroughly and explained the actual situation.

[^1]
## § 6.

## More general variations (variation of time).

In the foregoing sections, we have already occasionally remarked that we can go from the true path to a neighboring varied one by way of more general types of displacements $\delta x_{v}, \delta y_{v}$, $\delta z$ than the virtual ones, i.e., ones that are performed at constant time. We would now like to consider that more general type in more detail. We would now like to associate a space-time point $\left(x_{v}, y_{v}, z_{v}, t\right)$, as we would like to say briefly, on the true path with the space-time point ( $x_{v}$ $+\Delta x_{v}, y_{v}+\Delta y_{v}, z_{v}+\Delta z_{v}, t+\Delta t$. The set of all of the latter points defines the varied path under our new rule for variation and for that we would like to expressly reserve the notation $\Delta$. We would again like to clarify the association of the two paths in this graphically by examining the relationship in the $x_{v} t$-plane, i.e., we shall consider the projection of the paths onto the $x$-axis.


Figure 2.
In Fig. 2, the point 1 means the point on the true path with the coordinates $x_{v}, y_{v}, z_{v}$ at time $t$. Up to now, we went from a point 1 to a point 2 (with the same value of $t$, i.e., by means of the operation $\delta$ ) that has the coordinates ( $x_{v}+\delta x_{v}, y_{v}+\delta y_{v}, z_{v}+\delta z_{v}$ ). The line segment (1,2) is accordingly equal to $\delta x_{v}$. However, we would now like to associate the point 1 with the point 4 that has time $t+\Delta t$ and coordinates $x_{v}+\Delta x_{v}, y_{v}+\Delta y_{v}, z_{v}+\Delta z_{v}$.

In addition, the point 3 on the true path is indicated in the figure, which is attained by the material point at time $t+\Delta t$. If we denote its velocity components by $\dot{x}_{v}, \dot{y}_{v}, \dot{z}_{v}$ then the point 3 will have the coordinates $x_{v}+\dot{x}_{v} \Delta t, y_{v}+\dot{y}_{v} \Delta t, z_{v}+\dot{z}_{v} \Delta t$. We will then get the following table for our four points:

| Point | Coordinates | Time |
| :---: | :---: | :--- |
| 1 | $x_{v}, y_{v}, z_{v}$ | $t$ |
| 2 | $x_{v}+\delta x_{v}, y_{v}+\delta y_{v}, z_{v}+\delta z_{v}$ | $t$ |
| 3 | $x_{v}+\dot{x}_{v} \Delta t, y_{v}+\dot{y}_{v} \Delta t, z_{v}+\dot{z}_{v} \Delta t$. | $t+\Delta t$ |
| 4 | $x_{v}+\Delta x_{v}, y_{v}+\Delta y_{v}, z_{v}+\Delta z_{v}$ | $t+\Delta t$ |

The difference between the ordinates of 2 and 1 is accordingly $\delta x_{v}$, and between 4 and 1 , it will be $\Delta x_{v}$.

We would like to exhibit the relationships between $\Delta x_{v}, \delta x_{v}$, and $\Delta t$. To that end, we observe that the point 4 emerges from 3 under the $\delta$ process, and likewise 2 from 1 . We can then express the $x$-coordinate of 4 as follows: $\left(x_{v}+\dot{x}_{v} \Delta t\right)+\delta\left(x_{v}+\dot{x}_{v} \Delta t\right)$. In general, we will then find the following values for the coordinates of 4 :

$$
x_{v}+\delta x_{v}+\dot{x}_{v} \Delta t, \quad y_{v}+\delta y_{v}+\dot{y}_{v} \Delta t, \quad z_{v}+\delta z_{v}+\dot{z}_{v} \Delta t
$$

when we neglect terms of higher order.
On the other hand, from the table above, its coordinates will be the following ones:

$$
x_{v}+\Delta x_{v}, \quad y_{v}+\Delta y_{v}, \quad z_{v}+\Delta z_{v} .
$$

A comparison will then give the following relations between $\Delta x_{v}, \delta x_{v}$, and $\Delta t$ :

$$
\left\{\begin{align*}
\Delta x_{v} & =\delta x_{v}+\dot{x}_{v} \Delta t  \tag{65}\\
\Delta y_{v} & =\delta y_{v}+\dot{y}_{v} \Delta t \\
\Delta z_{v} & =\delta z_{v}+\dot{z}_{v} \Delta t
\end{align*}\right.
$$

We would further like to apply the operations $\Delta$ and $\delta$ to a function $\varphi(x, y, z, t)$ and likewise exhibit the relations between $\Delta \varphi$ and $\delta \varphi$. From the definitions of the symbols, we have:

$$
\begin{align*}
& \Delta \varphi=\frac{\partial \varphi}{\partial x} \Delta x+\frac{\partial \varphi}{\partial y} \Delta y+\frac{\partial \varphi}{\partial z} \Delta z+\frac{\partial \varphi}{\partial t} \Delta t  \tag{66.a}\\
& \delta x=\frac{\partial \varphi}{\partial x} \delta x+\frac{\partial \varphi}{\partial y} \delta y+\frac{\partial \varphi}{\partial z} \delta z \tag{66.b}
\end{align*}
$$

If we add the time derivative to that:

$$
\begin{equation*}
\frac{d \varphi}{d t}=\dot{\varphi}=\frac{\partial \varphi}{\partial x} \dot{x}+\frac{\partial \varphi}{\partial y} \dot{y}+\frac{\partial \varphi}{\partial z} \dot{z}+\frac{\partial \varphi}{\partial t}, \tag{66.c}
\end{equation*}
$$

extend that equation by $\Delta t$, and subtract from (66.a) then we will have:

$$
\Delta \varphi-\dot{\varphi} \Delta t=\frac{\partial \varphi}{\partial x}(\Delta x-\dot{x} \Delta t)+\frac{\partial \varphi}{\partial y}(\Delta y-\dot{y} \Delta t)+\frac{\partial \varphi}{\partial z}(\Delta z-\dot{z} \Delta t),
$$

and from (65), the right-hand side of that is equal to:

$$
\frac{\partial \varphi}{\partial x} \delta x+\frac{\partial \varphi}{\partial y} \delta y+\frac{\partial \varphi}{\partial z} \delta z
$$

so from (66.b), it will be equal to $\delta \varphi$. We will then get:

$$
\begin{equation*}
\Delta \varphi=\delta \varphi+\dot{\varphi} \Delta t \tag{67}
\end{equation*}
$$

analogous to equation (65).

## § 7.

## General coordinates. True and non-holonomic coordinates.

The use of Cartesian coordinates is generally unsuitable, and for that reason one introduces so-called "general coordinates," i.e., quantities $q_{k}$ that are in one-to-one correspondence with the $x_{v}, y_{v}, z_{v}$. If holonomic equations of constraint, say $m$ of them, exist between the $3 n$ Cartesian coordinates $x_{v}, y_{v}, z_{v}$ then $(3 n-m)=N$ quantities $q_{k}$ can always be chosen such that no further equations of constraint will exist between them. If the equations of constraint were scleronomic then the relations that couple the $x_{v}, y_{v}, z_{v}$ to the $q_{k}$ would not include time $t$ explicitly, and one would also call the $q_{k}$ "scleronomic" coordinates, in a terminology that is easy to understand. By contrast, if the constraints that were originally present were rheonomic then time would also enter explicitly into the relations between the Cartesian coordinates and the $q_{k}$, which would then be called "rheonomic," as well. In both cases, we have, in any event, equations of the following type:
or

$$
\begin{align*}
& q_{k}=q_{k}\left(x_{1}, y_{1}, z_{1}, \ldots, x_{n}, y_{n}, z_{n}\right)  \tag{68.a}\\
& q_{k}=q_{k}\left(x_{1}, y_{1}, z_{1}, \ldots, x_{n}, y_{n}, z_{n}, t\right) .
\end{align*}
$$

In what follows, for the sake of simplicity, we will always restrict ourselves to scleronomic coordinates. Due to the form of the relation (68), the $q_{k}$ are also called holonomic or true coordinates. However, in some situations, it is preferable to introduce certain linear differential expressions in the true coordinates that we would like to abbreviate by $d^{\prime} \pi_{r}$ :

$$
\begin{equation*}
d^{\prime} \pi_{r}=\sum_{k=1}^{N} \alpha_{k r} d q_{k} \quad(r=1,2, \ldots, N) \tag{68}
\end{equation*}
$$

in which the $\alpha_{k r}$ are functions of the $q_{k}$.
If those coefficients $\alpha_{k r}$ satisfy the integrability conditions:

$$
\begin{equation*}
\frac{\partial \alpha_{k r}}{\partial q_{\lambda}}-\frac{\partial \alpha_{\lambda r}}{\partial q_{k}}=0 \quad(r=1,2, \ldots, N) \tag{69}
\end{equation*}
$$

then the $d^{\prime} \pi_{r}$ will be exact differentials, and the quantities $\pi_{r}$ can be obtained as functions of the $q_{k}$ from (68) by integration. In that case, the $\pi_{r}$ will be merely new general coordinates of the same character as the $q_{k}$. Naturally, the prime on the differential symbol in (68) can then be omitted in that case. By contrast, if the integrability conditions (69) are not fulfilled then there will exist no such quantities as " $\pi_{r}$ ", rather, the $d^{\prime} \pi_{r}$ will just be defined by (68). If one divides
(68) by $d t$ then the quantities $d q_{k} / d t=\dot{q}_{k}$ will appear on the right-hand side, namely, the generalized velocities. We would not like to denote the quantities $d^{\prime} \pi_{r} / d t$ on the left-hand side by $\dot{\pi}_{r}$, in order to not create the impression that we are dealing with time derivatives of the quantities " $\pi_{r}$ ". We shall denote the aforementioned quotients by $\pi_{r}$, to distinguish them. They obey the equations:

$$
\begin{equation*}
\dot{\pi}_{r}=\sum_{k=1}^{N} \alpha_{k r} \dot{q}_{k} \quad(r=1, \ldots, N), \tag{70}
\end{equation*}
$$

so they are linear combinations of the generalized velocity components and can even be employed as such generalized velocities. However, they differ from the $\dot{q}_{k}$ by the fact that the latter are derivatives of quantities $q_{k}$, while the former are not. Naturally, for that reason, one cannot employ the " $\pi_{r}$ " as generalized coordinates either. Only the $q_{k}$ and the $\dot{q}_{k}$ will appear together in the Lagrange equations of motion that will be discussed later. The difference between the $\dot{q}_{k}$ and the $\pi_{r}$ that was explained here makes it plausible from the outset that the quantities $\pi_{r}$ cannot be used as velocity components in the Lagrange equations since the associated quantities $\pi_{r}$ do not even exist. However, as we will explain later, the Lagrange equations can be extended in such a way that they will also remain useful in that case. One calls the quantities $d^{\prime} \pi_{r}$ (although it is not entirely appropriate) differentials of non-holonomic coordinates or quasi-coordinates $\left({ }^{1}\right)$. An example of those quasi-coordinates is the following one: As is known, one can determine the position of a rigid body that is fixed at a point by the socalled Euler angles $\varphi, \psi, \vartheta$, which are obviously true coordinates. $\varphi$ is called the angle of precession, $\psi$ is the angle of proper rotation, and $\vartheta$ is the pendulum angle. The quantities $\dot{\varphi}, \dot{\psi}$, $\dot{\vartheta}$ are correspondingly called velocity components. On the other hand, one can decompose the angular velocity of the rigid body around the instantaneous rotational axis into three components $\stackrel{*}{\pi}, \stackrel{*}{\chi}$, along three axes that are fixed in the body. They are naturally functions of the $\dot{\varphi}, \dot{\psi}, \dot{\vartheta}$. Indeed, the mechanics of rigid bodies tells us that the following relations will exist:

$$
\left\{\begin{array}{l}
\stackrel{*}{\pi}=\sin \psi \sin \vartheta \cdot \dot{\varphi}+\cos \psi \cdot \dot{\vartheta}+0 \cdot \dot{\psi}  \tag{71}\\
\stackrel{*}{\chi}=\cos \psi \sin \vartheta \cdot \dot{\varphi}-\sin \psi \cdot \dot{\vartheta}+0 \cdot \dot{\psi} \\
\stackrel{*}{\rho}=\cos \vartheta \cdot \dot{\varphi}+0 \quad \cdot \dot{\vartheta}+1 \cdot \dot{\psi}
\end{array}\right.
$$

[^2]In this case, the $\alpha_{k r}$ have the following values:
and they do not fulfill the integrability conditions (69). For example, one has:

$$
\frac{\partial \alpha_{11}}{\partial \vartheta}-\frac{\partial \alpha_{21}}{\partial \varphi}=\frac{\partial}{\partial \vartheta}(\sin \psi \sin \vartheta)-\frac{\partial}{\partial \varphi}(\cos \psi)=\sin \psi \cos \vartheta \neq 0
$$

If we multiply equations (71) by $d t$ then the quantities ${ }^{*} d t,{ }^{*} \chi d t, \stackrel{*}{\rho} d t$ will be inexact differentials that we would like to denote by $d^{\prime} \pi, d^{\prime} \chi, d^{\prime} \rho$ :

$$
\left\{\begin{array}{l}
d^{\prime} \pi=\sin \psi \sin \vartheta \cdot d \varphi+\cos \psi \cdot d \vartheta+0 \cdot d \psi  \tag{73}\\
d^{\prime} \chi=\cos \psi \sin \vartheta \cdot d \varphi-\sin \psi \cdot d \vartheta+0 \cdot d \psi \\
d^{\prime} \rho=\cos \vartheta \cdot d \varphi+0 \quad \cdot d \vartheta+1 \cdot d \psi
\end{array}\right.
$$

The quantities $d^{\prime} \pi, \ldots$ have precisely the property that we have just demanded of the differentials of quasi-coordinates. We shall return to this example in more detail later.

In what follows, we shall concern ourselves with an important property of the differentials of non-holonomic coordinates upon which precisely the aforementioned inapplicability of the ordinary Lagrange equations is based, in the final analysis.


Figure 3.a
Let $q_{k}$ be a true coordinate and let $\dot{q}_{k}$ be the associated velocity. We now treat the relationship between the expressions $\delta\left(\frac{d q_{k}}{d t}\right)$ and $\frac{d}{d t}\left(\delta q_{k}\right)$. We draw the curve $q_{k}(t)$ in the $q_{k} t$ -
plane. At the same time, we denote a neighboring curve by $q_{k}^{\prime}(t)$. If we associate both curves with each other in such a way that we let each point $q_{k}$ corresponds to the point $q_{k}^{\prime}$ that belongs to the same value of $t$, as is suggested in Fig. 3.a, then by the definition of the $\delta$ symbol, we will obviously have $q_{k}^{\prime}-q_{k}=\delta q_{k}$. Hence:

$$
\begin{equation*}
\frac{d}{d t}\left(\delta q_{k}\right)=\frac{d q_{k}^{\prime}}{d t}-\frac{d q_{k}}{d t}=\dot{q}_{k}^{\prime}-\dot{q}_{k} \tag{74}
\end{equation*}
$$



Figure 3.b
If we now draw the curves $\dot{q}_{k}(t)$ and $\dot{q}_{k}^{\prime}(t)$ in the $\dot{q}_{k} t$-plane then from the definition of the $\delta$ symbol, we will once more have:

$$
\begin{equation*}
\delta \dot{q}_{k}=\delta\left(\frac{d q_{k}}{d t}\right)=\dot{q}_{k}^{\prime}-\dot{q}_{k} \tag{75}
\end{equation*}
$$

as we read off directly from Fig. 3.b, and a comparison of (74) and (75) will yield the fundamental relation:

$$
\begin{equation*}
\frac{d}{d t}\left(\delta q_{k}\right)=\delta\left(\frac{d q_{k}}{d t}\right) \tag{76}
\end{equation*}
$$

i.e., the operations $\delta$ and $d / d t$ will commute for holonomic coordinates. One sees that this is essentially based upon the fact that $t$, the independent variable, remains unvaried. The commutability will not be true for the symbols $\Delta$ and $d / d t$, as one sees immediately.

Let non-holonomic coordinates be defined by the equations:

$$
\begin{equation*}
\delta \pi_{r}=\sum_{k=1}^{N} \alpha_{k r} \delta q_{k} \quad(r=1, \ldots, N) \tag{77}
\end{equation*}
$$

and the associated velocities by:

$$
\begin{equation*}
\stackrel{*}{\pi}_{r}=\sum_{k=1}^{N} \alpha_{k r} \dot{q}_{k}, \tag{78}
\end{equation*}
$$

respectively.
If we now take the time derivative of (77) and apply the operation $\delta$ to (78) then we will find that:

$$
\begin{aligned}
\frac{d}{d t}\left(\delta \pi_{r}\right) & =\sum_{k=1}^{N} \alpha_{k r} \frac{d \delta q_{k}}{d t}+\sum_{k=1}^{N} \frac{d \alpha_{k r}}{d t} \delta q_{k} \\
\delta \pi_{r}^{*} & =\sum_{k=1}^{N} \alpha_{k r} \delta\left(\frac{d q_{k}}{d t}\right)+\sum_{k=1}^{N} \delta \alpha_{k r} \cdot \frac{d q_{k}}{d t}
\end{aligned}
$$

Since one has $\frac{d \delta q_{k}}{d t}=\delta \frac{d q_{k}}{d t}$, from (76), subtraction will yield:

$$
\begin{equation*}
\delta{\stackrel{*}{\pi_{r}}}_{r} \frac{d}{d t}\left(\delta \pi_{r}\right)=\sum_{k}\left\{\delta \alpha_{k r} \dot{q}_{k}-\dot{\alpha}_{k r} \delta q_{k}\right\} . \tag{79}
\end{equation*}
$$

As calculation will show, the right-hand side of that will be non-zero for non-holonomic coordinates, i.e., the two operations in question do not commute. In order to make that obvious, we write:

$$
\dot{\alpha}_{k r}=\sum_{\lambda} \frac{\partial \alpha_{k r}}{\partial q_{\lambda}} \dot{q}_{\lambda}
$$

and when that is substituted in (79), that will produce:

$$
\delta \pi_{r}^{*}-\frac{d}{d t}\left(\delta \pi_{r}\right)=\sum_{k} \delta \alpha_{k r} \dot{q}_{k}-\sum_{k} \sum_{\lambda} \frac{\partial \alpha_{k r}}{\partial q_{\lambda}} \delta q_{k} \dot{q}_{\lambda}
$$

or when we permute the summation indices $k$ and $\lambda$ in the double sum (which is obviously permissible):

$$
\delta \stackrel{*}{\pi}_{r}-\frac{d}{d t}\left(\delta \pi_{r}\right)=\sum_{k} \dot{q}_{k}\left(\delta \alpha_{k r}-\sum_{k} \sum_{\lambda} \frac{\partial \alpha_{k r}}{\partial q_{\lambda}} \delta q_{k}\right) .
$$

Likewise, we have:

$$
\delta \alpha_{k r}=\sum_{\lambda} \frac{\partial \alpha_{k r}}{\partial q_{\lambda}} \delta q_{\lambda}
$$

so when that is substituted, we will ultimately have:

$$
\begin{equation*}
\delta \pi_{r}-\frac{d \delta \pi_{r}}{d t}=\sum_{k} \sum_{\lambda}\left\{\frac{\partial \alpha_{k r}}{\partial q_{\lambda}}-\frac{\partial \alpha_{\lambda r}}{\partial q_{k}}\right\} \dot{q}_{k} \delta q_{k} \tag{80}
\end{equation*}
$$

The expression in brackets $\left\{\frac{\partial \alpha_{k r}}{\partial q_{\lambda}}-\frac{\partial \alpha_{\lambda r}}{\partial q_{k}}\right\}$ that appears on the right is nothing but the lefthand side of the integrability equation (69), and since that is not fulfilled for non-holonomic coordinates, we will in fact have $\delta \pi_{r} \neq \frac{d \delta \pi_{r}}{d t}$. However, it is precisely the commutability of $\delta$ and $d / d t$ that is required in the derivation of the Lagrange equations, as we will explain more precisely later.

We can modify equation (80) somewhat. Namely, if we solve equation (78) for $\dot{q}_{k}$ and (77) for $\delta q_{k}$ then we can replace $\dot{q}_{k}$ and $\delta q_{\lambda}$ with $\stackrel{*}{r}_{r}$ and $\delta \pi_{r}$, resp., on the right-hand side of (80). Let the solution of the aforementioned equations be the following:

$$
\left\{\begin{align*}
\delta q_{\lambda} & =\sum_{\rho} \beta_{\rho \lambda} \delta \pi_{\rho},  \tag{81}\\
\dot{q}_{k} & =\sum_{\sigma} \beta_{\sigma \lambda} \pi_{\sigma}^{*}
\end{align*}\right.
$$

in which $\rho$ and $\sigma$ are two summation indices. (80) will then become:

$$
\begin{equation*}
\delta\left(\pi_{r}^{*}\right)=\frac{d}{d t}\left(\delta \pi_{r}\right)+\sum_{k} \sum_{\lambda} \sum_{\rho} \sum_{\sigma}\left\{\frac{\partial \alpha_{k r}}{\partial q_{\lambda}}-\frac{\partial \alpha_{\lambda r}}{\partial q_{k}}\right\} \beta_{\rho \lambda} \beta_{\sigma k} \pi_{\sigma}^{*} \delta \pi_{\rho}, \tag{82}
\end{equation*}
$$

or when one permutes $\rho$ and $\sigma$ :

$$
\begin{equation*}
\delta\left(\pi_{r}^{*}\right)=\frac{d}{d t}\left(\delta \pi_{r}\right)+\sum_{\sigma} \sum_{\rho}\left[\sum_{k} \sum_{\lambda} \beta_{\rho k} \beta_{\sigma \lambda}\left\{\frac{\partial \alpha_{k r}}{\partial q_{\lambda}}-\frac{\partial \alpha_{\lambda r}}{\partial q_{k}}\right\}\right] \pi_{\rho}^{*} \cdot \delta \pi_{\sigma} . \tag{82.a}
\end{equation*}
$$

One can make a slight alteration to (82.a).
The square bracket depends upon the indices $k, r, \lambda, \rho, \sigma$, of which those two symbols $k$ and $\lambda$ in the double summation that is to be performed are summation indices, while the indices $r, \rho$, $\sigma$ are running variables. For that reason, we would like to combine them into the notation ( $-\gamma_{r \rho \sigma}$ ) and then get:

$$
\begin{equation*}
\delta\left(\stackrel{*}{\pi}_{r}\right)=\frac{d}{d t}\left(\delta \pi_{r}\right)+\sum_{\sigma} \sum_{\rho} \gamma_{r \rho \sigma} \stackrel{*}{\pi}_{\rho} \delta \pi_{\sigma} \tag{83}
\end{equation*}
$$

The summarizing of the square bracket into $-\gamma_{r \rho \sigma}$ is justified by the fact that $\gamma_{r \rho \sigma}$ depends upon only the mutual couplings between the $\delta \pi_{r}$ and the $\delta q_{k}$.

