## Non-holonomic Relativity and the Unitary Theory of Einstein and Mayer

## By

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VRANCEAUNU [1] has recently established a non-holonomic unitary theory of fields in the schema of his theory of non-holonomic spaces [2].

The present note has the goal of making the relationship between this non-holonomic theory of relativity and the unitary theory of Einstein and MAYER [3] as clear as possible.

In order to establish the non-holonomic unitary theory of fields one may, as the author has already pointed out [4], start with the following hypotheses:

(A) A Riemannian space  $V_5$  admits an infinitesimal point translation:

(1) 
$$x^{\lambda^*} = x^{\lambda} + \phi^{\lambda} dt,$$
  $(\lambda, \mu, \nu ... = 1, 2, 3, 4, 5)$ 

that leaves invariant the fundamental quadratic form:

(2) 
$$d\sigma^2 = G_{\lambda\mu} \, dx^\lambda dx^\mu$$

in which  $\phi^{\lambda} = \delta_{5}^{\lambda}$  is a unit vector.

(B) Spacetime is identified with a non-holonomic space  $V_5^4$  that is defined in  $V_5$  by:

$$\phi_{\lambda} \, dx^{\lambda} = 0,$$

in which  $\phi_{\lambda} = G_{\lambda\mu} \phi^{\mu}$ .

If one considers only coordinate transformations of the form:

(4) 
$$\begin{cases} \overline{x}^{i} = \overline{x}^{i} (x^{j}) \\ \overline{x}^{5} = \overline{x}^{5} + f(x^{j}), \quad (i, j, k, ... = 1, 2, 3, 4), \end{cases}$$

then one concludes from these hypotheses that the components  $G_{\lambda\mu}$  of the fundamental tensor in any coordinate system are independent of the fifth coordinate  $x^5$  and  $G_{5\lambda} = \phi_{\lambda}$ ,  $G_{55} = 1$ .

Moreover, one may write (2) in the form:

(5) 
$$d\sigma^2 = g_{ij}dx^i dx^j + (\phi_i dx^i + dx^5)^2,$$

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in which  $g_{ij} = G_{ij} - \phi_i \phi_j$ .

Since the  $g_{ij}$  transform under the coordinate transformation (4) like the components of a tensor in a space  $V_4$  that is described by  $x^i$ , they constitute ten functions that represent the gravitational field, whereas the  $\phi_i$  are the components of a vector in  $V_4$  that represents the electromagnetic field.

The non-holonomic space  $V_4^5$  thus defined is totally geodesic. Since a geodesic in  $V_4^5$ , i.e., a geodesic that satisfies the equation:

$$\phi_{\lambda} \, dx^{\lambda} = 0,$$

is given by [5]:

$$\frac{d^2x^{\lambda}}{ds^2} + (\Gamma^{\lambda}_{\mu\nu} + \phi^{\lambda}\phi_{\mu,\nu})\frac{dx^{\mu}}{ds}\frac{dx^{\nu}}{ds} = 0,$$

in which:

(6)

(7) 
$$\Gamma^{\lambda}_{\mu\nu} = \frac{1}{2} G^{\lambda\omega} \left( \frac{\partial G_{\alpha\mu}}{\partial x^{\nu}} + \frac{\partial G_{\omega\nu}}{\partial x^{\mu}} - \frac{\partial G_{\mu\nu}}{\partial x^{\omega}} \right),$$
$$\phi_{\mu,\nu} = \frac{\partial \phi_{\mu}}{\partial x^{\nu}} - \phi_{\lambda} \Gamma^{\lambda}_{\mu\nu}$$
$$= \frac{1}{2} \left( \frac{\partial \phi_{\mu}}{\partial x^{\nu}} - \frac{\partial \phi_{\nu}}{\partial x^{\mu}} \right),$$

as a consequence (60) reduces to:

$$\frac{d^2x^{\lambda}}{ds^2} + \Gamma^{\lambda}_{\mu\nu}\frac{dx^{\mu}}{ds}\frac{dx^{\nu}}{ds} = 0;$$

however, these are the equations for geodesics in  $V_5$ .

Thus  $V_4^5$  is totally geodesic.

We define a tensor  $g_{\lambda\mu}$  in  $V_5$  by:

(8) 
$$g_{\lambda\mu} = G_{\lambda\mu} - \phi_{\lambda}\phi_{\mu}.$$

Since this tensor satisfies the identity:

$$g_{5\lambda}=g_{\lambda 5}=0,$$

one may consider that  $g_{\lambda i}$  transforms under a coordinate transformation (4) like the components of a tensor in  $V_5$  with respect to the index  $\lambda$ , and like the components of a vector in  $V_4$  with respect to *i*.