## § 8.

## Hamilton's principle of stationary action and its equivalence with d'Alembert's principle.

The left-hand side of equation (37) includes d'Alembert's principle in the form:

$$
\sum_{v=1}^{n}\left[\left(X_{v}-m_{v} \ddot{x}_{v}\right) \delta x_{v}+\left(Y_{v}-m_{v} \ddot{y}_{v}\right) \delta y_{v}+\left(Z_{v}-m_{v} \ddot{z}_{v}\right) \delta z_{v}\right] .
$$

That expression will vanish when $\delta x_{v}, \delta y_{v}, \delta z_{v}$ are virtual displacements. Naturally, it will be non-zero when that is not the case, in general. In what follows, we would like to leave it undecided whether the $\delta x_{v}, \delta y_{v}, \delta z_{v}$ are or are not virtual displacements. For our initial considerations, it will not matter whether the expression above vanishes, but only when in combination. We convert it into the following form:

$$
\begin{equation*}
\sum_{v}\left(X_{v} \delta x_{v}+Y_{v} \delta y_{v}+Z_{v} \delta z_{v}\right)-\sum_{v} m_{v}\left(\ddot{x}_{v} \delta x_{v}+\ddot{y}_{v} \delta y_{v}+\ddot{z}_{v} \delta z_{v}\right) \tag{84}
\end{equation*}
$$

As is known, the first summand:

$$
\begin{equation*}
\sum_{v}\left(X_{v} \delta x_{v}+Y_{v} \delta y_{v}+Z_{v} \delta z_{v}\right)=\delta^{\prime} A \tag{85}
\end{equation*}
$$

represents the work done by the forces $X_{v}, Y_{v}, Z_{v}$ under the arbitrary displacement $\delta x_{v}, \ldots$ The second sum can be converted term-by-term in the following way:

$$
\begin{aligned}
m_{v} \ddot{x}_{v} \delta x_{v} & =\frac{d}{d t}\left(m_{v} \dot{x}_{v} \delta x_{v}\right)-m_{v} \dot{x}_{v} \frac{d \delta x_{v}}{d t} \\
& =\frac{d}{d t}\left(m_{v} \dot{x}_{v} \delta x_{v}\right)-m_{v} \dot{x}_{v} \delta \dot{x}_{v} \\
& =\frac{d}{d t}\left(m_{v} \dot{x}_{v} \delta x_{v}\right)-\delta\left(\frac{1}{2} m_{v} \dot{x}_{v}^{2}\right) .
\end{aligned}
$$

We will then have:

$$
m_{v}\left(\ddot{x}_{v} \delta x_{v}+\ddot{y}_{v} \delta y_{v}+\ddot{z}_{v} \delta z_{v}\right)=\frac{d}{d t}\left[m_{v}\left(\dot{x}_{v} \delta x_{v}+\dot{y}_{v} \delta y_{v}+\dot{z}_{v} \delta z_{v}\right)\right]-\delta \frac{m_{v}}{2}\left(\dot{x}_{v}^{2}+\dot{y}_{v}^{2}+\dot{z}_{v}^{2}\right) .
$$

The last term is the variation of the kinetic energy $L_{v}$ of the $v^{\text {th }}$ mass-point then. Therefore, upon summing over all mass-points of the system, one will ultimately have:

$$
\begin{equation*}
m_{v}\left(\ddot{x}_{v} \delta x_{v}+\ddot{y}_{v} \delta y_{v}+\ddot{z}_{v} \delta z_{v}\right)=-\delta L+\frac{d}{d t}\left[\sum_{v} m_{v}\left(\dot{x}_{v} \delta x_{v}+\dot{y}_{v} \delta y_{v}+\dot{z}_{v} \delta z_{v}\right)\right] . \tag{86}
\end{equation*}
$$

Subtracting equation (86) from (85) will then produce the following identity:

$$
\begin{equation*}
\sum_{v=1}^{n}\left[\left(X_{v}-m_{v} \ddot{x}_{v}\right) \delta x_{v}+\cdots\right]=\delta L+\delta^{\prime} A-\frac{d}{d t}\left[\sum_{v} m_{v}\left(\dot{x}_{v} \delta x_{v}+\dot{y}_{v} \delta y_{v}+\dot{z}_{v} \delta z_{v}\right)\right] . \tag{87}
\end{equation*}
$$

Here, the d'Alembert expression on the left-hand side is transformed with no concern for whether the $\delta x_{v}, \ldots$ are or are not virtual displacements. Since a differential quotient with respect to time is found on the right-hand side, that suggests performing an integration over $t$ between two fixed limits $t_{0}$ and $t_{1}$. That immediately yields:

$$
\int_{t_{0}}^{t_{1}} \sum_{v=1}^{n}\left[\left(X_{v}-m_{v} \ddot{x}_{v}\right) \delta x_{v}+\cdots\right] d t=\int_{t_{0}}^{t_{1}}\left(\delta L+\delta^{\prime} A\right) d t-\left[\sum_{v} m_{v}\left(\dot{x}_{v} \delta x_{v}+\dot{y}_{v} \delta y_{v}+\dot{z}_{v} \delta z_{v}\right)\right]_{t_{0}}^{t_{1}}
$$

The term outside of the integral sign can now be made to vanish by a suitable assumption on the quantities $\delta x_{v}, \ldots$ at the times $t_{0}$ and $t_{1}$. Namely, if we assume from now on that the $\delta x_{v}, \delta y_{v}, \delta z_{v}$ are equal to zero at the two times $t_{0}$ and $t_{1}$ (the physical meaning of this will be discussed later) then the expression in question will vanish at the two limits, and will get simply:

$$
\begin{equation*}
\int_{t_{0}}^{t_{1}} \sum_{v=1}^{n}\left[\left(X_{v}-m_{v} \ddot{x}_{v}\right) \delta x_{v}+\cdots\right] d t=\int_{t_{0}}^{t_{1}}\left(\delta L+\delta^{\prime} A\right) d t \tag{88}
\end{equation*}
$$

Now, if displacements that were assumed to be arbitrary up to now are virtual then the lefthand side will vanish on the grounds of d'Alembert's principle, so the right-hand side will, as well, and the statement that:

$$
\begin{equation*}
\int_{t_{0}}^{t_{1}}\left(\delta L+\delta^{\prime} A\right) d t=0 \tag{89}
\end{equation*}
$$

will be completely equivalent in its content to d'Alembert's principle and the equations of motion. Equations (89) is called Hamilton's principle of stationary action.

In the special case where the forces can be derived from a potential $\left(X_{v}=-\frac{\partial \Phi\left(x_{1}, \ldots, x_{n}\right)}{\partial x_{v}}\right.$, $\ldots$..), one has $\delta^{\prime} A=-\delta \Phi$, and when one switches the order of variation $\delta$ and integration over $t$, one can write:

$$
\begin{equation*}
\delta \int_{t_{0}}^{t_{1}}(L-\Phi) d t=\delta \int_{t_{0}}^{t_{1}} H d t=0 \tag{90}
\end{equation*}
$$

According to Helmholtz, $H=L-\Phi$ is called the kinetic potential.
In the form (90), Hamilton's principle of stationary action will admit a simple formulation: namely, except for an irrelevant constant factor, $\int_{t_{0}}^{t_{1}} H d t$ is the temporal mean of the kinetic potential. We can then summarize the equations of motion in the simple statement:

The temporal mean of the kinetic potential is stationary for the actual motion, i.e., a maximum, a minimum, or a saddle point value.

Therefore, the actual motion (we shall again restrict our notation to the projection onto the xtplane) will be compared to those neighboring (i.e., varied) paths that have the same initial and final points at times $t_{0}$ and $t_{1}$ as the actual one, since the $\delta x_{v}, \ldots$ should indeed vanish at those two times. The transition from the actual to a varied path must always take place by way of a virtual displacement. If non-holonomic constraints are present then from the foregoing the varied motions that one compares to the actual motion will represent impossible motions since they will then contradict the equations of motion.


Figure 4.
If we were to demand (but this is not based upon facts) that the varied paths should also obey the non-holonomic constraints then that would exclude the use of Hamilton's principle since it would then produce false equations of motion for them. H. Hertz ${ }^{( }{ }^{1}$ ) was of that opinion, although he incorrectly believed that Hamilton's principle should be restricted to holonomic constraints, simply because demanded that the varied path should be compatible with the constraints, whereas the principles of mechanics indicate that this should be necessary for the

[^3]transitional path. In reality, as our derivation shows, Hamilton's principle is true for arbitrary constraints, precisely as d'Alembert was.

If one would like to employ Hamilton's principle in order to arrive at the equations of motion then one must merely follow the sequence of equations (89) to (84) in reverse order. In the next section, we will treat the derivation of the Lagrange equations from the principle of stationary action in that way.

## § 9.

## Lagrange's equations for holonomic and non-holonomic coordinates.

If we introduce general holonomic coordinates $q_{k}$ and $\dot{q}_{k}$ into $L$ and the expression $\delta^{\prime} A$, instead of Cartesian coordinates (which is how they were expressed up to now), then, as is known, $L$ will become a homogeneous quadratic form in the $\dot{q}_{k}$ whose coefficients are functions of the $q_{k}$. One can then say in full generality that $L$ depends upon the $q_{k}$ and $\dot{q}_{k}$ then:

$$
L=L\left(q_{k}, \dot{q}_{k}\right),
$$

and variation will then yield:

$$
\begin{equation*}
\delta L=\sum_{k=1}^{N} \frac{\partial L}{\partial q_{k}} \delta q_{k}+\sum_{k=1}^{N} \frac{\partial L}{\partial \dot{q}_{k}} \delta \dot{q}_{k} . \tag{91}
\end{equation*}
$$

The $\delta x_{v}, \delta y_{v}, \delta z_{v}$ enter into the expression $\delta^{\prime} A$, which are linear combinations of the $\delta q_{k}$ that might take the forms:

$$
\begin{aligned}
& \delta x_{\nu}=\sum_{k} \alpha_{k}^{v} \delta q_{k}, \\
& \delta y_{v}=\sum_{k} \beta_{k}^{v} \delta q_{k} \\
& \delta z_{v}=\sum_{k} \gamma_{k}^{v} \delta q_{k} .
\end{aligned}
$$

$X_{v} \delta x_{v}+Y_{v} \delta y_{v}+Z_{v} \delta z_{v}$ will then go to:

$$
X_{v} \delta x_{v}+Y_{v} \delta y_{v}+Z_{v} \delta z_{v}=X_{v} \sum_{k} \alpha_{k}^{v} \delta q_{k}+Y_{v} \sum_{k} \beta_{k}^{v} \delta q_{k}+Z_{v} \sum_{k} \gamma_{k}^{v} \delta q_{k},
$$

so ultimately $\delta^{\prime} A=\sum_{v=1}^{n}\left(X_{v} \delta x_{v}+\cdots\right)$ will go to:

$$
\delta^{\prime} A=\sum_{v=1}^{n} \sum_{k=1}^{N}\left(X_{v} \alpha_{k}^{v}+Y_{v} \beta_{k}^{v}+Z_{v} \gamma_{k}^{v}\right) \delta q_{k}=\sum_{k=1}^{N} \delta q_{k}\left\{\sum_{v=1}^{n}\left(X_{v} \alpha_{k}^{v}+Y_{v} \beta_{k}^{v}+Z_{v} \gamma_{k}^{v}\right)\right\}=\sum_{k=1}^{N} Q_{k} \delta q_{k} .
$$

The $Q_{k}$ in that is an abbreviation for $\sum_{v=1}^{n}\left(X_{v} \alpha_{k}^{v}+Y_{v} \beta_{k}^{v}+Z_{v} \gamma_{k}^{v}\right)$.
Its physical meaning is deduced from the fact that when it is multiplied by $\delta q_{k}$, it will have the dimension of work. Accordingly, $Q_{k}$ is a parallel to $X_{v}$, which also indeed produces a work when multiplied by $\delta x_{v}$, and for that reason, $Q_{k}$ will be referred to as the generalized force
component that is produced by the variation $\delta q_{k}$ of a coordinate $q_{k}$, while the remaining $q_{k}$ are held constant. One can therefore always write:

$$
\begin{equation*}
\delta^{\prime} A=\sum_{k=1}^{N} Q_{k} \delta q_{k}, \tag{92}
\end{equation*}
$$

so in the special case where a potential $\Phi$ exists, one will naturally have $Q_{k}=-\frac{\partial \Phi}{\partial q_{k}}$, so $\delta^{\prime} A=$ $-\delta \Phi=-\sum_{k} \frac{\partial \Phi}{\partial q_{k}} \delta q_{k}$.

If we substitute (91) and (92) in the principle of stationary action [equation (89)] then we will get:

$$
\begin{equation*}
\int_{t_{0}}^{t_{1}} d t \sum_{k} \frac{\partial L}{\partial q_{k}} \delta q_{k}+\int_{t_{0}}^{t_{1}} d t \sum_{k} \frac{\partial L}{\partial \dot{q}_{k}} \delta \dot{q}_{k}+\int_{t_{0}}^{t_{1}} d t \sum_{k} Q_{k} \delta q_{k}=0 . \tag{93}
\end{equation*}
$$

The second term in that can be converted by partial integration, in which we must make use of the commutability of $\delta$ and $d / d t$, which is true for only holonomic coordinates. That term will become, in turn (we drop the summation sign in the intermediate calculations, for simplicity):

$$
\int d t \frac{\partial L}{\partial \dot{q}_{k}} \delta \dot{q}_{k}=\int \frac{\partial L}{\partial \dot{q}_{k}} \frac{d \delta q_{k}}{d t} d t=\left[\frac{\partial L}{\partial \dot{q}_{k}} \delta q_{k}\right]_{t_{0}}^{t_{1}}-\int \frac{d}{d t}\left(\frac{\partial L}{\partial \dot{q}_{k}}\right) \delta q_{k} d t .
$$

Since the $\delta x_{v}, \delta y_{v}, \delta z_{v}$ vanish for $t_{0}$ and $t_{1}$, the same will be true for the $\delta q_{k}$, so the term outside of the integral sign will vanish, and equation (93) will then become:

$$
\begin{equation*}
\int_{t_{0}}^{t_{1}} d t \sum_{k=1}^{N}\left\{\frac{\partial L}{\partial q_{k}}-\frac{d}{d t}\left(\frac{\partial L}{\partial \dot{q}_{k}}\right)+Q_{k}\right\} \delta q_{k}=0 . \tag{94}
\end{equation*}
$$

Since the times $t_{0}$ and $t_{1}$ are completely arbitrary, but equation (94) is valid in full generality, the integral can vanish only when the integrand is annulled, and due to the complete independence of the $N$ quantities $\delta q_{k}$, it will decompose into $N$ independent equations:

$$
\begin{equation*}
\frac{\partial L}{\partial q_{k}}-\frac{d}{d t}\left(\frac{\partial L}{\partial \dot{q}_{k}}\right)=Q_{k} \quad(k=1,2, \ldots, N) . \tag{95}
\end{equation*}
$$

Those are the known Lagrange equations of the second kind, whose validity, as one sees, depends essentially upon the commutability of $\delta$ and $d / d t$, i.e., upon the use of holonomic coordinates.

We would now like to carry out the same calculation $\left({ }^{1}\right)$, but under the assumption that we now express the $\delta q_{k}$ and $\dot{q}_{k}$ as non-holonomic coordinate differentials $\delta \pi_{r}$ and the associated velocities $\stackrel{*}{\pi}_{r}$, resp. In that way, $L\left(q_{k}, \dot{q}_{k}\right)$ might go to $L^{\prime}\left(q_{k}, \dot{\pi}_{r}\right)$. If one replaces the $\delta q_{k}$ in the given way in the expression $\sum_{k} Q_{k} \delta q_{k}$ then $\delta^{\prime} A$ can obviously be put into the form:

$$
\begin{equation*}
\delta^{\prime} A=\sum_{r=1}^{N} \Pi_{r} \delta \pi_{r} \tag{96}
\end{equation*}
$$

then, in which the $\Pi_{r}$ is the generalized force component that generates the displacement $\delta \pi_{r}$, while all other $\delta \pi$ are equal to zero.

In Hamilton's principle, we will then have:

$$
\int_{t_{0}}^{t_{1}} d t \delta L^{\prime}+\int_{t_{0}}^{t_{1}} d t \sum_{r=1}^{N} \Pi_{r} \delta \pi_{r}=0
$$

or, since we have $\delta L^{\prime}=\sum \frac{\partial L^{\prime}}{\partial q_{k}} \delta q_{k}+\sum_{r} \frac{\partial L^{\prime}}{\partial{ }_{*}^{*}} \delta \stackrel{*}{r}_{r}$ :

$$
\begin{equation*}
\int_{t_{0}}^{t_{1}} d t \sum_{k} \frac{\partial L^{\prime}}{\partial q_{k}} \delta q_{k}+\int_{t_{0}}^{t_{1}} d t \sum_{r} \frac{\partial L^{\prime}}{\partial \pi_{r}^{*}} \delta \stackrel{*}{\pi}_{r}+\int_{t_{0}}^{t_{1}} d t \sum_{r=1}^{N} \Pi_{r} \delta \pi_{r}=0 \tag{97}
\end{equation*}
$$

We convert all terms in such a way that they will contain the factor $\delta \pi_{r}$. We start with the first one. From equation (81), one has (while changing only the summation sign):

$$
\begin{equation*}
\delta q_{k}=\sum_{r=1}^{N} \beta_{r k} \delta \pi_{r} \tag{98}
\end{equation*}
$$

As a result, the first term can be written:

$$
\int_{t_{0}}^{t_{1}} d t \sum_{r=1}^{N}\left\{\sum_{k=1}^{N} \frac{\partial L^{\prime}}{\partial q_{k}} \beta_{r k}\right\} \delta \pi_{r} .
$$

Now, if the " $\pi_{r}$ " were holonomic coordinates, so from (98), the $\beta_{r k}$ would be the partial derivatives $\partial q_{k} / \partial \pi_{r}$, then one would have:

$$
\sum_{k} \frac{\partial L^{\prime}}{\partial q_{k}} \beta_{r k}=\sum_{k} \frac{\partial L^{\prime}}{\partial q_{k}} \frac{\partial q_{k}}{\partial \pi_{r}}=\frac{\partial L^{\prime}}{\partial \pi_{r}}
$$

${ }^{(1)}$ For this, cf., C. Schaefer, Phys. Zeit. 19 (1918), pp. 406.

We cannot use this notation here (or only in a figurative sense), so we must then make it known that this is an improper notation by enclosing it in parentheses. Thus, we will have:

$$
\left(\frac{\partial L^{\prime}}{\partial \pi_{r}}\right) \quad \text { as an abbreviation for: } \quad \sum_{k} \frac{\partial L^{\prime}}{\partial q_{k}} \beta_{r k}
$$

With that, we can write the first term as:

$$
\begin{equation*}
\int_{t_{0}}^{t_{1}} d t \sum_{k=1}^{N} \frac{\partial L^{\prime}}{\partial q_{k}} \delta q_{k}=\int_{t_{0}}^{t_{1}} d t \sum_{r=1}^{N}\left(\frac{\partial L^{\prime}}{\partial \pi_{r}}\right) \delta \pi_{r} . \tag{99}
\end{equation*}
$$

In the second term, since the commutation of $\delta$ and $d / d t$ is no longer allowed, we must employ the general equation (83):

$$
\int_{t_{0}}^{t_{1}} d t \sum_{r} \frac{\partial L^{\prime}}{\partial \pi_{r}} \delta \stackrel{*}{\pi_{r}}=\int d t \sum_{r} \frac{\partial L^{\prime}}{\partial \pi_{r}} \frac{d \delta \pi_{r}}{d t}-\int d t \sum_{r} \frac{\partial L^{\prime}}{\partial \pi_{r}^{*}} \sum_{\rho} \sum_{\sigma} \gamma_{r \rho \sigma} \stackrel{*}{\pi}_{\rho} \delta \pi_{\sigma} .
$$

Since the $\delta \pi_{r}$ will vanish at the limits of the integral, and because it is true for the $\delta q_{k}$, the first term on the right in that will produce the value $-\int d t \sum_{r} \frac{d}{d t}\left(\frac{\partial L^{\prime}}{\partial \pi_{r}}\right) \delta \pi_{r}$ by ordinary partial integration. We commute the summation signs for $r$ and $\sigma$ in the second term on the right, in order to get $\delta \pi_{r}$, instead of $\delta \pi_{\sigma}$, as in the other expressions. That will then yield:

$$
\begin{equation*}
\int d t \sum_{r} \frac{\partial L^{\prime}}{\partial \pi_{r}} \delta \pi_{r}^{*}=-\int d t \sum_{r} \frac{d}{d t}\left(\frac{\partial L^{\prime}}{\partial \pi_{r}}\right) \delta \pi_{r}-\int d t \sum_{r} \sum_{\rho} \sum_{\sigma} \frac{\partial L^{\prime}}{\partial \pi_{r}} \gamma_{r \rho \sigma} \stackrel{*}{r}_{r} \delta \pi_{r} . \tag{100}
\end{equation*}
$$

If one substitutes (99) and (100) in the starting equation (97) then it will follow that:

$$
\int_{t_{0}}^{t_{1}} \sum_{r=1}^{N}\left[\left(\frac{\partial L^{\prime}}{\partial \pi_{r}}\right)-\frac{d}{d t}\left(\frac{\partial L^{\prime}}{\partial \pi_{r}}\right)-\sum_{\rho} \sum_{\sigma} \gamma_{r \rho \sigma} \frac{\partial L^{\prime}}{\partial \pi_{\sigma}^{*}} \pi_{\sigma}^{*}+\Pi_{r}\right] \delta \pi_{r}=0,
$$

from which it will follow by the usual argument that was also used above (viz., the $\delta \pi_{r}$ are mutually independent, just like the $\delta q_{k}$ ) that $\left({ }^{1}\right)$ :

[^4]\[

$$
\begin{equation*}
\frac{d}{d t}\left(\frac{\partial L^{\prime}}{\partial \pi_{r}}\right)-\left(\frac{\partial L^{\prime}}{\partial \pi_{r}}\right)+\sum_{\rho} \sum_{\sigma} \gamma_{r \rho \sigma} \frac{\partial L^{\prime}}{\partial \pi_{\sigma}^{*}} \pi_{\sigma}^{*}=\Pi_{r} \quad(r=1,2, \ldots, N) \tag{101}
\end{equation*}
$$

\]

Since the $\gamma_{r \rho \sigma}$ are proportional to the left-hand sides of the integrability conditions (69), the equation above will reduce immediately to the ordinary Lagrange one when the coordinates are holonomic, as it must. Equation (101) is the form of it that is valid for non-holonomic coordinates. In what follows, we will refer to it by the name of the "extended Lagrange equation."

## § 10.

## Applying the extended Lagrange equations to the Euler equations of a rigid body.

We would like to take the Euler equations of a rigid body as an example of the application of the generalized Lagrange equations.

If we express the kinetic energy of a rigid body that is fixed at a point in terms of the components $\pi, \chi, \rho$ of the angular velocities around three axes that are fixed in the body then, as is known, that will give:

$$
\begin{equation*}
L^{\prime}=\frac{1}{2} A \pi^{2}+\frac{1}{2} B \stackrel{*}{2}^{2}+\frac{1}{2} C{\stackrel{*}{\rho^{2}}}^{2} \tag{102}
\end{equation*}
$$

in which $A, B, C$ are the principal moments of inertia for the fixed point. The quantity of work $\delta^{\prime} A$ can then be expressed in terms of the infinitesimal rotation angles $\delta \pi, \delta \chi, \delta \rho$ as:

$$
\begin{equation*}
\delta^{\prime} A=\Pi \delta \pi+\mathrm{X} \delta \chi+\mathrm{P} \delta \rho \tag{103}
\end{equation*}
$$

in which $\Pi, \mathrm{X}, \mathrm{P}$ are the moments of the external forces about the principal axes of inertia.
In equation (71), we already gave the relations between the velocities $\stackrel{*}{\pi}, \stackrel{*}{\chi}, \stackrel{*}{\rho}$, and those of the Euler angles $\dot{\varphi}, \dot{\chi}, \dot{\vartheta}$ :

$$
\left\{\begin{array}{l}
\stackrel{*}{\pi}=\sin \psi \sin \vartheta \cdot \dot{\varphi}+\cos \psi \cdot \dot{\vartheta}+0 \cdot \dot{\psi}  \tag{104}\\
\stackrel{*}{\chi}=\cos \psi \sin \vartheta \cdot \dot{\varphi}-\sin \psi \cdot \dot{\vartheta}+0 \cdot \dot{\psi} \\
\stackrel{*}{\rho}=\cos \vartheta \cdot \dot{\varphi}+0 \quad \cdot \dot{\vartheta}+1 \cdot \dot{\psi}
\end{array}\right.
$$

With our previous notation, the coefficients $\alpha_{k r}$ will then have the values:

$$
\left\{\begin{array}{lll}
\alpha_{11}=\sin \psi \sin \vartheta, & \alpha_{21}=\cos \psi, & \alpha_{31}=0  \tag{105}\\
\alpha_{12}=\cos \psi \sin \vartheta, & \alpha_{22}=-\sin \psi, & \alpha_{32}=0 \\
\alpha_{13}= & \cos \vartheta, & \alpha_{23}=0,
\end{array} \alpha_{33}=1 .\right.
$$

If one solves equations (104) then one will get:

$$
\left\{\begin{array}{l}
\dot{\varphi}=\frac{\sin \psi}{\sin \vartheta} \cdot \pi^{*}+\frac{\cos \psi}{\sin \vartheta} \cdot \chi^{*}+0 \cdot \stackrel{*}{\rho}^{*}  \tag{106}\\
\dot{\vartheta}=\cos \psi \dot{\pi}-\sin \psi \cdot \dot{\chi}^{*}+0 \cdot{ }^{*} \rho \\
\dot{\psi}=-\frac{\cos \vartheta \sin \psi}{\sin \vartheta} \cdot \dot{\pi}^{\pi}-\frac{\cos \vartheta \cos \psi}{\sin \vartheta} \cdot \dot{\chi}^{*}+1 \cdot \stackrel{*}{\rho}^{*}
\end{array}\right.
$$

with the coefficients $\beta_{r \kappa}$ :

$$
\left\{\begin{array}{lll}
\beta_{11}=\frac{\sin \psi}{\sin \vartheta}, & \beta_{21}=\frac{\cos \psi}{\sin \vartheta}, & \beta_{31}=0  \tag{107}\\
\beta_{12}=\cos \psi, & \beta_{22}=-\sin \psi, & \beta_{32}=0 \\
\beta_{13}=-\frac{\cos \vartheta \sin \psi}{\sin \vartheta}, & \beta_{23}=-\frac{\cos \vartheta \cos \psi}{\sin \vartheta}, & \beta_{33}=1
\end{array}\right.
$$

The quantities $\alpha_{k r}$ obviously do not fulfill the integrability conditions (69), so $d \pi, d \chi, d \rho$ are the differentials of non-holonomic coordinates, i.e., in our previous notation, they would be denoted by $d \pi_{1}, d \pi_{2}, d \pi_{3}$.

If one would then like to employ the expression (102) for the kinetic energy, as well as (103) for the elementary work $\delta^{\prime} A$, in order to derive the equations of motion for the rigid body from Lagrange's then one must apply the extended Lagrange equations (101), not the simple ones (95).

The latter would imply the following:

$$
\left(\frac{\partial L^{\prime}}{\partial \pi}\right)=\left(\frac{\partial L^{\prime}}{\partial \chi}\right)=\left(\frac{\partial L^{\prime}}{\partial \rho}\right)=0
$$

since $L^{\prime}$ does not depend do not depend upon the coordinates at all, and furthermore:

$$
\frac{\partial L^{\prime}}{\partial \stackrel{*}{\pi}}=A \stackrel{*}{\pi}, \quad \frac{\partial L^{\prime}}{\partial \stackrel{*}{\chi}}=B \stackrel{*}{\chi}, \quad \frac{\partial L^{\prime}}{\partial \stackrel{*}{\rho}}=C \stackrel{*}{\rho},
$$

so one would arrive at the following equations:

$$
\left\{\begin{array}{l}
A \frac{d^{*}}{d t}=\Pi \\
B \frac{d^{*} \chi}{d t}=\mathrm{X}  \tag{108}\\
C \frac{d^{\rho}}{d t}=\mathrm{P}
\end{array}\right.
$$

while Euler's read:

$$
\begin{align*}
& A \frac{d \stackrel{*}{\pi}}{d t}+(C-B) \stackrel{*}{\chi} \stackrel{*}{\rho}=\Pi \\
& B \frac{d \stackrel{*}{\chi}}{d t}+(A-C) \stackrel{*}{\rho} \stackrel{*}{\pi}=\mathrm{X}  \tag{109}\\
& C \frac{d \stackrel{*}{\rho}}{d t}+(B-A) \stackrel{*}{\pi} \stackrel{*}{\chi}=\mathrm{P}
\end{align*}
$$

We must then employ only the extended equations, which will then add the term:

$$
\sum_{\rho=1}^{3} \sum_{\sigma=1}^{3} \gamma_{r \rho \sigma} \frac{\partial L^{\prime}}{\partial \pi_{\sigma}} \stackrel{*}{\sigma}_{\sigma}
$$

to the left-hand sides of equation (108).
One will next find the following values for the $\gamma_{r \rho \sigma}$ by somewhat-tedious, but elementary, calculations:
(109.a) $\left|\begin{array}{lll}\gamma_{111}=0 & \gamma_{211}=0 & \gamma_{311}=0 \\ \gamma_{121}=0 & \gamma_{221}=0 & \gamma_{321}=-1 \\ \gamma_{131}=0 & \gamma_{231}=1 & \gamma_{331}=0\end{array}\right|,\left|\begin{array}{lll}\gamma_{112}=0 & \gamma_{212}=0 & \gamma_{312}=1 \\ \gamma_{122}=0 & \gamma_{222}=0 & \gamma_{322}=0 \\ \gamma_{132}=-1 & \gamma_{232}=0 & \gamma_{332}=0\end{array}\right|,\left|\begin{array}{lll}\gamma_{113}=0 & \gamma_{213}=-1 & \gamma_{313}=0 \\ \gamma_{123}=1 & \gamma_{223}=0 & \gamma_{323}=0 \\ \gamma_{133}=0 & \gamma_{233}=0 & \gamma_{333}=0\end{array}\right|$.

One sees that all quantities $\gamma_{r \rho \sigma}$ with two or three equal indices are equal to zero, while all of the ones with three different indices will be equal to -1 or +1 . Let us now calculate the nineterm sums:

$$
\sum_{\rho=1}^{3} \sum_{\sigma=1}^{3} \gamma_{r \rho \sigma} \frac{\partial L^{\prime}}{\partial \pi_{\sigma}}{ }^{*}{ }_{\sigma}
$$

for each value of $r$. Here, that will simplify to two terms since from the matrices in (109.a), only two of any nine values of each $\gamma$ will be non-zero. For $r=1$, i.e., for $\stackrel{*}{\pi}_{1}=\stackrel{*}{\pi}$, that will give the following expression:

$$
\stackrel{*}{\pi_{3}} \frac{\partial L^{\prime}}{\partial \pi_{2}} \gamma_{132}+\stackrel{*}{\pi}_{2} \frac{\partial L^{\prime}}{\partial \pi_{3}} \gamma_{123}=\stackrel{*}{\pi}_{2} \frac{\partial L^{\prime}}{\partial \pi_{3}}-\stackrel{*}{\pi_{3}} \frac{\partial L^{\prime}}{\partial \pi_{2}}=\stackrel{*}{\chi} \frac{\partial L^{\prime}}{\partial \stackrel{*}{\rho}}-\stackrel{*}{\rho} \frac{\partial L^{\prime}}{\partial \stackrel{*}{\chi}}
$$

If one substitutes those values then one will find that the additional term is:

$$
(C-B) \stackrel{*}{\chi} \stackrel{*}{\rho},
$$

with which, the incorrect equations (108) will, in fact, go to the correct Euler equations (109). Q.E.D.

## § 11.

## The Hölder transformation ( ${ }^{1}$ ).

In order to prove the complete equivalence of the principle of stationary action with d'Alembert's, in § 8, we started from the expression on the left-hand side of d'Alembert's principle:

$$
\sum_{v=1}^{n}\left[\left(X_{v}-m_{v} \ddot{x}_{v}\right) \delta x_{v}+\left(Y_{v}-m_{v} \ddot{y}_{v}\right) \delta y_{v}+\left(Z_{v}-m_{v} \ddot{z}_{v}\right) \delta z_{v}\right]
$$

in which $\delta x_{v}, \delta y_{v}, \delta z_{v}$ are not initially virtual displacements, so the expression itself needs to be non-zero. The variations that enter into it would all be characterized by the symbol $\delta$, i.e., time would not be varied along with everything else.

We would now like to transform that expression, but apply the $\Delta$ process everywhere, instead of $\delta$.

That first requires a prefatory remark. Whereas we have proved that $\frac{\delta d x}{d t}=\frac{d \delta x}{d t}$, the commutability of $\delta$ and $d / d t$, which is based upon the fact that time is not varied, will no longer be true for $\Delta$ and $d / d t$, which we already pointed out at that time. We will then be dealing with the relationship between $\Delta\left(\frac{d x}{d t}\right)$ and $\frac{d(\Delta x)}{d t}$. In order to answer that question, it is appropriate to introduce a new independent variable $\Theta$, which we will make depend upon both $x$ and $t$; it shall also not be varied under the $\Delta$ process. We shall then reduce the presently-complicated case to the previously-simpler one by that parametric representation of $x$ and $t$. We have to set: $\frac{d x}{d t}=$ $\frac{x^{\prime}}{t^{\prime}}$, when we denote the derivatives with respect to the parameter $\Theta$ by a prime $\left(\frac{d x}{d \Theta}=x^{\prime}\right.$, $\left.\frac{d t}{d \Theta}=t^{\prime}\right)$.

We find, in succession, that:

$$
\Delta\left(\frac{x^{\prime}}{t^{\prime}}\right)=\frac{t^{\prime} \Delta x^{\prime}-x^{\prime} \Delta t^{\prime}}{t^{\prime 2}}=\frac{\Delta x^{\prime}}{t^{\prime}}-\frac{x^{\prime}}{t^{\prime}} \cdot \frac{\Delta t^{\prime}}{t^{\prime}}
$$

or
${ }^{(1)}$ O. Hölder, "Über die Prinzipien von Hamilton und Maupertuis," Nachr. kgl. Ges. Wiss. Göttingen, math.phys. Klasse (1896), pp. 122, et seq.

$$
\Delta\left(\frac{x^{\prime}}{t^{\prime}}\right)=\frac{\Delta\left(\frac{d x}{d \Theta}\right)}{\left(\frac{d t}{d \Theta}\right)}-\frac{\left(\frac{d x}{d \Theta}\right)}{\left(\frac{d t}{d \Theta}\right)} \cdot \frac{\Delta\left(\frac{d t}{d \Theta}\right)}{\left(\frac{d t}{d \Theta}\right)}
$$

or, since $\Delta\left(\frac{d x}{d \Theta}\right)=\frac{d}{d \Theta} \Delta x$, etc., since $\Theta$ is not varied in that, one will have:

$$
\Delta\left(\frac{d x}{d t}\right)=\Delta\left(\frac{x^{\prime}}{t^{\prime}}\right)=\frac{\frac{d \Delta x}{d \Theta}}{\frac{d t}{d \Theta}}-\left(\frac{d x}{d t}\right) \cdot \frac{\frac{d \Delta t}{d \Theta}}{\frac{d t}{d \Theta}}=\frac{d \Delta x}{d t}-\dot{x} \frac{d \Delta t}{d t}
$$

We ultimately find that:

$$
\begin{equation*}
\Delta\left(\frac{d x}{d t}\right)=\frac{d}{d t}(\Delta x)-\dot{x} \frac{d \Delta t}{d t} . \tag{110}
\end{equation*}
$$

Moreover, one can also derive that equation from (67) directly.
With that preliminary remark, we go on to derive the transformation of the d'Alembert expression. Let the kinetic energy of a system of $n$ mass-points be:

$$
L=\sum_{v=1}^{n} \frac{1}{2} m_{v}\left(\dot{x}_{v}^{2}+\dot{y}_{v}^{2}+\dot{z}_{v}^{2}\right) .
$$

We then form $\Delta L$ :

$$
\begin{equation*}
\Delta L=\Delta \sum_{v=1}^{n} \frac{1}{2} m_{v}\left(\dot{x}_{v}^{2}+\dot{y}_{v}^{2}+\dot{z}_{v}^{2}\right)=\sum_{v=1}^{n} m_{v}\left(\dot{x}_{v} \Delta \dot{x}_{v}+\dot{y}_{v} \Delta \dot{y}_{v}+\dot{z}_{v} \Delta \dot{z}_{v}\right), \tag{111}
\end{equation*}
$$

or from (110):

$$
\begin{aligned}
\Delta L & =\sum_{v=1}^{n} m_{v}\left[\dot{x}_{v}\left(\frac{d \Delta x_{v}}{d t}-\dot{x}_{v} \frac{d \Delta t}{d t}\right)+\dot{y}_{v}\left(\frac{d \Delta y_{v}}{d t}-\dot{y}_{v} \frac{d \Delta t}{d t}\right)+\dot{z}_{v}\left(\frac{d \Delta z_{v}}{d t}-\dot{z}_{v} \frac{d \Delta t}{d t}\right)\right] \\
& =\sum_{v=1}^{n} m_{v}\left(\dot{x}_{v} \frac{d \Delta x_{v}}{d t}+\dot{y}_{v} \frac{d \Delta y_{v}}{d t}+\dot{z}_{v} \frac{d \Delta z_{v}}{d t}\right)-\sum_{v=1}^{n} \frac{1}{2} m_{v}\left(\dot{x}_{v}^{2}+\dot{y}_{v}^{2}+\dot{z}_{v}^{2}\right) \frac{d \Delta t}{d t}
\end{aligned}
$$

or if one observes the value of $L$ :

$$
\begin{equation*}
\Delta L+2 L \frac{d \Delta t}{d t}=\sum_{v=1}^{n} m_{v}\left(\dot{x}_{v} \frac{d \Delta x_{v}}{d t}+\dot{y}_{v} \frac{d \Delta y_{v}}{d t}+\dot{z}_{v} \frac{d \Delta z_{v}}{d t}\right) . \tag{112}
\end{equation*}
$$

The right-hand side of that can be further treated - term-by-term - according to the following rule:

$$
\dot{x}_{v} \frac{d \Delta x_{v}}{d t}=\frac{d}{d t}\left(\dot{x}_{v} \Delta x_{v}\right)-\ddot{x}_{v} \Delta x_{v}, \quad \text { etc. }
$$

It will then follow that:

$$
\Delta L+2 L \frac{d \Delta t}{d t}=\frac{d}{d t}\left[\sum_{v=1}^{n} m_{v}\left(\dot{x}_{v} \Delta x_{v}+\dot{y}_{v} \Delta y_{v}+\dot{z}_{v} \Delta z_{v}\right)\right]-\sum_{v=1}^{n} m_{v}\left(\ddot{x}_{v} \Delta x_{v}+\ddot{y}_{v} \Delta y_{v}+\ddot{z}_{v} \Delta z_{v}\right),
$$

or

$$
\begin{equation*}
-\sum_{v=1}^{n} m_{v}\left(\ddot{x}_{v} \Delta x_{v}+\ddot{y}_{v} \Delta y_{v}+\ddot{z}_{v} \Delta z_{v}\right)=\Delta L+2 L \frac{d \Delta t}{d t}-\frac{d}{d t}\left[\sum_{v=1}^{n} m_{v}\left(\dot{x}_{v} \Delta x_{v}+\dot{y}_{v} \Delta y_{v}+\dot{z}_{v} \Delta z_{v}\right)\right] . \tag{113}
\end{equation*}
$$

That suggests an integration over time. If we simultaneously determine the $\Delta x_{v}, \Delta y_{v}, \Delta z_{v}$ (as in Hamilton's principle) such that they will vanish at the limits of the integral at the times $t_{0}$ and $t_{1}$ then the term that emerges from the integral will drop out, and what will remain:

$$
\begin{equation*}
-\int_{t_{0}}^{t_{1}} d t \sum_{v=1}^{n} m_{v}\left(\ddot{x}_{v} \Delta x_{v}+\ddot{y}_{v} \Delta y_{v}+\ddot{z}_{v} \Delta z_{v}\right)=\int_{t_{0}}^{t_{1}}\left(\Delta L+2 L \frac{d \Delta t}{d t}\right) d t . \tag{114}
\end{equation*}
$$

If we further denote the expression $\sum_{v=1}^{n}\left(X_{v} \Delta x_{v}+Y_{v} \Delta y_{v}+Z_{v} \Delta z_{v}\right)$ by $\Delta^{\prime} A$ and integrate over time from $t_{0}$ to $t_{1}$ then we will have:

$$
\begin{equation*}
\int_{t_{0}}^{t_{1}} d t \sum_{v=1}^{n}\left(X_{v} \Delta x_{v}+Y_{v} \Delta y_{v}+Z_{v} \Delta z_{v}\right)=\int_{t_{0}}^{t_{1}} \Delta^{\prime} A d t \tag{115}
\end{equation*}
$$

Adding (114) and (115) to the left-hand side of the d'Alembert expression, integrating over $t$, will then yield the desired transformation of the right-hand side:

$$
\begin{equation*}
\int_{t_{0}}^{t_{1}} d t \sum_{v=1}^{n}\left[\left(X_{v}-m_{v} \ddot{x}_{v}\right) \Delta x_{v}+\left(Y_{v}-m_{v} \ddot{y}_{v}\right) \Delta y_{v}+\left(Z_{v}-m_{v} \ddot{z}_{v}\right) \Delta z_{v}\right]=\int_{t_{0}}^{t_{1}} d t\left(\Delta L+\Delta^{\prime} A+2 L \frac{d \Delta t}{d t}\right) \tag{116}
\end{equation*}
$$

We would like to call this equation the Hölder transformation. It represents a generalization of equation (88) and differs from it only by the fact that the more general $\Delta$ process was applied in it. If we then set $\Delta t=0$ in equation (116), the $\Delta x_{v}, \ldots$ will become identical to the $\delta x_{v}, \ldots$, and (116) will go to (88), from which the principle of stationary action will then follow.

The meaning of the Hölder transformation lies in the following fact:
Naturally, the variation that appears in (116) is much more general than the previous one, due to the appearance of $\Delta t$. We can then prescribe any sort of relation between the quantities $\Delta x_{v}$, $\Delta y_{v}, \Delta z_{v}$, and $\Delta t$, i.e., suitably restrict the most-general variations that enter into (116). For every
chosen restriction, we will get a new dynamical principle from (116). The most radical restriction would be $\Delta t=0$, and that would then imply just Hamilton's principle of stationary action.

In the next section, we shall discuss those constraints that will make (116) lead to the socalled principle of least action that was first presented by Euler and Maupertuis but formulated precisely by Lagrange.

## § 12.

## The various forms of the principle of least action.

We would next like to treat equation (116) further without introducing any specialization into it.