Thus, the two quantities:

(9)  $g_{\lambda}^{i} = g^{ij}g_{j\lambda}$   $(g^{ij} g_{ij} = \delta_{k}^{i}),$ and: (10)  $g_{i}^{\lambda} = G^{\lambda\mu}g_{\mu i}$   $(G^{\lambda\mu}G_{\mu\nu} = \delta_{\nu}^{\lambda}),$ are mixed tensors. As one may easily verify, the  $G^{\lambda\mu}$  satisfy the equations:

(11)  
$$\begin{cases} G^{ij} = g^{ij}, \\ G^{i5} = -g^{ij}\phi_j, \\ G^{55} = 1 + g^{ij}\phi_i\phi_j, \end{cases}$$

and therefore  $g_{\lambda}^{i}$  and  $g_{i}^{\lambda}$  have the following explicit forms:

(12) 
$$\begin{cases} g_i^{\lambda} = \delta_i^{\lambda} - \phi^{\lambda} \phi_i, \\ g_{\lambda}^i = \delta_{\lambda}^i, \end{cases}$$

from which one easily obtains the identities:

(13) 
$$g_{ij} = g_i^{\lambda} g_j^{\mu} G_{\lambda\mu},$$

(14) 
$$G_{\lambda\mu} = g^i_{\lambda} g^j_{\mu} g_{ij} + \phi_{\lambda} \phi_{\mu},$$

(15) 
$$g^i_{\lambda}g^j_{\mu} = \delta^j_i,$$

(16) 
$$g_i^{\lambda}g_{\mu}^{i} = \delta_{\mu}^{\lambda} - \phi^{\lambda}\phi_{\mu},$$

(17) 
$$\phi_{\lambda} g_{i}^{\lambda} = 0,$$

 $g_i^{\lambda}$ ,  $g_{\lambda}^{i}$ , and  $\phi_{\lambda}$  correspond to  $\gamma_i^{\lambda}$ ,  $\gamma_{\lambda}^{j}$ , and  $A_{\lambda}$ , respectively, in the previously-cited memoir of EINSTEIN and MAYER.

When a vector  $V^{\lambda}$  in  $V_5$  satisfies:

$$\phi_{\lambda} V^{\lambda} = 0,$$

one says that it is in  $V_4^5$ . A covariant vector  $W_{\lambda}$  is in  $V_4^5$  when it satisfies:

$$W_{\lambda} \phi^{\lambda} = 0.$$

These definitions may be extended to affinors. For example,  $g_{\lambda\mu}$ , as defined by (8), is certainly in  $V_4^5$ , since:

$$g_{\lambda\mu}\phi^{\lambda}\phi^{\mu}=g_{55}=0.$$

When one is given a vector  $v_i$  in  $V_4$ , i.e., a space framed by the coordinates  $x^1$ ,  $x^2$ ,  $x^3$ , and  $x^4$ , one may form a vector in  $V_5$  by  $g_i^{\lambda} v^i$ , and this vector is in  $V_4^5$ :

$$\phi_{\lambda}(g_i^{\lambda}v^i)=0.$$

If one is given a vector  $v^i$  in  $V_4$  and a scalar  $\rho$  then one may form a vector  $V^{\lambda}$  that is not in  $V_4^5$  from:

(18) 
$$g_i^{\lambda} v^i + \rho \phi^{\lambda}.$$

We shall consider the covariant derivative to be geometrically of this type.

A connection in the space  $V_5$  is well defined by the CHRISTOFFEL symbols that one forms from  $G_{\lambda\mu}$ :

$$\Gamma^{\lambda}_{\mu\nu} = \frac{1}{2} G^{\lambda\omega} \left( \frac{\partial G_{\omega\mu}}{\partial x^{\nu}} + \frac{\partial G_{\omega\nu}}{\partial x^{\mu}} - \frac{\partial G_{\mu\nu}}{\partial x^{\omega}} \right),$$

which have the following values:

(19)  
$$\begin{cases} \Gamma_{\mu}^{\lambda} = \Gamma_{5\mu}^{\lambda} = \phi_{\mu}^{\lambda}, \\ \Gamma_{jk}^{i} = \begin{cases} i \\ jk \end{cases} + \phi_{j}^{i}\phi_{k} + \phi_{k}^{i}\phi_{j}, \\ \Gamma_{jk}^{5} = -\phi_{i}\Gamma_{jk}^{i} + \frac{1}{2}\left(\frac{\partial\phi_{j}}{\partial x^{k}} + \frac{\partial\phi_{k}}{\partial x^{j}}\right), \end{cases}$$

in which:

(20)  
$$\begin{cases} \phi_{\lambda\mu} = \frac{1}{2} \left( \frac{\partial \phi_{\lambda}}{\partial x^{\mu}} - \frac{\partial \phi_{\mu}}{\partial x^{\lambda}} \right), \\ \phi_{\mu}^{\lambda} = G^{\lambda\nu} \phi_{\nu\mu}, \\ \phi_{j}^{i} = g^{ik} \phi_{kj}, \\ \left\{ \begin{array}{c} i \\ jk \end{array} \right\} = \frac{1}{2} g^{ih} \left( \frac{\partial g_{hj}}{\partial x^{k}} + \frac{\partial g_{hk}}{\partial x^{j}} - \frac{\partial g_{jk}}{\partial x^{h}} \right). \end{cases}$$

Take a vector  $V^{\mu}$  whose components are independent of the variable  $x^5$  and consider a displacement in  $V_4^5$ :

$$\partial V^{\nu} = \frac{\partial V^{\nu}}{\partial x^{\mu}} dx^{\mu} + \Gamma^{\nu}_{\lambda\mu} V^{\lambda} dx^{\mu}$$

in which:

$$\frac{\partial V^{\nu}}{\partial x^{\mu}}dx^{\mu}=\frac{\partial V^{\nu}}{\partial x^{i}}dx^{i},$$

and:

$$dx^5 = -\phi_i \, dx^i,$$

since the vector must be in  $V_4^5$ , which is defined by  $\phi_{\lambda} dx^{\lambda} = 0$ .

One therefore has:

(21) 
$$\delta V^{\nu} = \left(\frac{\partial V^{\nu}}{\partial x^{i}} + \Pi^{\nu}_{\lambda i} V^{\lambda}\right) dx^{i},$$

in which:

(22) 
$$\Pi^{\nu}_{\lambda i} = \Gamma^{\nu}_{\lambda i} - \Gamma^{\nu}_{\lambda 5} \phi_i$$

In the sequel, we let ";" denote the covariant derivative with respect to the connection parameters  $\Pi_{\lambda i}^{\nu}$ . For a mixed tensor, one defines the covariant derivative as follows:

(22) 
$$T_{\lambda i;j}^{\nu} = \frac{\partial T_{\mu i}^{\lambda}}{\partial x^{j}} + T_{\nu i}^{\lambda} \Pi_{\mu j}^{\nu} - T_{\mu i}^{\nu} \Pi_{\nu j}^{\lambda} - T_{\mu k}^{\lambda} \begin{cases} k \\ ij \end{cases}.$$

Since:

$$\Pi^{\lambda}_{\mu i} = \Gamma^{\lambda}_{\mu i} - \Gamma^{\lambda}_{\mu 5} \phi_i,$$

one has, for  $G_{\lambda\mu}$ :

$$G_{\lambda\mu\,i} = \frac{\partial G_{\lambda\mu}}{\partial x^{i}} + G_{\nu\mu}\Pi^{\nu}_{\lambda i} - G_{\lambda\nu}\Pi^{\nu}_{\mu i}$$
$$= \left(\frac{\partial G_{\lambda\mu}}{\partial x^{i}} + G_{\nu\mu}\Gamma^{\nu}_{\lambda i} - G_{\lambda\nu}\Gamma^{\nu}_{\mu i}\right) + (\phi_{\lambda\mu} + \phi_{\mu\lambda})\phi_{i}$$

Both of the two terms in the right-hand side are annulled, due to the definitions, just like  $\Gamma^{\lambda}_{\mu\nu}$  and  $\phi_{\lambda\mu}$ , and one thus obtains:

$$G_{\lambda\mu\,i}=0$$

This is the first condition that was posed by Einstein and Mayer (loc. cit., pp. 548 (I)), which geometrically signifies that the length of a vector does not change when it is parallel displaced with respect to  $\Pi_{\mu i}^{\lambda}$ .