We first set:

$$
\Delta x_{v}=\delta x_{v}+\dot{x}_{v} \Delta t, \quad \text { etc. }
$$

in the left-hand side of it and obtain:

$$
\begin{align*}
& \int_{t_{0}}^{t_{1}} d t \sum_{v=1}^{n}\left[\left(X_{v}-m_{v} \ddot{x}_{v}\right) \delta x_{v}+\left(Y_{v}-m_{v} \ddot{y}_{v}\right) \delta y_{v}+\left(Z_{v}-m_{v} \ddot{z}_{v}\right) \delta z_{v}\right] \\
& +\int_{t_{0}}^{t_{1}} d t \sum_{v=1}^{n}\left[X_{v} \delta x_{v}+Y_{v} \delta y_{v}+Z_{v} \delta z_{v}\right]  \tag{117}\\
& -\int_{t_{0}}^{t_{1}} d t \sum_{v=1}^{n} m_{v}\left[\ddot{x}_{v} \dot{x}_{v}+\ddot{y}_{v} \dot{y}_{v}+\ddot{z}_{v} \dot{z}_{v}\right] \Delta t=\int_{t_{0}}^{t_{1}} d t\left[\Delta L+\Delta^{\prime} A+2 L \frac{d \Delta t^{\prime}}{d t}\right] .
\end{align*}
$$

Now, any sort of constraints - say, $m$ of them - might be prescribed, perhaps non-holonomic, rheonomic ones, in order to remain as general as possible:

$$
\begin{equation*}
\sum_{v=1}^{n}\left(a_{k v}^{\prime} d x_{v}+b_{k v}^{\prime} d y_{v}+c_{k v}^{\prime} d z_{v}+a_{v}^{\prime} d t\right)=0 \quad(k=1,2, \ldots, m) \tag{118}
\end{equation*}
$$

If we subject the $\delta x_{v}, \delta y_{v}, \delta z_{v}$, which were completely-free up to now, to the constraints that are valid for virtual displacements:

$$
\sum_{v=1}^{n}\left(a_{k v}^{\prime} \delta x_{v}+b_{k v}^{\prime} \delta y_{v}+c_{k v}^{\prime} \delta z_{v}\right)=0 \quad(k=1,2, \ldots, m)
$$

then the first term on the right-hand side of (117) will vanish, according to d'Alembert's principle. Moreover, the square bracket in the second term obviously represents the work done per unit time $d^{\prime} A / d t$, and the one in the third term represents the change in kinetic energy per unit time $d L / d t$. If we introduce that into (117) then we will find the following expression for the virtual displacements, which is completely equivalent to d'Alembert's principle, so to the equations of motion:

$$
\begin{equation*}
\int_{t_{0}}^{t_{1}} d t\left[\Delta L+\Delta^{\prime} A+2 L \frac{d \Delta t}{d t}-\frac{d^{\prime} A}{d t} \Delta t+\frac{d L}{d t} \Delta t\right]=0 . \tag{119}
\end{equation*}
$$

That equation, in turn, represents a very general principle of dynamics from which the equations of motion can be obtained by purely-formal process.

From what we said before, we can now restrict the variations that enter into (119) in a more suitable way. Every such auxiliary condition will then produce a special dynamical principle.

Now, Lagrange ( ${ }^{1}$ ) has shown that one will arrive at the principle of least action that Euler first expressed in some special cases when one demands the validity of the energy principle for the transitional path that is composed of virtual displacements, i.e., when one demands that one should have:

$$
\begin{equation*}
\delta L=\delta^{\prime} A \tag{120}
\end{equation*}
$$

That does not say, by any means, that the energy principle should be valid for the true or varied path, which is indeed not the case for rheonomic constraints.

If we now assume, for the sake of simplicity, that the constraints are scleronomic (whether holonomic or non-holonomic) then it will follow that the energy equation $\frac{d L}{d t}=\frac{d^{\prime} A}{d t}$ is, in fact, true for the true motion, and (119) will simplify to:

$$
\begin{equation*}
\int_{t_{0}}^{t_{1}} d t\left[\Delta L+\Delta^{\prime} A+2 L \frac{d \Delta t}{d t}\right]=0 . \tag{121}
\end{equation*}
$$

If we introduce the Lagrange "transition condition" (120) into that, which we can also write in the following form, in which $\delta$ is expressed as $\Delta$ [according to equation (67)]:

$$
\begin{equation*}
\Delta L-\frac{d L}{d t} \Delta t=\Delta^{\prime} A-\frac{d A}{d t} \Delta t \tag{120.a}
\end{equation*}
$$

then (121) will imply that:

$$
\int_{t_{0}}^{t_{1}} d t\left[2 \Delta L+2 L \frac{d \Delta t}{d t}\right]=2 \int[\Delta L d t+L d \Delta t]=0 .
$$

However, $\Delta L d t+L d \Delta t$ is equal to $\Delta(L d t)$, so one can drop the factor of 2 and write:

$$
\int_{t_{0}}^{t_{1}} \Delta(L d t)=0
$$

or after switching the integral sign with $\Delta$ :

[^5]\[

$$
\begin{equation*}
\Delta \int_{t_{0}}^{t_{1}} L d t=0 \tag{122}
\end{equation*}
$$

\]

That equation is called the principle of least action, and we shall return to its formulation in more detailed words.

Up to now, it was derived under the assumption that the constraints were scleronomic. If we now drop that restricting assumption then the law of energy $\frac{d L}{d t}=\frac{d^{\prime} A}{d t}$ will no longer be fulfilled for the true motion, and we must once more start from the general equation (119):

$$
\begin{equation*}
\int_{t_{0}}^{t_{1}} d t\left[\Delta L+\Delta^{\prime} A+2 L \frac{d \Delta t}{d t}-\frac{d^{\prime} A}{d t} \Delta t+\frac{d L}{d t} \Delta t\right]=0 \tag{119}
\end{equation*}
$$

into which we have again introduced the Lagrange transition condition in the form (120.a). It will then follow that:

$$
\int_{t_{0}}^{t_{1}} d t\left[2 \Delta L-\frac{d L}{d t} \Delta t+\frac{d^{\prime} A}{d t} \Delta t+2 L \frac{d \Delta t}{d t}-\frac{d^{\prime} A}{d t} \Delta t+\frac{d L}{d t} \Delta t\right]=0
$$

or

$$
\int[\Delta L d t+L d \Delta t]=0
$$

or:

$$
\begin{equation*}
\Delta \int_{t_{0}}^{t_{1}} L d t=0 \tag{122}
\end{equation*}
$$

as above, i.e., the principle of least action, which is now proved under the most general assumptions.

Except for an irrelevant constant factor, $\int_{t_{0}}^{t_{0}} L d t$ is the temporal mean value of the kinetic energy. We can then say:

For the true motion, the temporal mean value of the kinetic energy is stationary, i.e., a maximum, minimum, or saddle-point value.

The true path is then compared to those varied paths that have the same starting point and end point as the former has at time $t_{0}$ and $t_{1}$, and in addition satisfy the condition that the energy principle should be true under the transition from the true to the varied motion.

The first statement would be meaningless without that additional condition.
More precisely, we would like to refer to (122) as the broader form of the principle of least action. We once more stress that it is not assumed in it that the energy principle is valid for the true (and varied) path, but its validity is required of only the transitional motion.

We will get a narrower form for the aforementioned principle when we also assume the energy principle for the true motion, and indeed in the form (as an energy integral):

$$
\begin{equation*}
L+\Phi=h . \tag{123}
\end{equation*}
$$

According to Jacobi $\left({ }^{1}\right)$, one can proceed as follows: Time $t$ and the time differential $d t$ are eliminated from equation (122). For the sake of generality, we then set $L$ equal to:

$$
\begin{equation*}
2 L=\sum_{i=1}^{N} \sum_{k=1}^{N} a_{i k} \frac{d q_{i}}{d t} \frac{d q_{k}}{d t} \tag{124}
\end{equation*}
$$

in terms of generalized coordinates $q_{k}$ and velocities $\dot{q}_{k}$ (namely, as a homogeneous quadratic form in the latter) then that will imply $d t^{2}$ in the form:

$$
d t^{2}=\frac{\sum_{i, k} a_{i k} d q_{i} d q_{k}}{2 L}
$$

or from (123):

$$
\begin{equation*}
d t=\sqrt{\frac{\sum a_{i k} d q_{i} d q_{k}}{2(h-\Phi)}} \tag{125}
\end{equation*}
$$

Since $\Phi$ depends upon only $q_{k}, d t$ is then expressed in terms of purely geometrical quantities.
We likewise replace $L$ with $(h-\Phi)$ in (122) and get:

$$
\begin{equation*}
\Delta \int \sqrt{\frac{h-\Phi}{2}} \cdot \sqrt{\sum a_{i k} d q_{i} d q_{k}}=0 \tag{126}
\end{equation*}
$$

instead of (122). Naturally, geometric data is substituted at the limits of the integral at times $t_{0}$ and $t_{1}$, namely, one gives the configuration of the system at those times.

Equation (126) is the Jacobi form of the principle of least action, and from the above, it is valid only when the energy integral is valid for the true path, i.e., only for the so-called complete systems. However, it is precisely in mechanics that one must chiefly deal with incomplete systems, and the applicability of (126) will then be restricted. Jacobi's mistake, if one might call it that, consisted of the fact that he believed that this was the only form that was valid, while in reality, (122) is valid much more generally.

Now, the fact that the Jacobi principle assumes the energy principle for the true path demands that the energy principle must also be fulfilled for the Lagrange transition condition (120). That is because energy has the value $L+\Phi=h$ for each point $P$ of the true path, i.e., it is equal to a constant. If I go to the associated point $P^{\prime}$ of the varied path then $\delta h=0$, according to

[^6]Lagrange, i.e., the point $P^{\prime}$ will correspond to the same value of energy as $P$. One can then arrange for that to be true for every point of the true and varied path, and since all points of the true path have the same value of $h$, that will also be true for the varied path. That is the same line of reasoning by which we previously proved that when a holonomic equation of constraint is fulfilled by the true path, that will also be the case for the varied path since that is the case for the transitional path.

The equations of motion are implied by the broader form (122) when one goes through the line of reasoning that was used to derive it in reverse order. We shall not go into that. Rather, we would like to show that the Jacobi form (126) does, in fact, lead to the equations of motion for a complete system. Indeed, we would like to implement that for Cartesian coordinates, which we would like to denote here, not by $x_{1}, y_{1}, z_{1}, \ldots, x_{n}, y_{n}, z_{n}$, but in general by $x_{i}$, which runs through $x_{1}, x_{2}, x_{3}, x_{4}, x_{5}, \ldots$, for the sake of convenience. $L$ will then have the following form:

$$
L=\frac{1}{2} \sum m_{i}\left(\frac{d x_{i}}{d t}\right)^{2},
$$

and it will follow from (126) that:

$$
\begin{equation*}
\Delta \int \sqrt{\frac{h-\Phi}{2}} \sqrt{\sum m_{i} d x_{i}^{2}}=0 . \tag{127}
\end{equation*}
$$

Here, it would be convenient to regard the $x_{i}$ as functions of a parameter $\Theta$ that is not affected by the variation. We therefore write (127) as:

$$
\begin{equation*}
\Delta \int_{t_{0}}^{t_{1}} \sqrt{\frac{h-\Phi}{2}} \sqrt{\sum m_{i}\left(\frac{d x_{i}}{d \Theta}\right)^{2}} d \Theta=0 \tag{127.a}
\end{equation*}
$$

If we denote the derivatives by a prime, as before, then that will give the general form of the variational principle that is included in (127.a)

$$
\begin{equation*}
\Delta \int_{\Theta_{0}}^{\Theta_{1}} F\left(x_{i}, x_{i}^{\prime}\right) d \Theta \tag{128}
\end{equation*}
$$

in which we have set:

$$
\begin{equation*}
F\left(x_{i}, x_{i}^{\prime}\right)=\sqrt{\frac{h-\Phi}{2}} \sqrt{\sum_{i} m_{i} x_{i}^{\prime 2}} . \tag{129}
\end{equation*}
$$

Upon developing (128), it will follow that:

$$
\begin{equation*}
\int_{t_{0}}^{t_{1}} d \Theta \Delta F=\int_{t_{0}}^{t_{1}} d \Theta\left[\sum_{i} \frac{\partial F}{\partial x_{i}} \Delta x_{i}+\sum_{i} \frac{\partial F}{\partial x_{i}^{\prime}} \Delta x_{i}^{\prime}\right]=0 . \tag{130}
\end{equation*}
$$

The second term with $\Delta x_{i}^{\prime}$ can be converted by partial integration:

$$
\int_{t_{0}}^{t_{1}} \frac{\partial F}{\partial x_{i}^{\prime}} \Delta x_{i}^{\prime} d \Theta=\int_{t_{0}}^{t_{1}} \frac{\partial F}{\partial x_{i}^{\prime}} \frac{d \Delta x_{i}^{\prime}}{d \Theta} d \Theta=\left[\frac{\partial F}{\partial x_{i}^{\prime}} \Delta x_{i}\right]_{\Theta_{0}}^{\Theta_{1}}-\int_{\Theta_{0}}^{\Theta_{1}} \frac{d}{d \Theta}\left(\frac{\partial F}{\partial x_{i}^{\prime}}\right) \Delta x_{i} d \Theta .
$$

The term outside the integral drops out since the $\Delta x_{i}$ should vanish at the limits, as before. Therefore, (130) will become:

$$
\begin{equation*}
\int_{\Theta_{0}}^{\Theta_{1}} d \Theta \sum_{i}\left[\frac{\partial F}{\partial x_{i}}-\frac{d}{d \Theta}\left(\frac{\partial F}{\partial x_{i}^{\prime}}\right)\right] \Delta x_{i}=0 . \tag{131}
\end{equation*}
$$

If we take the simplest case in which no constraints exist, then it will follow by the argument that has already been applied frequently that:

$$
\begin{equation*}
\frac{\partial F}{\partial x_{i}}-\frac{d}{d \Theta}\left(\frac{\partial F}{\partial x_{i}^{\prime}}\right)=0 \tag{132}
\end{equation*}
$$

for all values of $i$.
One will see that those are actually the equations of motion for a complete system when one performs the differentiations on the $F$ in (129). One will initially get:

$$
\begin{equation*}
\frac{\partial F}{\partial x_{i}}=\frac{-\frac{1}{2} \frac{\partial \Phi}{\partial x_{i}} \sum m_{i} x_{i}^{\prime 2}}{2 \sqrt{\frac{h-\Phi}{2}} \cdot \sum m_{i} x_{i}^{\prime 2}} \tag{133}
\end{equation*}
$$

However, one had:

$$
2 L=\sum m_{i}\left(\frac{d x_{i}}{d \Theta}\right)^{2}\left(\frac{d \Theta}{d t}\right)^{2}=\sum m_{i} x_{i}^{\prime 2}\left(\frac{d \Theta}{d t}\right)^{2}
$$

hence:

$$
\sum m_{i} x_{i}^{\prime 2}=2 L\left(\frac{d t}{d \Theta}\right)^{2}=2(h-\Phi)\left(\frac{d t}{d \Theta}\right)^{2}
$$

When that is substituted in (133), that will give:

$$
\begin{equation*}
\frac{\partial F}{\partial x_{i}}=-\frac{1}{2} \frac{\partial \Phi}{\partial x_{i}}\left(\frac{d t}{d \Theta}\right) \tag{134}
\end{equation*}
$$

Moreover, from (129):

$$
\frac{\partial F}{\partial x_{i}^{\prime}}=\frac{\sqrt{\frac{h-\Phi}{2}} m_{i} x_{i}^{\prime}}{\sqrt{\sum_{i} m_{i} x_{i}^{\prime 2}}}=\sqrt{\frac{\frac{1}{2}(h-\Phi)}{2(h-\Phi)\left(\frac{d t}{d \Theta}\right)^{2}}} m_{i} x_{i}^{\prime}=\frac{1}{2} m_{i} \frac{d x_{i}}{d \Theta} \frac{d \Theta}{d t}=\frac{1}{2} m_{i} \frac{d x_{i}}{d t}
$$

so one will finally have:

$$
\frac{d}{d \Theta}\left(\frac{\partial F}{\partial x_{i}^{\prime}}\right)=\frac{1}{2} m_{i} \frac{d}{d \Theta}\left(\frac{d x_{i}}{d t}\right)
$$

When (134) and (135) are substituted in (132), that will yield:

$$
\frac{\partial \Phi}{\partial x_{i}} \frac{d t}{d \Theta}+m_{i} \frac{d}{d \Theta}\left(\frac{d x_{i}}{d t}\right)=0
$$

or

$$
\frac{\partial \Phi}{\partial x_{i}}+m_{i} \frac{d}{d \Theta}\left(\frac{d x_{i}}{d t}\right) \frac{d \Theta}{d t}=0
$$

or finally:

$$
m_{i} \frac{d^{2} x_{i}}{d t^{2}}=-\frac{\partial \Phi}{\partial x_{i}}
$$

which was to be proved.
One can put the principle of least action into yet another form. If one takes the velocity of the $i^{\text {th }}$ mass-point to be $v_{i}=d s_{i} / d t$, in which $d s_{i}$ means the path element of that mass-point, then equation (122) can be written:

$$
\Delta \int L d t=\frac{1}{2} \Delta \int \sum m_{i} v_{i} \frac{d s_{i}}{d t} d t=\frac{1}{2} \Delta \int \sum m_{i} v_{i} d s_{i}=0
$$

or after cancelling the constant factor:

$$
\begin{equation*}
\Delta \int \sum m_{i} v_{i} d s_{i}=0 \tag{136}
\end{equation*}
$$

That equation gave the principle its name, since one cares to refer to the expression $m_{i} v_{i} d s_{i}$ as the "action." Of course, it is by no means the case that the total action $\int \sum m_{i} v_{i} d s_{i}$ must be a minimum, as the epithet "least" action might suggest. However, such questions, which are interesting in the calculus of variations, play no role in dynamics.

In the special case of one mass-point, it will then follow that:

$$
\begin{equation*}
\Delta \int_{s_{0}}^{s_{1}} m v d s=0 \tag{137}
\end{equation*}
$$

and if it is also force-free then since its velocity $v$ would be constant then, it will follow that:

$$
\begin{equation*}
\Delta \int_{s_{0}}^{s_{1}} d s=0 \tag{138}
\end{equation*}
$$

That equation admits a simple physical interpretation:
$\int_{s_{0}}^{s_{1}} d s$ is the length of the path that is laid through the force-free particles between the points $s_{0}$ and $s_{1}$. From (138), the length shall always have an extremal value, and in the simplest case (which will always occur when the points $s_{0}$ and $s_{1}$ lie close enough to each other), it will be a minimum. However, the shortest path between two points is a straight line. Thus, (138) immediately implies the statement that a force-free mass-point moves rectilinearly.

In the foregoing, it was tacitly assumed that the force-free mass-point is subject to no kinematical constraints. However, if such a thing is prescribed then, in general, the mass-point will naturally no longer describe a straight line on the surface that it prescribes. However, it is natural that the curve on the surface that is described between two sufficiently-close points $s_{0}$ and $s_{1}$ will also have the "minimal property" that is expressed by equation (138). Such curves are called geodetic or, according to Hertz, straightest. It is easy to obtain their equation. We regard the $x_{i}$ in (138) as functions of a parameter $\Theta$ in terms of which that equation can be written, and since $d s^{2}=d x_{1}^{2}+d x_{2}^{2}+d x_{3}^{2}$, we will have:

$$
\begin{equation*}
\Delta \int_{\Theta_{0}}^{\Theta_{1}} \sqrt{x_{1}^{\prime 2}+x_{2}^{\prime 2}+x_{3}^{\prime 2}} d \Theta=0 \tag{138.a}
\end{equation*}
$$

If we further multiply the prescribed constraint equation $\varphi=0$ by $d \Theta$ and integrate between two limits, as in (138.a), then we will get:

$$
\begin{equation*}
\int_{\Theta_{0}}^{\Theta_{1}} \varphi d \Theta=0 \tag{138.b}
\end{equation*}
$$

That equation is fulfilled at the same time as (138.a). We multiply that equation by an unknown factor $\lambda$ and add it to the integral in (138.a). The variation of the new integral that thus arises must vanish then. That will then give:

$$
\Delta \int_{\Theta_{0}}^{\Theta_{1}}\left\{\sqrt{x_{1}^{\prime 2}+x_{2}^{\prime 2}+x_{3}^{\prime 2}}+\lambda \varphi\right\} d \Theta=\Delta \int_{\Theta_{0}}^{\Theta_{1}} F\left(x_{i}, x_{i}^{\prime}\right) d \Theta=0
$$

We will then have an integral of the form (128), and performing the variation will produce the equation:

$$
\frac{\partial F}{\partial x_{i}}=\frac{d}{d \Theta}\left(\frac{\partial F}{\partial x_{i}^{\prime}}\right)
$$

Evaluating that for, e.g., $x_{1}$ will then yield:

$$
\lambda \frac{\partial \varphi}{\partial x_{1}}=\frac{d}{d \Theta}\left(\frac{x_{1}^{\prime}}{\sqrt{x_{1}^{\prime 2}+x_{2}^{\prime 2}+x_{3}^{\prime 2}}}\right)
$$

If we now choose the parameter $\Theta$ to be the arc-length $s$ then we will obviously have $\sqrt{x_{1}^{\prime 2}+x_{2}^{\prime 2}+x_{3}^{\prime 2}}=1$, and the equation above will go to:

$$
\frac{d^{2} x}{d s^{2}}=\lambda \frac{\partial \varphi}{\partial x_{1}}
$$

to which two corresponding ones must be added for $x_{2}$ and $x_{3}$. They will be the equations for the geodetic line on the surface $\varphi=0$ that the mass-point describes.

The form (138) can also be obtained for a force-free system. To that end, we start from the Jacobi equation (127), in which the factor $\sqrt{(h-\Phi) / 2}$ will be constant, due to the assumed absence of forces, so it can be dropped. What will remain is:

$$
\Delta \int \sqrt{\sum m_{i} d x_{i}^{2}}=0
$$

Now, Heinrich Hertz ${ }^{1}$ ) defined:

$$
\begin{equation*}
d s^{2} \sum m_{i}=\sum m_{i} d x_{i}^{2} \tag{139}
\end{equation*}
$$

in his theory of mechanics.
$d s^{2}$ is then a quadratic mean value that Hertz referred to as the square of the path-element of the system. With that, the Jacobi form will again become:

$$
\begin{equation*}
\Delta \int d s=0 \tag{140}
\end{equation*}
$$

Hertz called that formula (140), in conjunction with his definition of $d s$ that is included in (139), the principle of the straightest path.