Having said this, we calculate the covariant derivative  $g_{i;j}^{\lambda}$  of  $g_i^{\lambda}$ . Before doing this, we write down the relationship between  $\Gamma_{\mu\nu}^{\lambda}$  and  $\begin{cases} a \\ bc \end{cases}$ :

(24) 
$$\Gamma^{\lambda}_{\mu\nu} = g^{\lambda}_{a} g^{b}_{\mu} g^{c}_{\nu} \left\{ \begin{matrix} a \\ bc \end{matrix} \right\} + \frac{1}{2} \phi^{\lambda} \left( \frac{\partial \phi_{\mu}}{\partial x^{\nu}} + \frac{\partial \phi_{\nu}}{\partial x^{\mu}} \right) + \phi^{\lambda}_{\mu} \phi_{\nu} + \phi^{\lambda}_{\nu} \phi_{\mu},$$

from which, one obtains:

(25) 
$$\Pi^{\lambda}_{\mu i} = g^{\lambda}_{a} g^{b}_{\mu} \begin{cases} a \\ bi \end{cases} + \frac{1}{2} \phi^{\lambda} \left( \frac{\partial \phi_{\mu}}{\partial x^{i}} + \frac{\partial \phi_{i}}{\partial x^{\mu}} \right) + \phi^{\lambda}_{i} \phi_{\nu} .$$

Now:

$$g_{i;j}^{\lambda} = \frac{\partial g_i^{\lambda}}{\partial x^j} + g_i^{\nu} \Pi_{\nu j}^{\lambda} - g_k^{\lambda} \begin{cases} k \\ ij \end{cases},$$

but, on the other hand:

$$\frac{\partial g_i^{\lambda}}{\partial x^j} = \frac{\partial}{\partial x^j} (\delta_i^{\lambda} - \phi^{\lambda} \phi_i)$$

$$= -\phi^{\lambda} \frac{\partial \phi_{i}}{\partial x^{j}},$$

$$g_{i}^{\nu} \Pi_{\nu j}^{\lambda} = g_{a}^{\lambda} \begin{cases} a \\ ij \end{cases} + \frac{1}{2} \phi^{\lambda} \left( \frac{\partial \phi_{i}}{\partial x^{j}} + \frac{\partial \phi_{j}}{\partial x^{i}} \right),$$

so therefore one has:

(26)

The geometrical significance of (26) is the following: If  $v^{i}$  is perculal displayed in K is a

 $g_{i:i}^{\lambda} = -\phi^{\lambda}\phi_{ij}.$ 

If  $v^i$  is parallel displaced in  $V_4$ , i.e.:

 $\delta v^i = 0,$ 

then the corresponding vector  $g_i^{\lambda} v^i$  in  $V_4^5$ , after being displaced in  $V_4^5$ , satisfies the equation:

$$\delta(g_i^{\lambda}v^i) = (\delta g_i^{\lambda}) v^i$$
$$= -\phi^{\lambda} (\phi_{ij} v^i dx^j).$$

Therefore, the covariant derivative  $\delta(g_i^{\lambda}v^i)$  is normal to  $V_4^5$ .

This is the second condition of Einstein and Mayer (loc. cit., pp. 549 (II)).

Furthermore,  $\phi_{ij}$  is an antisymmetric tensor in (26), so (26) also expresses the fact that:

If  $v^i$  is parallel displaced in its proper direction in  $V_4$  then the corresponding vector  $g_i^{\lambda}v^i$  is also parallel displaced in  $V_4^5$ , since:

$$\delta(g_i^{\lambda}v^i) = (\delta g_i^{\lambda}) v^i$$
  
=  $-\phi^{\lambda} \phi_{ij} v^i v^j dt$   
= 0.

This is the third condition of EINSTEIN and MAYER (loc. cit., pp. 549 (III)). Moreover, one has:

(29)  
$$\phi_{i}^{\lambda} = \frac{\partial \phi^{\lambda}}{\partial x^{i}} + \phi^{\nu} \Pi_{\nu i}^{\lambda}$$
$$= g_{j}^{\lambda} \phi_{i}^{j},$$

which geometrically signifies that the covariant derivative of  $\phi^{\lambda}$  is always found in  $V_4^5$ .

The equations of the trajectories will then be obtained in exactly the same manner as those of EINSTEIN and MAYER.

In spacetime, a curve is given by:

$$x^{i}(s),$$

but when one is dealing with a trajectory for an electrically charged particle one must also specify a scalar  $\rho = e/m$ , where *e* is the charge and *m* is the mass of that particle.