Now, it is interesting that one can also impose that same form when conservative forces are at work. To that end, one must extend Hertz's definition (139) to:

[^7]\[

$$
\begin{equation*}
d \sigma^{2} \sum m_{i}=\left(\frac{h-\Phi}{2}\right) \sum m_{i} d x_{i}^{2} . \tag{141}
\end{equation*}
$$

\]

Since one interprets $d \sigma$ as the path-element of the system (of course, the manner by which it is composed from the $d x_{i}$ will no longer be the Euclidian form that comes from the Pythagorean theorem), that will obviously give, in turn:

$$
\begin{equation*}
\Delta \int d \sigma=0 \tag{142}
\end{equation*}
$$

One can then reduce the motion of a system in Euclidian space under the influence of conservative forces to the motion of a force-free system in a general Euclidian manifold in a purely-formal way.

That viewpoint has a strong similarity to the appearance of non-Euclidian geometry in Einstein's general theory of relativity.

## § 13.

## Hamilton's canonical equations.

Hamilton transformed the Lagrange equations in a convenient way and thus obtained a particularly-symmetric system that has been confirmed in all more-detailed investigations in dynamics, electrodynamics, and statistical mechanics, and in particular, at the hands of Maxwell, Boltzmann, and Gibbs. We shall now move on to the derivation of those equations.

The Lagrange equations read:

$$
\begin{equation*}
\frac{d}{d t}\left(\frac{\partial L}{\partial \dot{q}_{k}}\right)-\frac{\partial L}{\partial q_{k}}=Q_{k} \tag{95}
\end{equation*}
$$

One now assumes the existence of a potential $\Phi$, which will make $Q_{k}$ go to $-\partial \Phi / \partial q_{k}$. One can then write (95) as:

$$
\frac{d}{d t}\left(\frac{\partial L}{\partial \dot{q}_{k}}\right)-\frac{\partial}{\partial q_{k}}(L-\Phi)=0,
$$

or also, since $\Phi$ depends upon only the $q_{k}$, but not upon the $\dot{q}_{k}$, when one adds the vanishing term $-\partial \Phi / \partial \dot{q}_{k}$ :

$$
\frac{d}{d t}\left(\frac{\partial(L-\Phi)}{\partial \dot{q}_{k}}\right)-\frac{\partial}{\partial q_{k}}(L-\Phi)=0
$$

or when one introduces the notation $H$ for the difference $L-\Phi$, which was previously called the kinetic potential:

$$
\begin{equation*}
\frac{d}{d t}\left(\frac{\partial H}{\partial \dot{q}_{k}}\right)-\frac{\partial H}{\partial q_{k}}=0 \quad(k=1,2, \ldots, N) . \tag{143}
\end{equation*}
$$

The expressions $\partial H / \partial \dot{q}_{k}$ have a simple physical meaning that we can clarify by a dimensional consideration. $H$ has the dimensions of an energy $\left[M L^{2} T^{-2}\right]$. If we assume, for the moment, that $q_{k}$ has the dimension of length, which is indeed the case for, e.g., Cartesian coordinates, then $\dot{q}_{k}$ will have the dimension of a velocity $\left[L T^{-1}\right], \partial H / \partial \dot{q}_{k}$ will have the dimension $\left[M L T^{-1}\right.$ ], i.e., the dimension of an impulse. In fact, for Cartesian coordinates, $\partial H / \partial \dot{q}_{k}$ will coincide with the impulse components, which one can confirm by calculation. Now, in just the same way that we referred to $q_{k}$ and $\dot{q}_{k}$ as generalized coordinates and generalized velocities, even when they did not have the dimensions of length and velocity, and just as we further referred to the quantities $Q_{k}$ as the generalized force components, we will call the quantities $\partial H / \partial \dot{q}_{k}$ the generalized impulse components and denote them by $p_{k}$ :

$$
\begin{equation*}
\frac{\partial H}{\partial \dot{q}_{k}}=p_{k} \quad(k=1,2, \ldots, N) . \tag{144}
\end{equation*}
$$

The Lagrange equations will then read:

$$
\begin{equation*}
\frac{\partial H}{\partial q_{k}}=\frac{d p_{k}}{d t} . \tag{145}
\end{equation*}
$$

We now consider a variation $\delta$ of $H\left(q_{k}, \dot{q}_{k}\right)$, which will read:

$$
\delta H=\sum_{k} \frac{\partial H}{\partial q_{k}} \delta q_{k}+\sum_{k} \frac{\partial H}{\partial \dot{q}_{k}} \delta \dot{q}_{k}
$$

when written out in detail, or from (144) and (145):

$$
\begin{equation*}
\delta H=\sum_{k} \frac{d p_{k}}{d t} \delta q_{k}+\sum_{k} p_{k} \delta \dot{q}_{k} . \tag{146}
\end{equation*}
$$

If one further forms:

$$
\delta \sum_{k} p_{k} \dot{q}_{k}=\sum_{k} p_{k} \delta \dot{q}_{k}+\sum_{k} \dot{q}_{k} \delta p_{k}
$$

then it will follow that:

$$
\begin{equation*}
\sum_{k} p_{k} \delta \dot{q}_{k}=\delta \sum_{k} p_{k} \dot{q}_{k}-\sum_{k} \dot{q}_{k} \delta p_{k}, \tag{147}
\end{equation*}
$$

and when that is substituted in (146), that will give:

$$
\delta H=\sum_{k} \frac{d p_{k}}{d t} \delta q_{k}+\delta \sum_{k} p_{k} \dot{q}_{k}-\sum_{k} \dot{q}_{k} \delta p_{k},
$$

or

$$
\begin{equation*}
\delta\left(H-\sum p_{k} \dot{q}_{k}\right)=\sum_{k} \frac{d p_{k}}{d t} \delta q_{k}-\sum_{k} \frac{d q_{k}}{d t} \delta p_{k} . \tag{148}
\end{equation*}
$$

If we now consider the function $H-\sum p_{k} \dot{q}_{k}=R$ to depend upon the quantities $p_{k}$ and $q_{k}$ then we can further write the variation of $R$, which is considered to be a function of just those variables, as:

$$
\begin{equation*}
\delta R=\sum_{k} \frac{\partial R}{\partial p_{k}} \delta p_{k}+\sum_{k} \frac{\partial R}{\partial q_{k}} \delta q_{k}, \tag{149}
\end{equation*}
$$

and a comparison of that with (148) will yield the double system of equations:

$$
\begin{equation*}
\frac{d p_{k}}{d t}=\frac{\partial R}{\partial q_{k}}, \quad-\frac{d q_{k}}{d t}=\frac{\partial R}{\partial q_{k}} \tag{150}
\end{equation*}
$$

whose distinctive symmetry is apparent.
Equations (150) are Hamilton's equations of dynamics. They are equivalent to the Lagrange equations in the form (143). Under some circumstances, one can further simplify Hamilton's equations when the function $R$, viz., the modified kinetic potential, takes on a fundamental physical meaning, namely, that of minus the total energy. That will be the case if and only if the kinetic energy $L$ is a homogeneous quadratic form in the velocity components $\dot{q}_{k}$, which does not always need to be the case. One calls the cases in which that requirement is fulfilled "natural" problems. If we now assume that we are dealing with such a thing then, from (144), we will have:

$$
\sum p_{k} \dot{q}_{k}=\sum \frac{\partial H}{\partial \dot{q}_{k}} \dot{q}_{k}
$$

or, since $H=L-\Phi$, in which $\Phi$ does not depend upon $\dot{q}_{k}$ :

$$
\sum p_{k} \dot{q}_{k}=\sum \frac{\partial L}{\partial \dot{q}_{k}} \dot{q}_{k} .
$$

Moreover, since $L$ is now a homogeneous quadratic form, from Euler's theorem on homogeneous functions:

$$
\sum_{k} \frac{\partial L}{\partial \dot{q}_{k}} \dot{q}_{k}=2 L .
$$

Thus, one has $\sum p_{k} \dot{q}_{k}=2 L$, and $R$ will become equal to:

$$
\begin{equation*}
R=H-\sum_{k} p_{k} \dot{q}_{k}=L-\Phi-2 L=-(L+\Phi)=-E, \tag{151}
\end{equation*}
$$

in succession, when one lets $E$ denote the total energy, which is regarded as a function of the socalled canonical variables $p_{k}$ and $q_{k}$. Equations (150) will then become:

$$
\begin{equation*}
-\frac{d p_{k}}{d t}=+\frac{\partial E}{\partial q_{k}}, \quad \frac{d q_{k}}{d t}=+\frac{\partial E}{\partial p_{k}} . \tag{152}
\end{equation*}
$$

That is the form of Hamilton's equations that is ordinarily employed, namely, the so-called canonical form of the equations of dynamics. However, the domain of their validity is restricted by the demand that one must be dealing with a "natural" problem.

It would be useful to clarify the meaning of the canonical equations in the example of Cartesian coordinates. We take, say, a mass-point that is pulled back to its rest position by an elastic force. For it, we will have:

$$
L=\frac{1}{2} m\left(\dot{x}^{2}+\dot{y}^{2}+\dot{z}^{2}\right), \quad \Phi=\frac{1}{2} k^{2}\left(x^{2}+y^{2}+z^{2}\right) .
$$

The impulses are:

$$
p_{1}=m \dot{x}, \quad p_{2}=m \dot{y}, \quad p_{3}=m \dot{z}
$$

so $E$ will be a function of the impulses and coordinates:

$$
E=\frac{1}{2 m}\left(p^{2}+p^{2}+p^{2}\right)+\frac{k^{2}}{2}\left(x^{2}+y^{2}+z^{2}\right) .
$$

From (152), one will then have:

$$
-\frac{d p_{1}}{d t}=-m \ddot{x}=\frac{\partial E}{\partial x}=k^{2} x
$$

etc., and likewise:

$$
\frac{d q_{1}}{d t}=\frac{d x}{d t}=\frac{p_{1}}{m} .
$$

The second equation in (152) then defines the impulse, while the first one yields the actual equations of motion.

## § 14.

## General variation of Hamilton's principal function.

Previously, in the principle of stationary action, one was led to the integral:

$$
\begin{equation*}
S=\int_{t_{0}}^{t_{1}}(L-\Phi) d t=\int_{t_{0}}^{t_{1}} H d t \tag{153}
\end{equation*}
$$

under the assumption that a potential energy $\Phi$ existed. The principle of stationary action, in turn, indeed consisted of saying that $\delta$ process would produce the value zero when it is applied to that integral under certain assumptions about the behavior at the limits.

Hamilton had varied the function $S$, which he called the principal function, in a general manner without prescribing any conditions on the limits. At the same time, the variation of time was also permitted. In brief, he had then applied the $\Delta$ process to $S$ and achieved some results that were important to dynamics in that way.

We thus form $\Delta S=\Delta \int_{t_{0}}^{t_{1}} H d t$. It will then follow that:

$$
\Delta \int_{t_{0}}^{t_{1}} H d t=\int_{t_{0}}^{t_{1}} \Delta H d t+\int_{t_{0}}^{t_{1}} H d \Delta t
$$

or when we consider $H$ to be a function of the $q_{k}$ and $\dot{q}_{k}$ (we can drop the summation signs up to the conclusion of the calculation, for the sake of clarity):

$$
\Delta S=\int \frac{\partial H}{\partial q_{k}} \Delta q_{k} d t+\int \frac{\partial H}{\partial \dot{q}_{k}} \Delta \dot{q}_{k} d t+\int H d \Delta t .
$$

However, from equation (110), one will have:

$$
\Delta \dot{q}_{k}=\Delta\left(\frac{d q_{k}}{d t}\right)=\frac{d}{d t}\left(\Delta q_{k}\right)-\dot{q}_{k} \frac{d \Delta t}{d t}
$$

since the independent variable $t$ is varied along with the others under the $\Delta$ process, and it will further follow that:

$$
\Delta S=\int \frac{\partial H}{\partial q_{k}} \Delta q_{k} d t+\int \frac{\partial H}{\partial \dot{q}_{k}} \frac{d \Delta q_{k}}{d t} d t-\int \frac{\partial H}{\partial \dot{q}_{k}} \dot{q}_{k} d \Delta t+\int H d \Delta t,
$$

or, after combining the third and fourth terms:

$$
\begin{equation*}
\Delta S=\int \frac{\partial H}{\partial q_{k}} \Delta q_{k} d t+\int \frac{\partial H}{\partial \dot{q}_{k}} \frac{d \Delta q_{k}}{d t} d t+\int\left(H-\frac{\partial H}{\partial \dot{q}_{k}} \dot{q}_{k}\right) d \Delta t . \tag{154}
\end{equation*}
$$

The second, and then the third, integrals are now converted by partial integration:

$$
\begin{equation*}
\int \frac{\partial H}{\partial \dot{q}_{k}} \frac{d \Delta q_{k}}{d t} d t=\left[\frac{\partial H}{\partial \dot{q}_{k}} \Delta q_{k}\right]_{t_{0}}^{t_{1}}-\int \frac{d}{d t}\left(\frac{\partial H}{\partial \dot{q}_{k}}\right) \Delta q_{k} d t \tag{155}
\end{equation*}
$$

as well as:

$$
\begin{equation*}
\int\left(H-\frac{\partial H}{\partial \dot{q}_{k}} \dot{q}_{k}\right) d \Delta t=\left[\left(H-\frac{\partial H}{\partial \dot{q}_{k}} \dot{q}_{k}\right) \Delta t\right]_{t_{0}}^{t_{1}}-\int \frac{d}{d t}\left(H-\frac{\partial H}{\partial \dot{q}_{k}} \dot{q}_{k}\right) \Delta t d t . \tag{156}
\end{equation*}
$$

If one substitutes both results in equation (154) then one will get:

$$
\left\{\begin{align*}
\Delta S= & \int \frac{\partial H}{\partial q_{k}} \Delta q_{k} d t-\int \frac{d}{d t}\left(\frac{\partial H}{\partial \dot{q}_{k}}\right) \Delta q_{k} d t-\int \frac{d}{d t}\left(H-\frac{\partial H}{\partial \dot{q}_{k}} \dot{q}_{k}\right) \Delta t d t  \tag{157}\\
& +\left[\left(H-\frac{\partial H}{\partial \dot{q}_{k}} \dot{q}_{k}\right) \Delta t+\frac{\partial H}{\partial \dot{q}_{k}} \Delta q_{k}\right]_{t_{0}}^{t_{1}}
\end{align*}\right.
$$

The differentiation with respect to $t$ is performed in the third term:

$$
\begin{equation*}
\frac{d}{d t}\left(H-\frac{\partial H}{\partial \dot{q}_{k}} \dot{q}_{k}\right)=\frac{\partial H}{\partial q_{k}} \dot{q}_{k}-\frac{d}{d t}\left(\frac{\partial H}{\partial \dot{q}_{k}}\right) \dot{q}_{k} \tag{158}
\end{equation*}
$$

When combined with (157), that will give:

$$
\begin{align*}
\Delta S= & \int \frac{\partial H}{\partial q_{k}} \Delta q_{k} d t-\int \frac{d}{d t}\left(\frac{\partial H}{\partial \dot{q}_{k}}\right) \Delta q_{k} d t-\int \frac{\partial H}{\partial q_{k}} \dot{q}_{k} \Delta t d t \\
& +\int \frac{d}{d t}\left(\frac{\partial H}{\partial \dot{q}_{k}}\right) \dot{q}_{k} \Delta t d t+\left[\left(H-\frac{\partial H}{\partial \dot{q}_{k}} \dot{q}_{k}\right) \Delta t+\frac{\partial H}{\partial \dot{q}_{k}} \Delta q_{k}\right]_{t_{0}}^{t_{1}} . \tag{159}
\end{align*}
$$

When we now reinsert the summation sign everywhere, a suitable combination of that will further imply that:

$$
\Delta S=\int \sum_{k}\left[\frac{\partial H}{\partial q_{k}}-\frac{d}{d t}\left(\frac{\partial H}{\partial \dot{q}_{k}}\right)\right]\left\{\Delta q_{k}-\dot{q}_{k} \Delta t\right\} d t+\left[\left(H-\sum_{k} \frac{\partial H}{\partial \dot{q}_{k}} \dot{q}_{k}\right) \Delta t+\sum_{k} \frac{\partial H}{\partial \dot{q}_{k}} \Delta q_{k}\right]_{t_{0}}^{t_{1}}
$$

Now, since $\Delta q_{k}-\dot{q}_{k} \Delta t=\delta q_{k}$, that can now be written more simply as:

$$
\begin{equation*}
\Delta S=\int \sum_{k}\left[\frac{\partial H}{\partial q_{k}}-\frac{d}{d t}\left(\frac{\partial H}{\partial \dot{q}_{k}}\right)\right] \delta q_{k} d t+\left[\left(H-\sum_{k} \frac{\partial H}{\partial \dot{q}_{k}} \dot{q}_{k}\right) \Delta t+\sum_{k} \frac{\partial H}{\partial \dot{q}_{k}} \Delta q_{k}\right]_{t_{0}}^{t_{1}} . \tag{160}
\end{equation*}
$$

That is Hamilton's fundamental formula, upon which his further conclusions were based.

## § 15.

## Hamilton's differential equation for the principal function and the integrals of the equation of motion.

We would now like to assume that the coordinates $q_{k}$ are mutually independent. Obviously, that means that the system is holonomic, because for such a thing, that can always be achieved by a suitable choice of coordinates. Lagrange's equations are then valid, and the first integral in (160) will vanish. What will remain is:

$$
\begin{equation*}
\Delta S=\left[\left(H-\sum_{k} \frac{\partial H}{\partial \dot{q}_{k}} \dot{q}_{k}\right) \Delta t+\sum_{k} \frac{\partial H}{\partial \dot{q}_{k}} \Delta q_{k}\right]_{t_{0}}^{t_{1}} . \tag{161}
\end{equation*}
$$

$H-\sum_{k} \frac{\partial H}{\partial \dot{q}_{k}} \dot{q}_{k}=H-\sum_{k} p_{k} \dot{q}_{k}$ is nothing but the function $R$ that was introduced in $\S$ 13. If we then assume that we are dealing with a "natural" problem then $R=-E$, in which $E$ means the total energy. With that, (161) will become:

$$
\begin{equation*}
\Delta S=E_{0} \Delta t_{0}-E_{1} \Delta t_{1}-\sum p_{k}^{1} \Delta q_{k}^{1}-\sum p_{k}^{0} \Delta q_{k}^{0} \tag{162}
\end{equation*}
$$

when written out in detail. In that, we have denoted the values of the impulse components at times $t_{0}$ and $t_{1}$ by $p_{k}^{0}$ and $p_{k}^{1}$, resp., and the indices on the other quantities mean something analogous.

Naturally, we again associate the true path with a varied path by way of the $\Delta$ symbol, which has not been restricted in any way up to now. However, following Hamilton, we would now like to establish the following:

1. The neighboring path shall likewise begin at the time-point $t=t_{0}$, i.e., we shall always set $\Delta t_{0}=0$ in what follows.
2. The neighboring paths shall be generated only by varying the initial conditions $q_{k}^{0}$ and $p_{k}^{0}$ on the true path, i.e., we imagine that a family of varied paths exists that are subject to the same forces and constraints as the true path, except that the initial values are not $q_{k}^{0}$ and $p_{k}^{0}$, but $q_{k}^{0}+\Delta q_{k}^{0}$ and $p_{k}^{0}+\Delta p_{k}^{0}$, resp.

When we now drop the upper index of 1 , for simplicity's sake, (162) will then become:

$$
\begin{equation*}
\Delta S=-E \Delta t+\sum_{k} p_{k} \Delta q_{k}-\sum p_{k}^{0} \Delta q_{k}^{0} . \tag{163}
\end{equation*}
$$

We would now like to express $\Delta S$ in a certain way. The integrals of our system of $N$ degrees of freedom exhibit the coordinates $q_{k}$ and velocities $\dot{q}_{k}$ as functions of time $t$, the initial coordinates $q_{k}^{0}$, and the initial velocities $\dot{q}_{k}^{0}$ (the initial impulses $p_{k}^{0}$, resp.):

$$
\begin{align*}
& q_{k}=q_{k}\left(q_{k}^{0}, p_{k}^{0}, t\right),  \tag{164}\\
& \dot{q}_{k}=\dot{q}_{k}\left(q_{k}^{0}, p_{k}^{0}, t\right) . \tag{165}
\end{align*}
$$

The true motion will be determined completely by that. From the above, the varied motions will emerge from that when we substitute the values $q_{k}^{0}+\Delta q_{k}^{0}$ and $p_{k}^{0}+\Delta p_{k}^{0}$, in place of $q_{k}^{0}$ and $p_{k}^{0}$, resp. With those equations, we can express the $q_{k}$ and $\dot{q}_{k}$ as functions of time, the initial coordinates $q_{k}^{0}$, and the initial impulse $p_{k}^{0}$, and substitute those functions in the principal function, which was naturally considered to be a function of $q_{k}$ and $\dot{q}_{k}$ up to now. By means of the integrals, which are assumed to be known, we will then get $S$ as a function of the initial coordinates and impulses, as well as time $t$. Finally, we once more eliminate the initial impulses from $S$, and likewise by means of the integral (164). Indeed, they represent $N$ relations between the $(3 N+1)$ quantities $q_{k}, q_{k}^{0}, p_{k}^{0}, t$. We can then obtain the $N$ quantities $p_{k}^{0}$ as functions of the quantities $q_{k}, q_{k}^{0}, t$, i.e., the instantaneous and initial coordinates and time, with the help of those equations, substitute that in $S$, and ultimately eliminate the $p_{k}^{0}$ from $S$ in that way. After those operations, $S$ will be represented as something that depends upon the quantities $q_{k}^{0}$, the initial coordinates, the $q_{k}$, the instantaneous coordinates, and time $t$. Following Boltzmann, we would like to call the function $S$, when it is represented in that way, the principal function, as represented in the "Hamiltonian way," and denote it by the index $H$ :

$$
\begin{equation*}
S\left(q_{k}, q_{k}^{0}, t\right) \equiv S_{H} . \tag{166}
\end{equation*}
$$

$E$ will be represented similarly, and that is why we will also assign it the index $H$ accordingly.