One may then define a vector in  $V_5$  by:

(30) 
$$V^{\lambda} = g_{i}^{\lambda} \frac{dx^{i}}{ds} + \rho \phi^{\lambda}.$$

One poses the following conditions with respect to the trajectory:

The vector  $V^{\lambda}$  is parallel displaced along the curve in V<sub>4</sub>, from which, one has:

$$V^{\lambda}_{\ j}\frac{dx^{j}}{ds}=0.$$

Therefore, one has:

(31) 
$$g_{i;j}^{\lambda} \frac{dx^{i}}{ds} \frac{dx^{j}}{ds} + g_{i}^{\lambda} \left( \frac{d^{2}x^{i}}{ds^{2}} + \begin{cases} i \\ jk \end{cases} \frac{dx^{j}}{ds} \frac{dx^{k}}{ds} \end{pmatrix} + \phi^{\lambda} \rho_{;j} \frac{dx^{j}}{ds} + \rho \phi^{\lambda}_{;j} \frac{dx^{j}}{ds} = 0.$$

However, on the other hand, one has:

$$g_{i;j}^{\lambda} = -\phi^{\lambda} \phi_{ij},$$

$$\rho_{j} \frac{dx^{j}}{ds} = (\phi_{\lambda} V^{\lambda})_{jj} \frac{dx^{j}}{ds} = \phi_{\lambda;j} V^{\lambda} \frac{dx^{j}}{ds}$$

$$= g_{\lambda}^{a} \phi_{aj} (g_{i}^{\lambda} \frac{dx^{i}}{ds} + \rho \phi^{\lambda}) \frac{dx^{j}}{ds}$$

$$= \phi_{ij} \frac{dx^{i}}{ds} \frac{dx^{j}}{ds}$$

$$= 0,$$

so one finally has:

(32)  
$$\frac{d^2x^i}{ds^2} + \begin{cases} i\\ jk \end{cases} \frac{dx^j}{ds} \frac{dx^k}{ds} + \rho \phi_j^i \frac{dx^j}{ds} = 0,$$
$$\rho = \frac{e}{m} = \text{constant},$$

upon contracting (31) by  $g_{\lambda}^{h}$ .

In the memoir of EINSTEIN and MAYER, the [Lagrangian] density whose variation would give the field equations is nowhere to be found.

Here, one may take the variation:

(33) 
$$\delta \int K g^{\frac{1}{2}} dx^1 dx^2 dx^3 dx^4 = 0,$$

where:

$$\begin{split} g &= \mid g_{ij} \mid = \mid G_{\lambda\mu} \mid, \\ K &= G^{\lambda\mu} K_{\mu\lambda}, \qquad K_{\mu\lambda} = K^{\dots \ \omega}_{\omega\mu\lambda}, \\ K^{\dots \ \omega}_{\omega\mu\lambda} &= \frac{\partial \Gamma^{\nu}_{\lambda\omega}}{\partial x^{\mu}} - \frac{\partial \Gamma^{\nu}_{\lambda\mu}}{\partial x^{\omega}} + \Gamma^{\kappa}_{\lambda\omega} \Gamma^{\nu}_{\kappa\mu} - \Gamma^{\kappa}_{\lambda\mu} \Gamma^{\nu}_{\kappa\omega}. \end{split}$$

(33) gives us [6]:(34)in which:

$$abla_{\lambda\mu} - \phi_{\lambda} \phi_{\mu} 
abla = 0,$$
 $abla_{\lambda\mu} = K_{\lambda\mu} - rac{1}{2} G_{\lambda\mu} K,$ 

$$abla = G^{\lambda\mu} \nabla_{\lambda\mu}.$$

Upon contracting (34) by  $g_i^{\lambda} g_j^{\mu}$ , we have:

(35)  
$$R_{ij} + 2 \phi_i^k \phi_{kj} - \frac{1}{2} g_{ij} K = 0,$$
$$R_{ij} - \frac{1}{2} g_{ij} K + 2 \left[ \phi_i^k \phi_{kj} + \frac{1}{4} g_{ij} \phi_h^k \phi_k^h \right] = 0,$$

in which  $R_{ij}$  and R are the Ricci tensor and the scalar curvature formed from  $g_{ij}$ , respectively.

In (34), if we contract with  $g_i^{\lambda}$  and set  $\mu = 5$ , then we obtain:

$$\phi_{i;j}^k = 0.$$

 $\nabla_{55} - \nabla = 0,$ 

Finally, if we set  $\lambda = \mu = 5$  then we have:

but:

$$\nabla = G^{\lambda\mu} \nabla_{\lambda\mu} = G^{\lambda\mu} (K_{\lambda\mu} - \frac{1}{2} G_{\lambda\mu} K)$$
  
=  $-\frac{3}{2} K$ ,  
$$\nabla_{55} = K_{55} - \frac{1}{2} K$$
  
=  $\phi_i^{\ j} \phi_j^i - \frac{1}{2} K$ ,

so:

$$\nabla_{55} - \nabla = K + \phi_i^j \phi_j^i.$$

R = 0.

As a consequence: (36)

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