If we now go from the true path to the varied path in the manner that was described above then $S$ will go to $S+\Delta S$, since the $q_{k}^{0}$ will be converted into $q_{k}^{0}+\Delta q_{k}^{0}$ and the $p_{k}^{0}$, into $p_{k}^{0}+\Delta p_{k}^{0}$, while the upper limit $t$ will be converted into $t+\Delta t$.

From (166), $\Delta S_{H}$ can obviously be written:

$$
\begin{equation*}
\Delta S_{H}=\frac{\partial S_{H}}{\partial t} \Delta t+\sum_{k} \frac{\partial S_{H}}{\partial q_{k}} \Delta q_{k}+\sum_{k} \frac{\partial S_{H}}{\partial q_{k}^{0}} \Delta q_{k}^{0} . \tag{167}
\end{equation*}
$$

A comparison of that with (163) will then imply the following three equations:

$$
\left.\begin{array}{l}
\frac{\partial S_{H}}{\partial t}+E=0, \\
\frac{\partial S_{H}}{\partial q_{k}}=p_{k},  \tag{169}\\
\frac{\partial S_{H}}{\partial q_{k}^{0}}=-p_{k}^{0} .
\end{array}\right\} \quad(k=1,2, \ldots, N),
$$

From the foregoing, the partial derivatives of $S_{H}$ are understood as follows: $\partial S_{H} / \partial t$ formed with constant values of $q_{k}$ and $q_{k}^{0}$. In that case, $\left(\Delta S_{H}\right)_{q_{k}, q_{k}^{0}}$ is non-zero only due to the fact that the upper limit $t$ has been converted into $t+\Delta t$. The same is true of the remaining partial derivatives.

What do equations (168) to (170) mean then?
Equations (168) is a partial differential equation for $S_{H}$, namely, the so-called Hamilton differential equation. It will be satisfied identically by the function $S_{H}$ that we have constructed from the integrals of the equations of motion, which we have assumed to be known. $S_{H}$ includes the $N$ arbitrary, non-additive constants $q_{k}^{0}$, and an additive constant can be added to it, in addition, such that we will have $(N+1)$ independent constants in all, which is exactly as many as the independent variables in $\left(q_{k}, t\right)$. Due to that property, one calls $S_{H}$ a complete integral of the partial differential equation.

The derivatives $\partial S_{H} / \partial q_{k}$ and $\partial S / \partial q_{k}^{0}$ include no differential quotients with respect to time. By contrast, $p_{k}$ is a linear combination of the general velocity components $\dot{q}_{k}$, so it will then include the first derivatives with respect to $t$; the $p_{k}^{0}$ are constants. We can then characterize equations (169) as follows:

They are $N$ equations that express the velocities $\dot{q}_{k}$ in terms of constants and time: They are then the $N$ first ("intermediate") integrals of the equations of motion. Likewise, equations (170) are $N$ relations that express the general coordinates $q_{k}$ in terms of constants and time: They are $N$ second integrals of motion.

It might be good for us to explain the type of substitutions and constructions that have occurred here by the simplest-possible example. We deal with the free case. In that case, the integrals are known to read:

$$
\left\{\begin{array}{l}
\dot{x}=\dot{x}_{0}-g t,  \tag{171}\\
x=x_{0}+\dot{x}_{0} t-\frac{1}{2} g t^{2} .
\end{array}\right.
$$

The kinetic and potential energy will then be:

$$
L=\frac{1}{2} m \dot{x}^{2}=\frac{1}{2} m\left(\dot{x}_{0}^{2}+\dot{y}_{0}^{2}+\dot{z}_{0}^{2}\right),
$$

$$
\Phi=m g x=m g\left(x_{0}+\dot{x}_{0} t-\frac{1}{2} g t^{2}\right),
$$

while the total energy $E$ and the kinetic potential $\Phi$ are:

$$
\left\{\begin{array}{l}
E=L+\Phi=\frac{1}{2} m \dot{x}_{0}^{2}+m g x_{0},  \tag{172}\\
H=L-\Phi=\left(\frac{1}{2} m \dot{x}_{0}^{2}-m g x_{0}\right)-2 m g \dot{x}_{0} t+m g^{2} t^{2} .
\end{array}\right.
$$

Finally, when we set the lower limit $t_{0}=0$, for simplicity, $S=\int H d t$ will become:

$$
\begin{equation*}
S=\left(\frac{1}{2} m \dot{x}_{0}^{2}-m g x_{0}\right) t-m g \dot{x}_{0} t^{2}+\frac{1}{3} m g^{2} t^{3} . \tag{173}
\end{equation*}
$$

However, neither $E$ nor $S$ have been expressed in the Hamiltonian way up to now. To that end, we must represent the $\dot{x}_{0}$ in terms of $x$ and $t$ by using the second equation in (171):

$$
\begin{equation*}
\dot{x}_{0}=\frac{x-x_{0}+\frac{1}{2} g t^{2}}{t} . \tag{174}
\end{equation*}
$$

With that, we will then get:

$$
\begin{aligned}
& E_{H}=\frac{1}{2} m \frac{\left(x-x_{0}+\frac{1}{2} g t^{2}\right)^{2}}{t^{2}}+m g x_{0}, \\
& S_{H}=\frac{m}{2 t}\left(x-x_{0}+\frac{1}{2} g t^{2}\right)^{2}-m g x_{0} t-m g t\left(x-x_{0}+\frac{1}{2} g t^{2}\right)+\frac{1}{3} m g^{2} t^{2} .
\end{aligned}
$$

Those values of $S_{H}$ and $E_{H}$ must fulfill Hamilton's partial differential equation (168). It is:

$$
\frac{\partial S}{\partial t}=m\left(x-x_{0}+\frac{1}{2} g t^{2}\right) g-\frac{m}{2 t^{2}}\left(x-x_{0}+\frac{1}{2} g t^{2}\right)^{2}-m g x_{0}-m g\left(x-x_{0}\right)-\frac{3}{2} m g^{2} t^{2}+m g^{2} t^{2} .
$$

If one now forms equation (168) with the value of $E_{H}$ above then one will find that:

$$
\begin{aligned}
\left\{m g\left(x-x_{0}+\frac{1}{2} g t^{2}\right)\right. & \left.-\frac{m}{2 t^{2}}\left(x-x_{0}+\frac{1}{2} g t^{2}\right)^{2}-m g x_{0}-m g x+m g x_{0}-\frac{1}{2} m g^{2} t^{2}\right\} \\
+ & \left\{\frac{m}{2 t^{2}}\left(x-x_{0}+\frac{1}{2} g t^{2}\right)^{2}+m g x_{0}\right\}=0
\end{aligned}
$$

and that equation will be fulfilled identically since the terms on the left-hand side cancel pairwise.

Similarly, one will have, e.g., with equation (170): $\frac{\partial S_{H}}{\partial x_{0}}=-m \dot{x}_{0}$.

One will find that:

$$
\frac{\partial S_{H}}{\partial x_{0}}=-\frac{m}{t}\left(x-x_{0}+\frac{1}{2} g t^{2}\right),
$$

and from (174), that is, in fact, equal to $-m \dot{x}_{0}$, which was to be proved.

## § 16.

## Jacobi's converse of Hamilton's theorem on the actual determination of the integrals.

How can one apply those arguments then? If we have only Hamilton's theorem, which is included in equations (168) to (170), then its usefulness would be very limited in terms of obtaining the integrals of the equations of motion, since we had to assume they existed in order to arrive at the theorem.

We would like to see what we can exhibit when we do not possess those integrals.
We know the expression for the kinetic energy as a function of the $q_{k}$ and $\dot{q}_{k}$, or $q_{k}$ and $p_{k}$, respectively, since we have $\partial L / \partial \dot{q}_{k}=p_{k}$; we likewise know $\Phi$ as a function of the $q_{k}$. We can then construct the total energy $E=L+\Phi$, as well as the kinetic potential $H=L-\Phi$ as a function of the $q_{k}$. However, we can eliminate the $p_{k}$ from $E$ by employing equation (169), i.e., we can set $p_{k}=\partial S / \partial q_{k}$. We will then get $E$ as a function of the $q_{k}$ and $\partial S / \partial q_{k}$, which we would like to suggest by $E\left(q_{k}, \frac{\partial S}{\partial q_{k}}\right)$.

Hamilton's partial differential equation (168) for $S$ will then read:

$$
\begin{equation*}
\frac{\partial S}{\partial t}+E\left(q_{k}, \frac{\partial S}{\partial q_{k}}\right)=0 \tag{175}
\end{equation*}
$$

We shall drop the index $H$ from now on.
Up to now, we could calculate only one solution that satisfied that equation identically from the known integrals of the equations of motion, and indeed a complete solution with $N$ independent constants $q_{k}^{0}$. Now, when the integrals are unknown, the same will be true for $S$, and we must employ the differential equation (175) precisely in order to determine $S$.

Now, Jacobi ${ }^{1}{ }^{1}$ ) proved the following theorem:
If an arbitrary complete integral of equation (175) is known, i.e., a solution with $N$ independent, arbitrary, multiplicative constants $\alpha_{1}, \alpha_{2}, \ldots, \alpha_{N}$ (which do not, by any means, need to be identical to the $q_{k}^{0}$ ), then one will get the integrals of the motion when one forms equations (169) and (170) with that arbitrary solution, i.e., one determines the impulses $p_{k}$ by means of the equations:

$$
\begin{equation*}
\frac{\partial S}{\partial q_{k}}=p_{k} \quad(k=1,2, \ldots, N) \tag{176}
\end{equation*}
$$

[^8](they are the $N$ first or intermediate integrals), and one further sets the derivatives of $S$ with respect to the constants $\alpha_{k}$ equal to $N$ new arbitrary constants $\beta_{k}$ :
\[

$$
\begin{equation*}
\frac{\partial S}{\partial \alpha_{k}}=\beta_{k} \quad(k=1,2, \ldots, N) . \tag{177}
\end{equation*}
$$

\]

Those are the $N$ second integrals of the motion.
In the cases that occur quite frequently in which $E$ does not depend upon time explicitly, one can simplify that result when one "separates" the time variable $t$ from $S$. In that case, one sets:

$$
S=-\alpha_{1} t+W\left(q_{k}, \alpha_{2}, \alpha_{3}, \ldots, \alpha_{N}\right),
$$

in which $\alpha_{1}$ is one of the constants, and $W$ now includes only $(N-1)$ of the independent constants $\alpha_{1}, \alpha_{2}, \ldots, \alpha_{N}$. One will then have:

$$
\begin{gathered}
\frac{\partial S}{\partial t}=-\alpha_{1}, \quad \frac{\partial S}{\partial q_{k}}=\frac{\partial W}{\partial q_{k}}, \quad \frac{\partial S}{\partial \alpha_{k}}=\frac{\partial W}{\partial \alpha_{k}} \quad(k=2, \ldots, N), \\
\frac{\partial S}{\partial \alpha_{1}}=-t+\frac{\partial W}{\partial \alpha_{1}} .
\end{gathered}
$$

With that, Hamilton's partial differential equation (175) for $S$ will go to the following one for W:

$$
\begin{equation*}
E\left(q_{k}, \frac{\partial W}{\partial q_{k}}\right)=\alpha_{1} \tag{178}
\end{equation*}
$$

and Jacobi's integral equations (176) and (177) will then become:

$$
\begin{array}{ll}
\frac{\partial W}{\partial q_{k}}=p_{k} & (N \text { equations }), \\
\frac{\partial W}{\partial \alpha_{k}}=\beta_{k} & (N-1 \text { equations }), \\
\frac{\partial W}{\partial \alpha_{1}}=t+\beta_{1} & \tag{181}
\end{array}
$$

Before we go on to prove Jacobi's theorem, we would like to clarify its meaning in an example that is as simple as possible.

We shall treat the linear oscillation of a mass-point about a rest position under the influence of an elastic force. We are then given at the outset:

$$
\begin{aligned}
& L=\frac{1}{2} m \dot{x}^{2}=\frac{1}{2 m} p^{2}, \\
& \Phi=\frac{1}{2} k^{2} x^{2}, \\
& E=L+\Phi=\frac{1}{2 m} p^{2}+\frac{1}{2} k^{2} x^{2} .
\end{aligned}
$$

Since $E$ does not depend upon $t$, we can use the function $W$ everywhere, instead of $S$, and write:

$$
E=\frac{1}{2 m} p^{2}+\frac{1}{2} k^{2} x^{2}=\frac{1}{2 m}\left(\frac{d W}{d x}\right)^{2}+\frac{1}{2} k^{2} x^{2}
$$

Thus, from equation (178), the differential equation for $W$ will read:

$$
\frac{1}{2 m}\left(\frac{d W}{d x}\right)^{2}+\frac{1}{2} k^{2} x^{2}=\alpha_{1}
$$

Here, we need an integral with only one arbitrary multiplicative constant, and that constant is already given by $\alpha_{1}$. An additive constant will no longer need to be added then. We will then get:

$$
\left(\frac{d W}{d x}\right)^{2}=2 m \alpha_{1}\left(1-\frac{k^{2}}{2 \alpha} x^{2}\right)
$$

so

$$
\frac{d W}{d x}=\sqrt{2 m \alpha_{1}} \cdot \sqrt{1-\frac{k^{2}}{2 \alpha} x^{2}}
$$

Integration will yield:

$$
W=\frac{\alpha_{1} \sqrt{m}}{k}\left[\arcsin \frac{k}{\sqrt{2 \alpha_{1}}}+\frac{k x}{\sqrt{2 \alpha_{1}}} \cdot \sqrt{1-\frac{k^{2}}{2 \alpha} x^{2}}\right] .
$$

If we correspondingly differentiate (181) with respect to the constant $\alpha_{1}$ and set $\partial W / \partial \alpha_{1}-t$ equal to a new constant $\beta_{1}$ then it will follow from (181) that:

$$
\beta_{1}+t=\frac{\sqrt{m}}{k} \arcsin \frac{k x}{\sqrt{2 \alpha_{1}}}
$$

or

$$
\frac{k}{\sqrt{m}}\left(\beta_{1}+t\right)=\arcsin \frac{k}{\sqrt{2 \alpha_{1}}} x
$$

or

$$
x=\frac{\sqrt{2 \alpha_{1}}}{k} \sin \frac{k}{\sqrt{m}}\left(\beta_{1}+t\right)
$$

which is, in fact, the desired integral.
Naturally, in such a simple case, one would not appeal to the Hamilton-Jacobi method in order to find the integrals of the equation. That example should serve only to make it easier to understand the theory.

We now move on to the proof of Jacobi's theorem!
We have to show that the due to the equations:

$$
\frac{\partial S}{\partial q_{k}}=p_{k} \quad \text { and } \quad \frac{\partial S}{\partial \alpha_{k}}=\beta_{k} \quad(k=2, \ldots, N)
$$

in which $S$ is a complete integral of Hamilton's partial differential equation with $N$ arbitrary independent constants $\alpha_{k}$, the equations of motion, which we would like to take in, say, Hamilton's canonical form:

$$
\begin{equation*}
\frac{\partial E}{\partial p_{k}}=\dot{q}_{k}, \quad-\frac{\partial E}{\partial q_{k}}=\dot{p}_{k}, \tag{182}
\end{equation*}
$$

will be satisfied.
We start from equation (177): $\frac{\partial S}{\partial \alpha_{k}}=\beta_{k}$ and form the complete temporal derivative $\frac{d}{d t}\left(\frac{\partial S}{\partial \alpha_{k}}\right)$, whose value is zero, for all values of $k(=1, \ldots, N)$. Since $S$ is a function $t, q_{k}$, and $\alpha_{k}$, we will then have the following system of equations:

$$
\left\{\begin{array}{l}
\frac{\partial^{2} S}{\partial \alpha_{1} \partial t}+\frac{\partial^{2} S}{\partial \alpha_{1} \partial q_{1}} \dot{q}_{1}+\frac{\partial^{2} S}{\partial \alpha_{1} \partial q_{2}} \dot{q}_{2}+\cdots+\frac{\partial^{2} S}{\partial \alpha_{1} \partial q_{N}} \dot{q}_{N}=0  \tag{183}\\
\cdots \ldots \ldots \ldots \ldots \ldots \ldots \ldots \ldots \ldots \ldots \ldots \ldots \ldots \ldots \ldots \ldots \ldots \ldots \ldots \ldots \ldots \ldots \ldots \ldots \ldots \ldots \ldots \ldots \ldots \ldots \ldots \ldots \ldots \ldots \ldots \ldots \ldots \ldots \ldots \ldots \ldots \ldots \ldots \ldots \ldots \ldots \\
\frac{\partial^{2} S}{\partial \alpha_{N} \partial t}+\frac{\partial^{2} S}{\partial \alpha_{N} \partial q_{1}} \dot{q}_{1}+\frac{\partial^{2} S}{\partial \alpha_{N} \partial q_{2}} \dot{q}_{2}+\cdots+\frac{\partial^{2} S}{\partial \alpha_{N} \partial q_{N}} \dot{q}_{N}=0
\end{array}\right.
$$

We have to calculate the $\dot{q}_{k}$ from those equations and substitute that in the first canonical equation in (182): $\frac{\partial E}{\partial p_{k}}=\dot{q}_{k}$, which must then be satisfied identically. However, we can also, conversely, substitute the values $\frac{\partial E}{\partial p_{k}}$ for $\dot{q}_{k}$ that come from the canonical equation (182) in (183) and show that (183) will be satisfied in that way. We write the system (183) more simply as:

$$
\begin{equation*}
\frac{\partial^{2} S}{\partial t \partial \alpha_{k}}+\sum_{i} \frac{\partial^{2} S}{\partial \alpha_{k} \partial q_{i}} \dot{q}_{i}=0 \quad(k=1,2, \ldots, N), \tag{184}
\end{equation*}
$$

and when we make the stated substitution for $\dot{q}_{k}$, we will have:

$$
\begin{equation*}
\frac{\partial^{2} S}{\partial \alpha_{k} \partial t}+\sum_{i} \frac{\partial^{2} S}{\partial \alpha_{k} \partial q_{i}} \cdot \frac{\partial E}{\partial p_{i}}=0 \quad(k=1,2, \ldots, N) \tag{185}
\end{equation*}
$$

Since that equation is an identity, the following can be shown: $S$ is a solution of Hamilton's partial differential equation:

$$
\frac{\partial S}{\partial t}+E\left(q_{k}, \frac{\partial S}{\partial q_{k}}\right) \equiv 0
$$

which is satisfied identically for arbitrary values of the constants $\alpha_{k}$. If we differentiate with respect to $\alpha_{k}$ then it will follow that:

$$
\frac{\partial^{2} S}{\partial t \partial \alpha_{k}}+\frac{\partial E}{\partial \alpha_{k}} \equiv 0
$$

Now, the $\alpha_{k}$ are found in the expressions $\partial S / \partial q_{k}$ that are included in $E$. thus:

$$
\frac{\partial E}{\partial \alpha_{k}}=\sum_{i} \frac{\partial E}{\partial\left(\frac{\partial S}{\partial q_{i}}\right)} \frac{\partial^{2} S}{\partial q_{i} \partial \alpha_{k}} \equiv 0
$$

However, since one has $\partial S / \partial q_{i}=p_{i}$, according to Jacobi, that can also be written:

$$
\frac{\partial^{2} S}{\partial t \partial \alpha_{k}}+\sum_{i} \frac{\partial^{2} S}{\partial q_{i} \partial \alpha_{k}} \cdot \frac{\partial E}{\partial p_{i}} \equiv 0,
$$

which is satisfied identically and agrees with (185).
One can proceed similarly with the other system of equations. One starts from $\frac{\partial S}{\partial q_{k}}=p_{k}$ and forms $\frac{d}{d t}\left(\frac{\partial S}{\partial q_{k}}\right)$. That will produce a system for the $\dot{p}_{k}$ that is analogous to equation (183) and can be combined with the second canonical equation (182) in the same way. Those calculations are entirely analogous to the ones that were just represented and will be omitted here.

With that, Jacobi's theorem is proved, and we shall now address the question of how the integration of the partial differential equation for $S$ is to be achieved in practice.

## On the integration of Hamilton's partial differential equation.

The integration of Hamilton's partial differential equation encounters difficulties that can be overcome in only special cases. Namely, under certain conditions, one can separate all variables, just as we could previously separate the time variable $t$, i.e., one can decompose the function $W$ that remains after separating the time variable from $S$ and is a function of all $q_{k}$ into a sum of functions, each of which depends upon only one variable. The partial differential equation will then go to a series of ordinary ones. Under those suitable conditions, one can then set:

$$
\begin{equation*}
W\left(q_{1}, q_{2}, \ldots, q_{N}\right)=W\left(q_{1}\right)+W\left(q_{2}\right)+\ldots+W\left(q_{N}\right) . \tag{186}
\end{equation*}
$$

The question of when a Hamiltonian equation admits a separation of variables as in (186) has not been solved. However, Stäckel ( ${ }^{1}$ ), supported by previous work by Liouville and Staude, has succeeded in solving the problem under the assumption that the function $E\left(q_{k}, \frac{\partial W}{\partial q_{k}}\right)$, which is indeed of degree two in the $\frac{\partial W}{\partial q_{k}}$, includes only the squares of those quantities, but not their double products. Stäckel's assumption was then that $E$, which we would like to split into $L$ and $\Phi$ for the moment, can be written:

$$
\begin{equation*}
E\left(q_{k}, \frac{\partial W}{\partial q_{k}}\right)=\frac{1}{2} \sum_{k=1}^{N} A_{k}\left(q_{k}\right) \cdot\left(\frac{\partial W}{\partial q_{k}}\right)^{2}+\Phi\left(q_{k}\right) . \tag{187}
\end{equation*}
$$

As Stäckel showed, one can then write out the general solution to the mechanical problem immediately. The problems in which a separation of variables is possible on the basis of Stäckel's theorem have had a fundamental significance in recent times in the so-called quantum theory through the work of Schwarzschild $\left({ }^{2}\right)$ and Epstein $\left({ }^{3}\right)$.

For that reason, we would like to discuss at least one example of the separation of variables.
We again deal with a mass-point that performs elastic oscillations about its rest position in the $x y$-plane. It is therefore a generalization of the problem that we treated in the previous section. Now, for that case, we have:

$$
L=\frac{1}{2} m \dot{x}^{2}+\frac{1}{2} m \dot{y}^{2}=\frac{1}{2 m} p_{1}^{2}+\frac{1}{2 m} p_{2}^{2},
$$

[^9]$$
\Phi=\frac{1}{2} k^{2} x^{2}+\frac{1}{2} k^{2} y^{2} .
$$

For the total energy, it will follow that:

$$
E\left(p_{k}, q_{k}\right)=\frac{1}{2 m} p_{1}^{2}+\frac{1}{2 m} p_{2}^{2}+\frac{1}{2} k^{2} x^{2}+\frac{1}{2} k^{2} y^{2},
$$

and since that does not include $t$ explicitly, we can use the simplified differential equation for the function $W$ that is in equation (178). For our problem, it will read:

$$
\begin{equation*}
\frac{1}{2 m}\left(\frac{\partial W}{\partial x}\right)^{2}+\frac{1}{2 m}\left(\frac{\partial W}{\partial y}\right)^{2}+\frac{1}{2} k^{2} x^{2}+\frac{1}{2} k^{2} y^{2}=\alpha_{1} \tag{188}
\end{equation*}
$$

From the above, the following Ansatz is possible here with no further analysis:

$$
\begin{equation*}
W(x, y)=W_{1}(x)+W_{2}(y), \tag{189}
\end{equation*}
$$

which will make (188) go to:

$$
\begin{equation*}
\frac{1}{2 m}\left(\frac{d W_{1}}{d x}\right)^{2}+\frac{1}{2 m}\left(\frac{d W_{2}}{d y}\right)^{2}+\frac{1}{2} k^{2} x^{2}+\frac{1}{2} k^{2} y^{2}=\alpha_{1} \tag{190}
\end{equation*}
$$

and we can perhaps set:

$$
\begin{equation*}
\frac{1}{2 m}\left(\frac{d W_{1}}{d x}\right)^{2}+\frac{1}{2} k^{2} x^{2}=\frac{\alpha_{2}^{2}}{2 m} \tag{191}
\end{equation*}
$$

in which $\alpha_{2}$ is a new constant. What will then remain is the differential equation for $W_{2}$ :

$$
\begin{equation*}
\frac{1}{2 m}\left(\frac{d W_{1}}{d y}\right)^{2}+\frac{1}{2} k^{2} y^{2}=\alpha_{1}-\frac{\alpha_{2}^{2}}{2 m} \tag{192}
\end{equation*}
$$

We now have only the two ordinary differential equations (191) and (192) to deal with then. We find immediately from (191) that:

$$
\frac{d W_{1}}{d x}=\alpha_{2} \sqrt{1-\frac{m k^{2}}{\alpha_{2}^{2}} x^{2}}
$$

which can be integrated with no further assumptions:

$$
\begin{equation*}
W_{1}(x)=\frac{\alpha_{2}^{2}}{2 k \sqrt{m}}\left[\arcsin \frac{k \sqrt{m}}{\alpha_{2}} x+\frac{k \sqrt{m}}{\alpha_{2}} x \sqrt{1-\frac{m k^{2}}{\alpha_{2}^{2}} x^{2}}\right] \tag{193}
\end{equation*}
$$

It likewise follows from (192) that:

$$
\frac{d W_{2}}{d y}=\sqrt{2 m \alpha_{2}-\alpha_{2}^{2}} \cdot \sqrt{1-\frac{m k^{2} y^{2}}{2 m \alpha_{1}-\alpha_{2}^{2}}},
$$

and the integration likewise yields:

$$
\begin{equation*}
W_{2}(x)=\frac{2 m \alpha_{1}-\alpha_{2}^{2}}{2 k \sqrt{m}}\left[\arcsin \frac{k \sqrt{m}}{\sqrt{2 m \alpha_{1}-\alpha_{2}^{2}}} y+\frac{k \sqrt{m}}{\sqrt{2 m \alpha_{1}-\alpha_{2}^{2}}} y \sqrt{1-\frac{k^{2} m y^{2}}{2 m \alpha_{1}-\alpha_{2}^{2}}}\right] \tag{194}
\end{equation*}
$$

It will then follow that $W$ is:

$$
\begin{align*}
W(x, y) & =\frac{\alpha_{2}^{2}}{2 k \sqrt{m}}\left[\arcsin \frac{k \sqrt{m}}{\alpha_{2}} x+\frac{k \sqrt{m}}{\alpha_{2}} x \sqrt{1-\frac{m k^{2}}{\alpha_{2}^{2}} x^{2}}\right]  \tag{195}\\
& +\frac{2 m \alpha_{1}-\alpha_{2}^{2}}{2 k \sqrt{m}}\left[\arcsin \frac{k \sqrt{m}}{\sqrt{2 m \alpha_{1}-\alpha_{2}^{2}}} y+\frac{k \sqrt{m}}{\sqrt{2 m \alpha_{1}-\alpha_{2}^{2}}} y \sqrt{1-\frac{k^{2} m y^{2}}{2 m \alpha_{1}-\alpha_{2}^{2}}}\right]
\end{align*}
$$

We will now find the integrals from equations (180) and (181) when we differentiate the constants $\alpha_{1}$ and $\alpha_{2}$ and set the results equal to $t+\beta_{1}$ and $\beta_{2}$, resp., where $\beta_{1}$ and $\beta_{2}$ are two new constants. We will then find, by a simple, but somewhat tedious, calculation, that:

$$
\frac{\partial W}{\partial \alpha_{1}}=\beta_{1}+t=\frac{\sqrt{m}}{k} \arcsin \frac{k \sqrt{m}}{\sqrt{2 m \alpha_{1}-\alpha_{2}^{2}}} y,
$$

or

$$
\begin{equation*}
y=\frac{\sqrt{2 m \alpha_{1}-\alpha_{2}^{2}}}{k \sqrt{m}} \sin \frac{k}{\sqrt{m}}\left(t+\beta_{1}\right) \tag{196}
\end{equation*}
$$

with which one of the integrals has been found, and it actually includes two arbitrary constants (viz., an amplitude constant and a phase constant), as it must.

Likewise, one will have:

$$
\frac{\partial W}{\partial \alpha_{2}}=\frac{\alpha_{2}}{k \sqrt{m}}\left[\arcsin \frac{k \sqrt{m}}{\alpha_{2}} x-\arcsin \frac{k \sqrt{m}}{\sqrt{2 m \alpha_{1}-\alpha_{2}^{2}}} y\right]=\beta_{2}
$$

from which the following relation between $x$ and $y$ will arise by a simple calculation, which is then the equation of the path:

$$
\begin{equation*}
\frac{k^{2} m}{\alpha_{2}^{2}} x^{2}+\frac{2 k^{2} m}{\alpha_{2} \sqrt{2 m \alpha_{1}-\alpha_{2}^{2}}} x y+\frac{k^{2} m}{2 m \alpha_{1}-\alpha_{2}^{2}} y^{2}=\sin ^{2}\left(\frac{k \sqrt{m} \beta_{2}}{\alpha_{2}}\right), \tag{197}
\end{equation*}
$$

i.e., the path is an ellipse, as is known. One also finds $x$ as a function of $t$ directly upon combining that with (196), which is the ordinary representation of the second integral:

$$
\begin{equation*}
x=\frac{\alpha_{2}}{k \sqrt{m}} \cdot \sin \frac{k}{\sqrt{m}}\left(t+\beta_{1}+\frac{m \beta_{2}}{\alpha_{2}}\right), \tag{198}
\end{equation*}
$$

which likewise includes arbitrary constants. The fact that equation (179) is also fulfilled by the impulses is easy to see.

One finds other examples of the separation of variables in the cited papers by Epstein, as well as in Debye ( ${ }^{1}$ ) and Scherrer ( ${ }^{2}$ ), and they merit special interest due to the fact that they point to the significance of Hamilton-Jacobi theory for one of the burning questions of modern physics, as was mentioned before.

[^10]
## § 18.

## Gauss's principle of least constraint.

Gauss ${ }^{(1)}$ ) has added a new principle to dynamics that serves to achieve the same thing as d'Alembert's and the principles that are equivalent to it that we have treated already.

We can characterize the true motion by giving the coordinates $x_{v}, y_{v}, z_{v}$ of the $v^{\text {th }}$ masspoints as functions of time:

$$
x_{\nu}=\psi_{v}(t), \quad y_{v}=\chi_{v}(t), \quad z_{v}=\omega_{v}(t) .
$$

Those expressions constitute the true path. The true velocities and accelerations are:

$$
\begin{array}{lll}
\dot{x}_{v}=\dot{\psi}_{v}(t), & \dot{y}_{v}=\dot{\chi}_{v}(t), & \dot{z}_{v}=\dot{\omega}_{v}(t), \\
\ddot{x}_{v}=\ddot{\psi}_{v}(t), & \ddot{y}_{v}=\ddot{\chi}_{v}(t), & \ddot{z}_{v}=\ddot{\omega}_{v}(t) .
\end{array}
$$

We now associate the true path with a varied one for which the coordinates and velocities remain unchanged, i.e., we take:

$$
\delta x_{v}=\delta y_{v}=\delta z_{v}=\delta \dot{x}_{v}=\delta \dot{y}_{v}=\delta \dot{z}_{v}=0 .
$$

Only the acceleration components should suffer small variations:

$$
\delta \ddot{x}_{v}, \quad \delta \ddot{y}_{v}, \quad \delta \ddot{z}_{v} \quad(\neq 0)
$$

We call such a variation that affects only the acceleration a Gaussian variation $\left({ }^{2}\right)$ and denote that by putting the index $g$ in the symbol $\delta$ when a more precise characterization is desirable.

Now let $m$ constraint equations exist, say non-holonomic-rheonomic ones, to remain as general as possible:

$$
\begin{equation*}
\sum_{v=1}^{n}\left(a_{k v}^{\prime} d x_{v}+b_{k v}^{\prime} d y_{v}+c_{k v}^{\prime} d z_{v}+a_{k}^{\prime} d t\right)=0 \quad(k=1,2, \ldots, m) \tag{199}
\end{equation*}
$$

in which the coefficients $a_{k v}^{\prime}, b_{k v}^{\prime}, c_{k v}^{\prime}, a_{k}^{\prime}$ are functions of the coordinates and time $t$. The true velocity components will then obey the equations:

[^11]\[

$$
\begin{equation*}
\sum_{v=1}^{n}\left(a_{k v}^{\prime} \dot{x}_{v}+b_{k v}^{\prime} \dot{y}_{v}+c_{k v}^{\prime} \dot{z}_{v}+a_{k}^{\prime}\right)=0 \quad(k=1,2, \ldots, m) . \tag{200}
\end{equation*}
$$

\]

Upon differentiating that with respect to $t$, we will get the relations that the true accelerations must satisfy:

$$
\begin{equation*}
\sum_{v=1}^{n}\left(a_{k v}^{\prime} \ddot{x}_{v}+b_{k v}^{\prime} \ddot{y}_{v}+c_{k v}^{\prime} \ddot{z}_{v}\right)+\Phi_{k}\left(x_{v}, y_{v}, z_{v}, \dot{x}_{v}, \dot{y}_{v}, \dot{z}_{v}, t\right)=0 \quad(k=1,2, \ldots, m) \tag{201}
\end{equation*}
$$

in which we have combined all of the terms that do not contain second derivatives with respect to time in $\Phi_{k}$. $\Phi_{k}$ will then depend upon only time, the coordinates, and the velocities.

If we now impose the demand that Gauss did, namely, that the varied motion should likewise satisfy equations (201), which are true for the true accelerations, when the quantities $\ddot{x}_{v}, \ddot{y}_{v}, \ddot{z}_{v}$ are replaced with the varied ones $\ddot{x}_{v}+\delta \ddot{x}_{v}, \ddot{y}_{v}+\delta \ddot{y}_{v}, \ddot{z}_{v}+\delta \ddot{z}_{v}$ :

$$
\sum_{v=1}^{n}\left[a_{k v}^{\prime}\left(\ddot{x}_{v}+\delta \ddot{x}_{v}\right)+b_{k v}^{\prime}\left(\ddot{y}_{v}+\delta \ddot{y}_{v}\right)+c_{k v}^{\prime}\left(\ddot{z}_{v}+\delta \ddot{z}_{v}\right)+\Phi_{k}\left(x_{v}, y_{v}, z_{v}, \dot{x}_{v}, \dot{y}_{v}, \dot{z}_{v}, t\right)=0\right.
$$

and we subtract equation (201) from that then it will then follow that the quantities $\delta \ddot{x}_{v}, \delta \ddot{y}_{v}$, $\delta \ddot{z}_{v}$ are subject to the constraints:

$$
\begin{equation*}
\sum_{v=1}^{n}\left(a_{k v}^{\prime} \delta \ddot{x}_{v}+b_{k v}^{\prime} \delta \ddot{y}_{v}+c_{k v}^{\prime} \delta \ddot{z}_{v}\right)=0 \quad(k=1,2, \ldots, m) \tag{202}
\end{equation*}
$$

The quantities $\delta \ddot{x}_{v}, \delta \ddot{y}_{v}, \delta \ddot{z}_{v}$ then fulfill the constraint equations in the same sense as was previously required of the virtual displacements $\delta x_{v}, \delta y_{v}, \delta z_{v}$. The conclusions that could be inferred from that property at the time must now be true for the $\delta \ddot{x}_{v}, \delta \ddot{y}_{v}, \delta \ddot{z}_{v}$ then.

From equation (41), the equations of motion for our system will then read:

$$
\left\{\begin{array}{l}
X_{v}-m_{v} \ddot{x}_{v}+\sum_{k=1}^{m} \lambda_{k} a_{k v}^{\prime}=0  \tag{203}\\
Y_{v}-m_{v} \ddot{y}_{v}+\sum_{k=1}^{m} \lambda_{k} b_{k v}^{\prime}=0 \\
Z_{v}-m_{v} \ddot{z}_{v}+\sum_{k=1}^{m} \lambda_{k} c_{k v}^{\prime}=0
\end{array}\right.
$$

If we successively extend those equations by $\delta \ddot{x}_{v}, \delta \ddot{y}_{v}, \delta \ddot{z}_{v}$ and sum over the entire system then we will get:

$$
\begin{aligned}
\sum_{v=1}^{n}\left[\left(m_{v} \ddot{x}_{v}-X_{v}\right) \delta \ddot{x}_{v}+\right. & \left.\left(m_{v} \ddot{y}_{v}-Y_{v}\right) \delta \ddot{y}_{v}+\left(m_{v} \ddot{z}_{v}-Z_{v}\right) \delta \ddot{z}_{v}\right] \\
& =\sum_{k=1}^{m} \lambda_{k}\left\{\sum_{v=1}^{n}\left(a_{k v}^{\prime} \delta \ddot{x}_{v}+b_{k v}^{\prime} \delta \ddot{y}_{v}+c_{k v}^{\prime} \delta \ddot{z}_{v}\right)\right\} .
\end{aligned}
$$

However, since the $\delta \ddot{x}_{v}, \delta \ddot{y}_{v}, \delta \ddot{z}_{v}$ must obey equations (202), in complete analogy with the virtual displacements, they will annul the right-hand side, and the result will be:

$$
\begin{equation*}
\sum_{v=1}^{n}\left[\left(m_{v} \ddot{x}_{v}-X_{v}\right) \delta \ddot{x}_{v}+\left(m_{v} \ddot{y}_{v}-Y_{v}\right) \delta \ddot{y}_{v}+\left(m_{v} \ddot{z}_{v}-Z_{v}\right) \delta \ddot{z}_{v}\right]=0, \tag{204}
\end{equation*}
$$

which are equivalent to the equations of motion, and obviously d'Alembert's principle, as well.
Now, Gauss defined the following expression:

$$
\begin{equation*}
\mathbf{Z}=\sum_{v=1}^{n} \frac{1}{m_{v}}\left[\left(m_{v} \ddot{x}_{v}-X_{v}\right)^{2}+\left(m_{v} \ddot{y}_{v}-Y_{v}\right)^{2}+\left(m_{v} \ddot{z}_{v}-Z_{v}\right)^{2}\right] \tag{205}
\end{equation*}
$$

to be the constraint (we will explain the rationale for that terminology later), and remarked that applying the operation $\delta_{g}$ to (205) and setting the result equal to zero would produce precisely equation (204), so the equations of motion:

$$
\begin{equation*}
\delta_{g} Z \equiv \sum_{v=1}^{n}\left[\left(m_{v} \ddot{x}_{v}-X_{v}\right)^{2}+\cdots\right]=0, \tag{206}
\end{equation*}
$$

such that the statement that $\delta_{g} Z=0$ will be equivalent to the equations of motion. Indeed, the constraint is always a minimum, since $\delta^{2} Z>0$, which one can convince oneself of by calculation. We then have Gauss's principle:

For the true motion, the constraint is a minimum for a given position and velocity.

Now, how does one come to call the expression (205) the "constraint"? Following Gauss, we consider the $v^{\text {th }}$ mass-particle at two successive moments in time $t$ and $t+d t$, and indeed for the time being, we assume that the equations of constraint have been suspended and only the explicit forces $X_{v}$,


Figure 6. $Y_{v}, Z_{v}$ are in effect. We then assume that the mass-
particle is under the simultaneous effect of the explicit forces and the equations of constraint (which resolve into forces of constraint), i.e., we then consider the true motion of the massparticle $\left(^{1}\right)$.

At time $t$, let the particle in question by at the point $A_{v}$ with the coordinates $x_{v}, y_{v}, z_{v}$ in the first way of considering the situation, and at time $t+d t$, let it be at $B_{v}$ with the coordinates $x_{v}+$ $\dot{x}_{v} d t+\frac{1}{2} \ddot{x}_{v} d t^{2}, \ldots, z_{v}+\dot{z}_{v} d t+\frac{1}{2} \ddot{z}_{v} d t^{2}$. That motion will be performed under the common action of the explicit forces and the equations of constraint, and under their combined influence, the true accelerations $\ddot{x}_{v}, \ddot{y}_{v}, \ddot{z}_{v}$ will occur.

Now, the three line segments $\overrightarrow{A_{v} B_{v}}, \overrightarrow{A_{v} C_{v}}$, and $\overrightarrow{B_{v} C_{v}}$ obviously have the following components:

$$
\begin{array}{llll}
\overrightarrow{A_{v} B_{v}}: & \dot{x}_{v} d t+\frac{1}{2} \frac{X_{v}}{m_{v}} d t^{2}, & \dot{y}_{v} d t+\frac{1}{2} \frac{Y_{v}}{m_{v}} d t^{2}, & \dot{z}_{v} d t+\frac{1}{2} \frac{Z_{v}}{m_{v}} d t^{2}, \\
\overrightarrow{A_{v} C_{v}}: & \dot{x}_{v} d t+\frac{1}{2} \ddot{x}_{v} d t^{2}, & \dot{y}_{v} d t+\frac{1}{2} \ddot{y}_{v} d t^{2}, & \dot{z}_{v} d t+\frac{1}{2} \ddot{z}_{v} d t^{2},
\end{array}
$$

so $\overrightarrow{B_{v} C_{v}}$ which is the vectorial difference of the two segments that were just considered, will have the components:

$$
\overrightarrow{B_{v} C_{v}}: \quad \frac{1}{2}\left(\ddot{x}_{v}-\frac{X_{v}}{m_{v}}\right) d t^{2}, \quad \frac{1}{2}\left(\ddot{y}_{v}-\frac{Y_{v}}{m_{v}}\right) d t^{2}, \quad \frac{1}{2}\left(\ddot{z}_{v}-\frac{Z_{v}}{m_{v}}\right) d t^{2} .
$$

The magnitude of the segment $\overrightarrow{B_{v} C_{v}}$ then measures the deviation from the true motion of the free one (as Gauss called it), i.e., the one that would be followed in the absence of constraint equations. Now, Gauss's principle demands that the sum $\sum_{v=1}^{n} m_{v}\left(\overrightarrow{B_{v} C_{v}}\right)^{2}$ should be a minimum for the true motion, i.e., that the mass $m_{\nu}$ multiplied by the square of that deviation should be as small as possible. One also sees that it is just that expression that one can use as a measure of the "constraint" (in the sense of the method of least-squares, which obviously guided Gauss's reasoning in presenting the principle). In a certain sense, the mass factors represent the "weight" of the deviation from the "free" motion in each case. If we form $\sum m_{v}\left(\overrightarrow{B_{v} C_{v}}\right)^{2}$ then will find that:

$$
\sum_{v=1}^{n} m_{v}\left[\left(\ddot{x}_{v}-\frac{X_{v}}{m_{v}}\right)^{2}+\cdots\right]=\sum_{v=1}^{n} \frac{1}{m_{v}}\left[\left(m_{v} \ddot{x}_{v}-X_{v}\right)^{2}+\cdots\right]
$$

[^12]when we omit the constants factors $\frac{1}{4} d t^{4}$, i.e., we will find just the expression (205), which defines the constraint.

Furthermore, one can easily get back to Hertz's principle of the straightest path from Gauss's principle of least constraint, although we will not go into the details of that here.

## § 19.

## The Gibbs-Appell form of the equations of motion.

The analytical formulation of Gauss's principle of least constraint that is included in equation (204) leads to a new form of the equations of motion that goes back to Gibbs ( ${ }^{1}$ ) and Appell ( ${ }^{2}$ ) and is suited to non-holonomic constraints, unlike the Lagrange equations.

Once more, we would like to consider our system of $m$ mass-points whose freedom is, however, restricted by $m$ holonomic equations of constraint, such that the number of degrees of freedom will amount to $3 n-m=N$.

In any event, we can then express the differentials of the coordinates $d x_{v}, d y_{v}, d z_{v}$ as linear functions of $N$ independent differentials $d q_{1}, d q_{2}, \ldots, d q_{N}$. If the equations are integrable then the coordinates $x_{v}, y_{v}, z_{v}$ are themselves representable as functions of $N$ independent parameters $q_{1}$, $\ldots, q_{N}$ that are then referred to as the true coordinates. However, if the integrability conditions are not fulfilled then the $N$ quantities $d q_{k}$ will represent non-holonomic differentials, which we previously denoted by $d \pi_{k}$. For the sake of generality, we would like to leave undecided which of the cases we are dealing with, and for that reason, we will pose the relations as follows:

$$
\left\{\begin{array}{l}
d x_{v}=a_{1 v} d q_{1}+a_{2 v} d q_{2}+\cdots+a_{N v} d q_{N},  \tag{207}\\
d y_{v}=b_{1 v} d q_{1}+b_{2 v} d q_{2}+\cdots+b_{N v} d q_{N}, \\
d z_{v}=c_{1 v} d q_{1}+c_{2 v} d q_{2}+\cdots+c_{N v} d q_{N} .
\end{array} \quad(v=1,2, \ldots, n)\right.
$$

We will then obtain the velocities $\dot{x}_{v}, \dot{y}_{v}, \dot{z}_{v}$ upon dividing that by $d t$ :

$$
\left\{\begin{array}{l}
\dot{x}_{v}=a_{1 v} \dot{q}_{1}+a_{2 v} \dot{q}_{2}+\cdots+a_{N v} \dot{q}_{N},  \tag{208}\\
\dot{y}_{v}=b_{1 v} \dot{q}_{1}+b_{2 v} \dot{q}_{2}+\cdots+b_{N v} \dot{q}_{N},
\end{array} \quad(v=1,2, \ldots, n)\right.
$$

Upon differentiating that with respect to $t$, we will ultimately obtain the following relations between the Cartesian components of the acceleration $\ddot{x}_{v}, \ddot{y}_{v}, \ddot{z}_{v}$ and the general ones $\ddot{q}_{k}$ :

$$
\left\{\begin{array}{l}
\ddot{x}_{v}=a_{1 v} \ddot{q}_{1}+a_{2 v} \ddot{q}_{2}+\cdots+a_{N v} \ddot{q}_{N}+\Phi_{1 v}\left(x_{v}, y_{v}, z_{v}, \dot{x}_{v}, \dot{y}_{v}, \dot{z}_{v}, t\right),  \tag{209}\\
\ddot{y}_{v}=b_{1 v} \ddot{q}_{1}+b_{2 v} \ddot{q}_{2}+\cdots+b_{N v} \ddot{q}_{N}+\Phi_{2 v}, \\
\ddot{z}_{v}=c_{1 v} \ddot{q}_{1}+c_{2 v} \ddot{q}_{2}+\cdots+c_{N v} \ddot{q}_{N}+\Phi_{3 v},
\end{array}\right.
$$

[^13]in which we have combined all of the terms that do not include second derivatives with respect to $t$ into the $\Phi$. If we form the Gaussian variation $\delta_{g}$ of the acceleration components, i.e., a variation that leaves $x_{v}, y_{v}, z_{v}, \dot{x}_{v}, \dot{y}_{v}, \dot{z}_{v}, t$ unchanged while affecting only the acceleration, then it will follow that:
\[

\left\{$$
\begin{array}{l}
\delta \ddot{x}_{v}=a_{1 v} \delta \ddot{q}_{1}+a_{2 v} \delta \ddot{q}_{2}+\cdots+a_{N v} \delta \ddot{q}_{N},  \tag{210}\\
\delta \ddot{y}_{v v}=b_{1 v} \delta \ddot{q}_{1}+b_{2 v} \delta \ddot{q}_{2}+\cdots+b_{N v} \delta \ddot{q}_{N}, \\
\delta \ddot{z}_{v}=c_{1 v} \delta \ddot{q}_{1}+c_{2 v} \delta \ddot{q}_{2}+\cdots+c_{N v} \delta \ddot{q}_{N} .
\end{array}
$$ \quad(v=1,2, ···, n)\right.
\]

The coefficients $a_{\lambda v}, b_{\lambda v}, c_{\lambda v}$ will then have the following meanings accordingly:

$$
\begin{equation*}
a_{\lambda \nu}=\frac{\partial \ddot{x}_{v}}{\partial \ddot{q}_{\lambda}}, \quad b_{\lambda v}=\frac{\partial \ddot{y}_{v}}{\partial \ddot{q}_{\lambda}}, \quad c_{\lambda v}=\frac{\partial \ddot{z}_{v}}{\partial \ddot{q}_{\lambda}} . \tag{211}
\end{equation*}
$$

We would like to introduce the expressions for $\delta \ddot{x}_{v}, \delta \ddot{y}_{v}, \delta \ddot{z}_{v}$ into equation (204) for the principle of least constraint. We first split that equation into two parts that correspond to the explicit forces and the inertial forces:

$$
\sum_{v=1}^{n}\left[X_{v} \delta \ddot{x}_{v}+Y_{v} \delta \ddot{y}_{v}+Z_{v} \delta \ddot{z}_{v}\right]=\sum_{v=1}^{n} m_{v}\left[\ddot{x}_{v} \delta \ddot{x}_{v}+\ddot{y}_{v} \delta \ddot{y}_{v}+\ddot{z}_{v} \delta \ddot{z}_{v}\right]
$$

If we perform the stated substitution here then we will get:

$$
\begin{aligned}
& \sum_{v=1}^{n}\left[X_{v} \sum_{\lambda=1}^{N} a_{\lambda v} \delta \ddot{q}_{\lambda}+Y_{v} \sum_{\lambda=1}^{N} b_{\lambda v} \delta \ddot{q}_{\lambda}+Z_{v} \sum_{\lambda=1}^{N} c_{\lambda v} \delta \ddot{q}_{\lambda}\right] \\
&=\sum_{v=1}^{n} m_{v}\left[\ddot{x}_{v} \sum_{\lambda=1}^{N} a_{\lambda v} \delta \ddot{q}_{\lambda}+\ddot{y}_{v} \sum_{\lambda=1}^{N} b_{\lambda v} \delta \ddot{q}_{\lambda}+\ddot{z}_{v} \sum_{\lambda=1}^{N} c_{\lambda v} \delta \ddot{q}_{\lambda}\right],
\end{aligned}
$$

or when rearranged:

$$
\sum_{\lambda=1}^{N}\left[\sum_{v=1}^{n}\left(X_{v} a_{\lambda v}+Y_{v} b_{\lambda v}+Z_{v} c_{\lambda v}\right)\right] \delta \ddot{q}_{\lambda}=\sum_{\lambda=1}^{N}\left[\sum_{v=1}^{n} m_{v}\left(\ddot{x}_{v} a_{\lambda v}+\ddot{y}_{v} b_{\lambda v}+\ddot{z}_{v} c_{\lambda v}\right)\right] \delta \ddot{q}_{\lambda} .
$$

However, due to the independence of the $\delta \ddot{q}_{\lambda}$, that equation will decompose into $N$ individual equations:

$$
\begin{equation*}
\sum_{v=1}^{n}\left(X_{v} a_{\lambda v}+Y_{v} b_{\lambda v}+Z_{v} c_{\lambda v}\right)=\sum_{v=1}^{n} m_{v}\left(\ddot{x}_{v} a_{\lambda v}+\ddot{y}_{v} b_{\lambda v}+\ddot{z}_{v} c_{\lambda v}\right) \quad(\lambda=1,2, \ldots, N) . \tag{212}
\end{equation*}
$$

The expression on the left has a simple mechanical meaning: Namely, it is the general force component $Q_{\lambda}$ that strives to vary the coordinate $q_{\lambda}$. One will see that most simply when one introduces the values of $\delta x_{v}, \delta y_{v}, \delta z_{v}$ from (207) into the expression for the work $\sum\left[X_{v} \delta x_{v}+\right.$
$\left.Y_{v} \delta y_{v}+Z_{\nu} \delta z_{v}\right]$ and arranges the result in terms of the $\delta q_{\lambda}$. The factor of $\delta q_{\lambda}$ is then just $Q_{\lambda}$. One likewise introduces the values of the coefficients $a_{\lambda v}, b_{\lambda v}, c_{\lambda v}$ from (211) into the right-hand side of the last equation. One will then get:

$$
\begin{equation*}
Q_{\lambda}=\sum_{v=1}^{n} m_{v}\left(\ddot{x}_{v} \frac{\partial \ddot{x}_{v}}{\partial \ddot{q}_{\lambda}}+\ddot{y}_{v} \frac{\partial \ddot{y}_{v}}{\partial \ddot{q}_{\lambda}}+\ddot{z}_{v} \frac{\partial \ddot{z}_{v}}{\partial \ddot{q}_{\lambda}}\right) \quad(\lambda=1,2, \ldots, N), \tag{213}
\end{equation*}
$$

or finally:

$$
\begin{equation*}
Q_{\lambda}=\frac{\partial}{\partial \ddot{q}_{\lambda}} \sum_{v=1}^{n} \frac{1}{2} m_{v}\left(\ddot{x}_{v}^{2}+\ddot{y}_{v}^{2}+\ddot{z}_{v}^{2}\right) \quad(\lambda=1,2, \ldots, N) . \tag{214}
\end{equation*}
$$

Here, the forces $Q_{\lambda}$ are expressed as the partial derivatives of a function that is formed in the same way as kinetic energy, except with the components of acceleration, instead of velocity components. We would like to introduce a special terminology for that function: We shall refer to:

$$
\begin{equation*}
T=\sum_{v=1}^{n} \frac{1}{2} m_{v}\left(\ddot{x}_{v}^{2}+\ddot{y}_{v}^{2}+\ddot{z}_{v}^{2}\right) \tag{215}
\end{equation*}
$$

as the Gibbs-Appell function. The equations of motion will then read:

$$
\begin{equation*}
Q_{\lambda}=\frac{\partial T}{\partial \ddot{q}_{\lambda}} . \tag{216}
\end{equation*}
$$

It is clear from the derivation that this will also be true when the coordinates are nonholonomic.

We would like to make an application of precisely that formula to the Euler equations of a rigid body that is fixed at a point. We have already treated that problem before with the help of the general Lagrange equations (§ 10).

The following equations in the form of (207) are true in this case:

$$
\left\{\begin{array}{l}
d x_{v}=z_{v} d \chi-y_{v} d \rho,  \tag{217}\\
d y_{v}=x_{v} d \rho-z_{v} d \pi \\
d z_{v}=y_{v} d \pi-x_{v} d \chi
\end{array}\right.
$$

The $x_{v}, y_{v}, z_{v}$ in that mean the coordinates of a mass-point in the rigid body in a coordinate system that is fixed in the body and whose origin coincides with the fixed point. We will impose further specializations later. Furthermore, $d \pi, d \chi, d \rho$ are the infinitesimal angles of rotation around the axes of that system. The velocities will follow upon dividing by $d t$ :

$$
\left\{\begin{array}{l}
\dot{x}_{v}=z_{v} \dot{\chi}-y_{v} \dot{\rho}  \tag{218}\\
\dot{y}_{v}=x_{v} \dot{\rho}-z_{v} \dot{\pi} \\
\dot{z}_{v}=y_{v} \dot{\pi}-x_{v} \dot{\chi}
\end{array}\right.
$$

In order to remain true to our previous notations, we shall not actually write $\dot{\pi}, \dot{\chi}, \dot{\rho}$, but rather $\stackrel{*}{\pi}, \stackrel{*}{\chi}, \stackrel{*}{\rho}$, since there are no such quantities as $\pi, \chi, \rho$ whose derivatives are the $\dot{\pi}, \dot{\chi}, \dot{\rho}$. Nonetheless, we would like to preserve the dot in the following calculations, for the sake of simplicity of notation, since hopefully no misunderstanding should arise.

We find that the acceleration components are:

$$
\begin{aligned}
& \ddot{x}_{v}=z_{v} \ddot{\chi}-y_{v} \ddot{\rho}+\dot{z}_{v} \dot{\chi}-\dot{y}_{v} \dot{\rho}, \\
& \ddot{y}_{v}=x_{v} \ddot{\rho}-z_{v} \ddot{\pi}+\dot{x}_{v} \dot{\rho}-\dot{z}_{v} \dot{\pi} \\
& \ddot{z}_{v}=y_{v} \ddot{\pi}-x_{v} \ddot{\chi}+\dot{y}_{v} \dot{\pi}-\dot{x}_{v} \dot{\chi} .
\end{aligned}
$$

If we once more substitute the value of $\dot{x}_{v}, \dot{y}_{v}, \dot{z}_{v}$ from (218) in that then that will yield:

$$
\left\{\begin{array}{l}
\ddot{x}_{v}=z_{v} \ddot{\chi}-y_{v} \ddot{\rho}+y_{v} \dot{\pi} \dot{\chi}-x_{v} \dot{\chi}^{2}-x_{v} \dot{\rho}^{2}+z_{v} \dot{\pi} \dot{\rho}  \tag{219}\\
\ddot{y}_{v}=x_{v} \ddot{\rho}-z_{v} \ddot{\pi}+z_{v} \dot{\chi} \dot{\rho}-y_{v} \dot{\rho}^{2}-y_{v} \dot{\pi}^{2}+x_{v} \dot{\chi} \dot{\pi} \\
\ddot{z}_{v}=y_{v} \ddot{\pi}-x_{v} \ddot{\chi}+x_{v} \dot{\rho} \dot{\pi}-z_{v} \dot{\pi}^{2}-z_{v} \dot{\chi}^{2}+y_{v} \dot{\rho} \dot{\chi}
\end{array}\right.
$$

If we now form $T$ then it would greatly simplify the calculations if we were to drop all terms in which no double dots appear, because those terms would make no contribution under Gaussian variation. We would then get, e.g., for $\ddot{x}_{v}^{2}$ :

$$
\begin{aligned}
\ddot{x}_{v}^{2} & =z_{v}^{2} \ddot{\chi}^{2}+y_{v}^{2} \ddot{\rho}^{2}-2 z_{v} y_{v} \ddot{\chi} \ddot{\rho}+2 z_{v} y_{v} \dot{\pi} \dot{\chi} \ddot{\chi}-2 x_{v} z_{v} \dot{\chi}^{2} \ddot{\chi}-2 x_{v} z_{v} \dot{\rho}^{2} \ddot{\chi} \\
& +2 z_{v}^{2} \dot{\pi} \dot{\rho} \ddot{\chi}-2 y_{v}^{2} \dot{\pi} \dot{\chi} \ddot{\rho}+2 x_{v} y_{v} \dot{\chi}^{2} \ddot{\rho}+2 x_{v} y_{v} \dot{\rho}^{2} \ddot{\rho}-2 y_{v} z_{v} \dot{\pi} \dot{\rho} \ddot{\rho}+\cdots
\end{aligned}
$$

The remaining components of the acceleration follow from that by cyclic permutation of the symbols $x, y, z$, as well as $\pi, \chi, \rho$.

If we now form $\sum m_{v}\left(\ddot{x}_{v}^{2}+\ddot{y}_{v}^{2}+\ddot{z}_{v}^{2}\right)$ then terms with the factors $\sum m_{v} x_{v}^{2}, \sum m_{v} y_{v}^{2}$, $\sum m_{v} z_{v}^{2}$, and ones with the factors $\sum m_{v} x_{v} y_{v}, \sum m_{v} y_{v} z_{v}, \sum m_{v} z_{v} x_{v}$ will appear. However, the latter are known to be the so-called moments of deviation about the axes, and we will succeed in making them drop out when we let the coordinate system that is used in the derivation of the Euler equations coincide with the principal axes of inertia through the fixed point. We will then have, quite simply:

$$
\begin{gathered}
2 T=\ddot{\pi}^{2} \sum_{v} m_{v}\left(y_{v}^{2}+z_{v}^{2}\right)+\ddot{\chi}^{2} \sum_{v} m_{v}\left(z_{v}^{2}+x_{v}^{2}\right)+\ddot{\rho}^{2} \sum_{v} m_{v}\left(y_{v}^{2}+x_{v}^{2}\right) \\
+2 \dot{\pi} \dot{\rho} \ddot{\chi} \sum_{v} m_{v}\left(z_{v}^{2}-x_{v}^{2}\right)+2 \dot{\chi} \dot{\pi} \ddot{\rho} \sum_{v} m_{v}\left(x_{v}^{2}-y_{v}^{2}\right)+2 \dot{\rho} \dot{\chi} \ddot{\pi} \sum_{v} m_{v}\left(y_{v}^{2}-x_{v}^{2}\right)+\cdots
\end{gathered}
$$

Now, when the principal moments of inertia are denoted by $A, B, C$, as usual:

$$
A=\sum m_{v}\left(y_{v}^{2}+z_{v}^{2}\right), \quad B=\sum m_{v}\left(z_{v}^{2}+x_{v}^{2}\right), \quad C=\sum m_{v}\left(y_{v}^{2}+x_{v}^{2}\right),
$$

and one makes suitable subtractions of them:

$$
B-A=\sum m_{v}\left(x_{v}^{2}-y_{v}^{2}\right), \quad C-B=\sum m_{v}\left(y_{v}^{2}-z_{v}^{2}\right), \quad A-C=\sum m_{v}\left(z_{v}^{2}-x_{v}^{2}\right),
$$

the expression for $T$ will become (naturally, except for terms that carry no double dots):

$$
\begin{equation*}
T=\frac{1}{2} A \ddot{\pi}^{2}+\frac{1}{2} B \ddot{\chi}^{2}+\frac{1}{2} C \ddot{\rho}^{2}+(A-C) \dot{\pi} \dot{\rho} \ddot{\chi}+(B-A) \dot{\chi} \dot{\pi} \ddot{\rho}+(C-B) \dot{\rho} \dot{\chi} \ddot{\pi} . \tag{220}
\end{equation*}
$$

If we now denote the generalized force components by $\Pi, \mathrm{X}, \mathrm{P}$, as before, in the sense that $\Pi \delta \pi+\mathrm{X} \delta \chi+\mathrm{P} \delta \rho$ means the work done under the infinitesimal rotation $\delta \pi, \delta \chi, \delta \rho$ then we will have from (216) and (220) that:

$$
\begin{aligned}
& \Pi=\frac{\partial T}{\partial \ddot{\pi}}=A \ddot{\pi}+(C-A) \dot{\chi} \dot{\rho}, \\
& \mathrm{X}=\frac{\partial T}{\partial \ddot{\chi}}=B \ddot{\chi}+(A-C) \dot{\rho} \dot{\pi}, \\
& \mathrm{P}=\frac{\partial T}{\partial \ddot{\rho}}=C \ddot{\rho}+(B-A) \dot{\pi} \dot{\chi},
\end{aligned}
$$

i.e., the Euler equations. It is remarkable that the Gibbs-Appell form of the equations of motion allows one to achieve that result with significantly less computational work that one expends with the extended Lagrange equations for non-holonomic coordinates.


[^0]:    $\left({ }^{1}\right)$ On this, cf., O. Hölder, "Über die Prinzipien von Hamilton und Maupertuis," Sitz.-Ber. d. Ges. d. Wiss. zu Göttingen (1896), pp. 122, et seq.

[^1]:    $\left.{ }^{1}{ }^{1}\right)$ Cf., infra, pp. 36.
    $\left(^{2}\right)$ Hölder, loc. cit.

[^2]:    $\left({ }^{1}\right)$ The recognition of this fact, as well as the term "non-holonomic coordinates," goes back to Boltzmann [Wien. Sitz.-Ber. 111 (1902), pp. 1603].

[^3]:    ( ${ }^{1}$ ) H. Hertz, Prinzipien der Mechanik, pp. 22, et seq.

[^4]:    ( ${ }^{1}$ ) This equation was first proved by L. Boltzmann [Wien. Sitz.-Ber. 111 (1902), pp. 1603; also Ges. Abh. III, pp. 682]. Later, G. Hamel [Math. Ann. 59 (1904), pp. 416, et seq.] devoted an extended investigation to that question.

[^5]:    $\left({ }^{1}\right)$ Cf., also the presentation by H. von Helmholtz, "Zur Geschichte des Prinzips der kleinsten Aktion," Ges. Abhandl., Bd. III, pp. 249.

[^6]:    $\left.{ }^{( }{ }^{1}\right)$ C. G. J. Jacobi, Vorlesungen über Dynamik, Werke, Supplementband, 1884. Lecture 6.

[^7]:    ( ${ }^{1}$ ) H. Hertz, Die Prinzipien der Mechanik, pp. 69.

[^8]:    ( ${ }^{1}$ ) C. G. J. Jacobi, Vorlesungen über Dynamik, Lecture 20.

[^9]:    ( ${ }^{1}$ ) P. Stäckel, Über die Integration der Hamilton-Jacobischen Differentialgleichung mittels Separation der Variabeln, Habilitationsschrift, Halle, 1891; one might also, cf., the very clear and simple presentation in Charlier, Mechanik des Himmels, Bd. I, pp. 77, et seq.
    $\left(^{2}\right)$ K. Schwarzschild, Sitz.-Ber. d. Berl. Akad. d. Wiss. (1916), pp. 548.
    $\left({ }^{3}\right)$ P. S. Epstein, Phys. Zeit. 17 (1916), pp. 148; Ann. Phys. (Leipzig) 50 (1916), pp. 489; ibid., (1916), pp. 815; ibid., 51 (1916), pp. 168.

[^10]:    $\left.{ }^{1}{ }^{1}\right)$ P. Debye, Nachr. kgl. Ges. Wiss. zu Göttingen, math.-phys. Klasse (1916), pp. 142.
    $\left(^{2}\right)$ P. Scherrer, ibid. (1916), pp. 154.

[^11]:    $\left.{ }^{1}{ }^{1}\right)$ C. F. Gauss, "Über ein neues allgemeines Grundgesetz der Mechanik," J. reine angew. Math. 4 (1829); also Werke, Bd. V, pp. 23, et seq., 1877.
    $\left(^{2}\right)$ That terminology goes back to Boltzmann (Vorl. über d. Prinzipe der Mechanik, Bd. I, pp. 209.)

[^12]:    ${ }^{(1)}$ ) On this, one can confer, say: Boltzmann, Vorlesungen über die Prinzipe der Mechanik, Bd. I, 1897, pp. 212.

[^13]:    $\left(^{1}\right)$ J. W. Gibbs, "On the fundamental formulae of Dynamics," Am. J. Math. 2 (1879), pp. 49, et seq.; Scientific Papers, vol. II, pp. 1, et seq.
    $\left(^{2}\right)$ P. Appell, "Sur une forme générale des équations de la dynamique," C. R. Acad. Sci. Paris 129 (1899), pp. 317; also Traité de mécanique rationelle, $2^{\text {nd }}$ ed., t. II, 1904, pp. 292.

